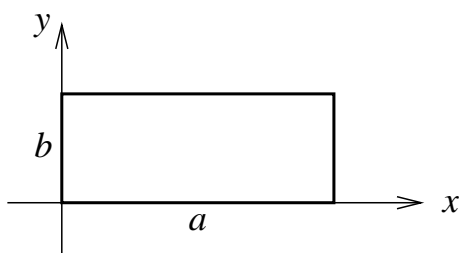


Midterm — Solutions

This midterm is a two hour open book, open notes exam. Do all three problems.

- [35 pts] 1. Consider the propagation of waves in a rectangular waveguide with sides of lengths a and b .



- [15] a) Show that (for $m > 0$ and $n > 0$) a TE_{mn} and TM_{mn} mode can be superposed to make the x component of the magnetic field vanish, $H_x = 0$.

For a right-moving TE_{mn} mode defined by

$$\psi^{\text{TE}} = H_0 \cos \frac{m\pi x}{a} \cos \frac{n\pi y}{b} \quad (1)$$

the x component of the magnetic field is given by

$$\vec{H}_t = \frac{ik}{\gamma^2} \vec{\nabla}_t \psi^{\text{TE}} \quad \Rightarrow \quad H_x^{\text{TE}} = -\frac{ik}{\gamma^2} H_0 \frac{m\pi}{a} \sin \frac{m\pi x}{a} \cos \frac{n\pi y}{b} \quad (2)$$

Likewise, for a right-moving TM_{mn} mode defined by

$$\psi^{\text{TM}} = E_0 \sin \frac{m\pi x}{a} \sin \frac{n\pi y}{b} \quad (3)$$

the x component of the magnetic field is

$$\vec{H}_t = \frac{i\epsilon\omega}{\gamma^2} \hat{z} \times \vec{\nabla}_t \psi^{\text{TM}} \quad \Rightarrow \quad H_x^{\text{TM}} = -\frac{i\epsilon\omega}{\gamma^2} E_0 \frac{n\pi}{b} \sin \frac{m\pi x}{a} \cos \frac{n\pi y}{b}$$

We note that in both cases, the modes have identical eigenvalues

$$\gamma^2 = \pi^2 \left[\left(\frac{m}{a} \right)^2 + \left(\frac{n}{b} \right)^2 \right] \quad (4)$$

and identical functional behavior for H_x . This allows us to superpose and cancel the x component of the magnetic field

$$0 = H_x = H_x^{\text{TE}} + H_x^{\text{TM}} = -\frac{i}{\gamma^2} \left(kH_0 \frac{m\pi}{a} + \epsilon\omega E_0 \frac{n\pi}{b} \right) \sin \frac{m\pi x}{a} \cos \frac{n\pi y}{b}$$

The solution is

$$E_0 = -\frac{k}{\epsilon\omega} \frac{m/a}{n/b} H_0$$

Of course, even though $H_x = 0$, the other components of \vec{E} and \vec{H} do not necessarily vanish. Furthermore, the two modes that are superposed must be moving in the same direction in order to have $H_x = 0$ everywhere along the z direction.

- [15] b) Compute the transmitted power in this superposition mode. (Recall that TE and TM modes are orthogonal in the sense that $\frac{1}{2} \int \hat{z} \cdot (\vec{E}_{\text{TE}} \times \vec{H}_{\text{TM}}^*) da = 0$ and similarly with TE and TM interchanged.)

The power is obtained from the Poynting vector

$$P = \int_A \hat{z} \cdot \vec{S} da = \frac{1}{2} \int_A \hat{z} \cdot (\vec{E} \times \vec{H}^*) da = \frac{1}{2} \int_A \hat{z} \cdot [(\vec{E}^{\text{TE}} + \vec{E}^{\text{TM}}) \times (\vec{H}^{\text{TE}} + \vec{H}^{\text{TM}})^*] da$$

However, since the modes are orthogonal, the power simply decomposes into an incoherent sum $P = P^{\text{TE}} + P^{\text{TM}}$ where

$$P^{\text{TE}} = \frac{1}{2} \int_A \hat{z} \cdot (\vec{E}^{\text{TE}} \times \vec{H}^{\text{TE}*}) da$$

$$P^{\text{TM}} = \frac{1}{2} \int_A \hat{z} \cdot (\vec{E}^{\text{TM}} \times \vec{H}^{\text{TM}*}) da$$

