

2D Numerical Calculation of a Collapsar

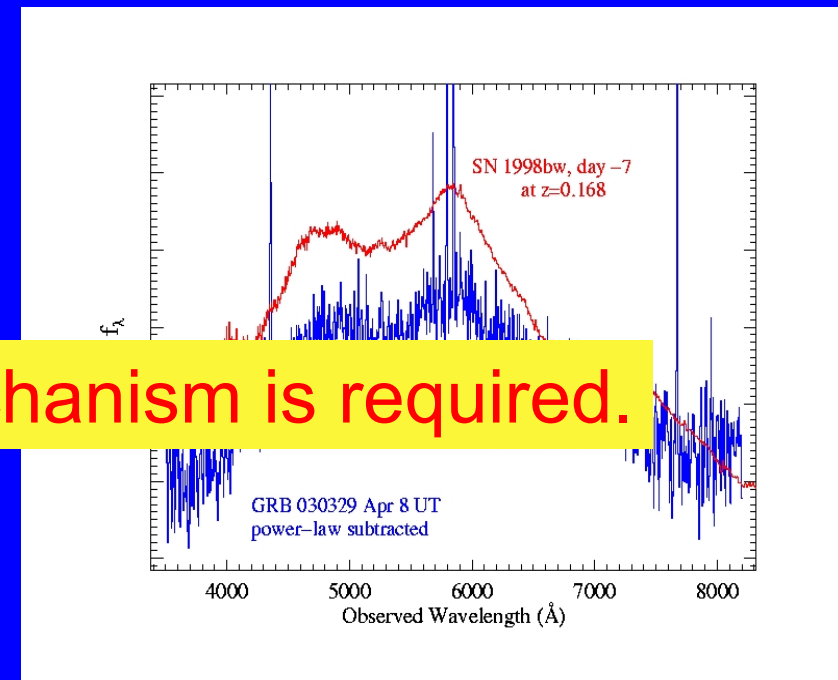