We use the standard expressions for transmitted power in a waveguide

$$P^{\text{TE}} = \frac{\mu\omega k}{2\gamma^2} \int_A |\psi^{\text{TE}}|^2 da = \frac{ab}{8} \frac{\mu\omega k}{\gamma^2} |H_0|^2$$

$$P^{\text{TM}} = \frac{\epsilon\omega k}{2\gamma^2} \int_A |\psi^{\text{TM}}|^2 da = \frac{ab}{8} \frac{\epsilon\omega k}{\gamma^2} |E_0|^2$$

where ψ^{TE} and ψ^{TM} are given in (1) and (3), respectively. Hence

$$P = \frac{ab}{8} \frac{\omega k}{\gamma^2} (\epsilon |E_0|^2 + \mu |H_0|^2) = \frac{ab}{8} \frac{\mu\omega k}{\gamma^2} \left[1 + \frac{k^2}{\mu\epsilon\omega^2} \left(\frac{m/a}{n/b} \right)^2 \right] |H_0|^2$$

Finally, noting that $k^2 = \mu\epsilon\omega^2 - \gamma^2$ as well as the definition of γ^2 in (4), the power expression may be rewritten as

$$P = \frac{ab}{8} \frac{\mu\omega k}{(n\pi/b)^2} \left[1 - \frac{(m\pi/a)^2}{\mu\epsilon\omega^2} \right] |H_0|^2$$

This may alternatively be written in terms of E_0 as

$$P = \frac{ab}{8} \frac{\epsilon\omega k}{(m\pi/a)^2} \left[1 + \frac{(n\pi/b)^2}{k^2} \right] |E_0|^2$$

[5] c) Can we still have $H_x = 0$ when $m = 0$ or $n = 0$?

If either $m = 0$ or $n = 0$, then the TM mode does not exist. In this case we cannot superpose the two modes. However, the TE mode itself may have a vanishing H_x . For H_x^{TE} given in (2), this may occur when $m = 0$. In particular, the TE_{0n} mode has $H_x = 0$, while TE_{m0} cannot have $H_x = 0$. So it is possible to have $H_x = 0$ when $m = 0$ but not when $n = 0$.

[35 pts] 2. A non-conducting sphere ($\epsilon = \epsilon_0$ and $\mu = \mu_0$) of radius a carries a uniform charge density ρ_0 throughout its volume. The sphere is centered at the origin of the coordinate system and oscillates back and forth about the z -axis with an angular velocity $\vec{\omega} = \hat{z} \phi_0 \omega \cos \omega t$.

[5] a) Show that the current density may be expressed as

$$\vec{J} = \Re[\hat{\phi} \rho_0 \phi_0 \omega r \sin \theta e^{-i\omega t}]$$

The velocity of a point \vec{r} within the sphere is given by

$$\vec{v} = \vec{\omega} \times \vec{r} = \hat{z} \times \vec{r} \phi_0 \omega \cos \omega t = \hat{\phi} \phi_0 \omega r \sin \theta \cos \omega t$$

As a result, the current density is

$$\vec{J} = \rho_0 \vec{v} = \hat{\phi} \rho_0 \phi_0 \omega r \sin \theta \cos \omega t = \Re[\hat{\phi} \rho_0 \phi_0 \omega r \sin \theta e^{-i\omega t}]$$

This current is only non-vanishing for $r < a$. For this time-harmonic current density, we may define the complex current as

$$\vec{J}(\vec{x}) = \hat{\phi} \rho_0 \phi_0 \omega r \sin \theta \quad (5)$$

[20] b) Compute the multipole radiation coefficients $a_E(l, m)$ and $a_M(l, m)$. Note that the integral $\int x^{l+2} j_l(x) dx = x^{l+2} j_{l+1}(x)$ may be helpful.

We start with the electric multipoles $a_E(l, m)$. Note from (5) that both $\rho = (1/i\omega) \vec{\nabla} \cdot \vec{J} = 0$ and $\vec{r} \cdot \vec{J} = 0$. Since

$$a_E(l, m) = \frac{k^2}{i\sqrt{l(l+1)}} \int Y_{lm}^* \left[c\rho \frac{\partial}{\partial r} [r j_l(kr)] + ik(\vec{r} \cdot \vec{J}) j_l(kr) \right] d^3x$$

we immediately conclude that all electric multipoles vanish

$$a_E(l, m) = 0$$

For the magnetic multipoles $a_M(l, m)$, we first compute

$$\vec{r} \times \vec{J} = \vec{r} \times \hat{\phi} \rho_0 \phi_0 \omega r \sin \theta = -\hat{\theta} \rho_0 \phi_0 \omega r^2 \sin \theta$$

and

$$\vec{\nabla} \cdot (\vec{r} \times \vec{J}) = \frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} \sin \theta (\vec{r} \times \vec{J})_\theta = -\frac{1}{r \sin \theta} \frac{\partial}{\partial \theta} \rho_0 \phi_0 \omega r^2 \sin^2 \theta = -2\rho_0 \phi_0 \omega r \cos \theta$$

This allows us to compute

$$\begin{aligned} a_M(l, m) &= \frac{k^2}{i\sqrt{l(l+1)}} \int Y_{lm}^* [\vec{\nabla} \cdot (\vec{r} \times \vec{J}) j_l(kr)] d^3x \\ &= \frac{k^2}{i\sqrt{l(l+1)}} (-2\rho_0 \phi_0 \omega) \int Y_{lm}^* r \cos \theta j_l(kr) r^2 dr d\Omega \\ &= \frac{2ik^2 \rho_0 \phi_0 \omega}{\sqrt{l(l+1)}} \sqrt{\frac{4\pi}{3}} \int Y_{lm}^* Y_{10} r^3 j_l(kr) dr d\Omega \\ &= \frac{2ik^2 \rho_0 \phi_0 \omega}{\sqrt{l(l+1)}} \sqrt{\frac{4\pi}{3}} \delta_{l,1} \delta_{m,0} \int_0^a r^3 j_l(kr) dr \\ &= i\sqrt{\frac{8\pi}{3}} k^2 \rho_0 \phi_0 \omega \delta_{l,1} \delta_{m,0} \int_0^a r^3 j_1(kr) dr \end{aligned}$$

where we have used the fact that $\cos \theta = P_1(\cos \theta) = \sqrt{3/4\pi} Y_{10}(\theta, \phi)$ as well as the orthonormality of the spherical harmonics. The radial integral may be performed according to

$$\int_0^a r^3 j_1(kr) dr = \frac{1}{k^4} \int_0^{ka} x^3 j_1(x) dx = \frac{x^3 j_2(x)}{k^4} \Big|_0^{ka} = \frac{(ka)^3 j_2(ka)}{k^4}$$

Hence the only non-vanishing magnetic multipole is

$$a_M(1, 0) = i\sqrt{\frac{8\pi}{3}} \rho_0 \phi_0 \omega ka^3 j_2(ka)$$

[10] c) Find the time-averaged power radiated per unit solid angle $dP/d\Omega$.

Since the radiation is purely magnetic dipole, the angular power distribution is given by

$$\begin{aligned} \frac{dP}{d\Omega} &= \frac{Z_0}{2k^2} |a_M(1, 0)|^2 |X_{10}|^2 \\ &= \frac{Z_0}{2k^2} \frac{8\pi}{3} \rho_0^2 \phi_0^2 \omega^2 k^2 a^6 j_2^2(ka) \left[\frac{3}{8\pi} \sin^2 \theta \right] \\ &= \frac{Z_0}{2} \rho_0^2 \phi_0^2 \omega^2 a^6 j_2^2(ka) \sin^2 \theta \end{aligned}$$

[30 pts] 3. Three *small* uniform and non-permeable ($\mu = \mu_0$) dielectric spheres of radii b and relative dielectric constant $\epsilon_r \approx 1$ are located along the z axis, centered at positions $z = -a$, $z = 0$ and $z = a$.

[15]

- a) Using the Born approximation, compute the unpolarized differential scattering cross section $d\sigma/d\Omega$. Assume $kb \ll 1$, but allow a to be arbitrarily small or large.

The Born approximation is given by

$$\begin{aligned}\vec{\epsilon}^* \cdot \vec{f} &= \frac{k^2}{4\pi} \int d^3x e^{i\vec{q} \cdot \vec{x}} \left[\vec{\epsilon}^* \cdot \vec{\epsilon}_0 \frac{\delta\epsilon}{\epsilon_0} + (\hat{n} \times \vec{\epsilon}^*) \cdot (\hat{n} \times \vec{\epsilon}^*) \frac{\delta\mu}{\mu_0} \right] \\ &= \frac{k^2}{4\pi} \int d^3x e^{i\vec{q} \cdot \vec{x}} \vec{\epsilon}^* \cdot \vec{\epsilon}_0 \frac{\delta\epsilon}{\epsilon_0}\end{aligned}$$

where $\vec{q} = \vec{k}_0 - \vec{k}$ and where we have used the fact that the dielectric is non-permeable. Since $\delta\epsilon$ is only non-vanishing in the interior of the dielectric spheres, the volume integral may be restricted to the three small spheres located at $\vec{x} \approx 0$, $\vec{x} \approx a\hat{z}$ and $\vec{x} \approx -a\hat{z}$. Since each sphere has a volume of $(4/3)\pi b^3$, we obtain

$$\begin{aligned}\vec{\epsilon}^* \cdot \vec{f} &\approx \frac{k^2}{4\pi} \left(\frac{4}{3}\pi b^3 \right) (\epsilon_r - 1) (\vec{\epsilon}^* \cdot \vec{\epsilon}_0) (1 + e^{iaq_z} + e^{-iaq_z}) \\ &= \frac{1}{3} k^2 b^3 (\epsilon_r - 1) (\vec{\epsilon}^* \cdot \vec{\epsilon}_0) (1 + 2\cos(aq_z))\end{aligned}$$

where

$$q_z = k_{0z} - k_z = k(\cos\theta_0 - \cos\theta)$$

Here θ_0 is the polar angle of the incident wave, and θ is the polar angle of the scattered wave. The differential scattering cross section is then

$$\frac{d\sigma}{d\Omega} = |\vec{\epsilon}^* \cdot \vec{f}|^2 \approx k^4 b^6 \left(\frac{\epsilon_r - 1}{3} \right)^2 |\vec{\epsilon}^* \cdot \vec{\epsilon}_0|^2 [1 + 2\cos(ka(\cos\theta_0 - \cos\theta))]^2$$

For unpolarized scattering, the polarization average is given by

$$|\vec{\epsilon}^* \cdot \vec{\epsilon}| \rightarrow \frac{1}{2}(1 + \cos^2\gamma)$$

where γ is the angle between \vec{k}_0 and \vec{k} (ie the incident and scattered wave). Hence

$$\frac{d\sigma}{d\Omega} \approx k^4 b^6 \left(\frac{\epsilon_r - 1}{3} \right)^2 \frac{1 + \cos^2\gamma}{2} [1 + 2\cos(ka(\cos\theta_0 - \cos\theta))]^2 \quad (6)$$

where $\cos\gamma$ is given in spherical coordinates by

$$\cos\gamma = \cos\theta \cos\theta_0 + \sin\theta \sin\theta_0 \cos(\phi - \phi_0)$$

[5]

- b) Show that, for $ka \ll 1$, the cross section is nine times as large as that for a single sphere.

For $ka \ll 1$, we may approximate

$$1 + 2\cos(ka(\cos\theta_0 - \cos\theta)) \approx 3$$

Substituting this into (6) gives

$$\frac{d\sigma}{d\Omega} \approx 9k^4b^6 \left(\frac{\epsilon_r - 1}{3} \right)^2 \frac{1 + \cos^2 \gamma}{2}$$

which is nine times as large as the single sphere result

$$\frac{d\sigma(\text{one sphere})}{d\Omega} = k^4b^6 \left(\frac{\epsilon_r - 1}{3} \right)^2 \frac{1 + \cos^2 \gamma}{2}$$

Note that for N small spheres, the long wavelength scattering cross section would be N^2 as large. At long wavelengths, the scattering is coherent, and the amplitude scales as total dielectric volume (so the cross section scales as the square of the volume).

- [10] c) Now suppose the incident plane wave is traveling in the $+x$ direction. For what values of a does the scattered power vanish along the z axis? Give your answer in terms of the wavelength $\lambda = 2\pi/k$.

If the incident wave travels along the $+x$, this corresponds to taking $\theta_0 = \pi/2$ and $\phi_0 = 0$. Inserting this into (6) gives

$$\frac{d\sigma}{d\Omega} \approx k^4b^6 \left(\frac{\epsilon_r - 1}{3} \right)^2 \frac{1 + \sin^2 \theta \cos^2 \phi}{2} [1 + 2 \cos(ka \cos \theta)]^2$$

In order to examine the scattered power along the z axis, we take $\theta = 0$ or π . In either case, the result is

$$\frac{d\sigma}{d\Omega} \approx k^4b^6 \left(\frac{\epsilon_r - 1}{3} \right)^2 \frac{1}{2} [1 + 2 \cos ka]^2$$

This vanishes when

$$1 + 2 \cos ka = 0 \quad \Rightarrow \quad \cos ka = -\frac{1}{2} \quad \Rightarrow \quad ka = \pm \frac{2\pi}{3} + 2n\pi$$

Using $k = 2\pi/\lambda$ gives

$$a = (n \pm \frac{1}{3})\lambda$$

This result can also be obtained by simply demanding destructive interference of the scattered waves from the centers of the three (small) spheres.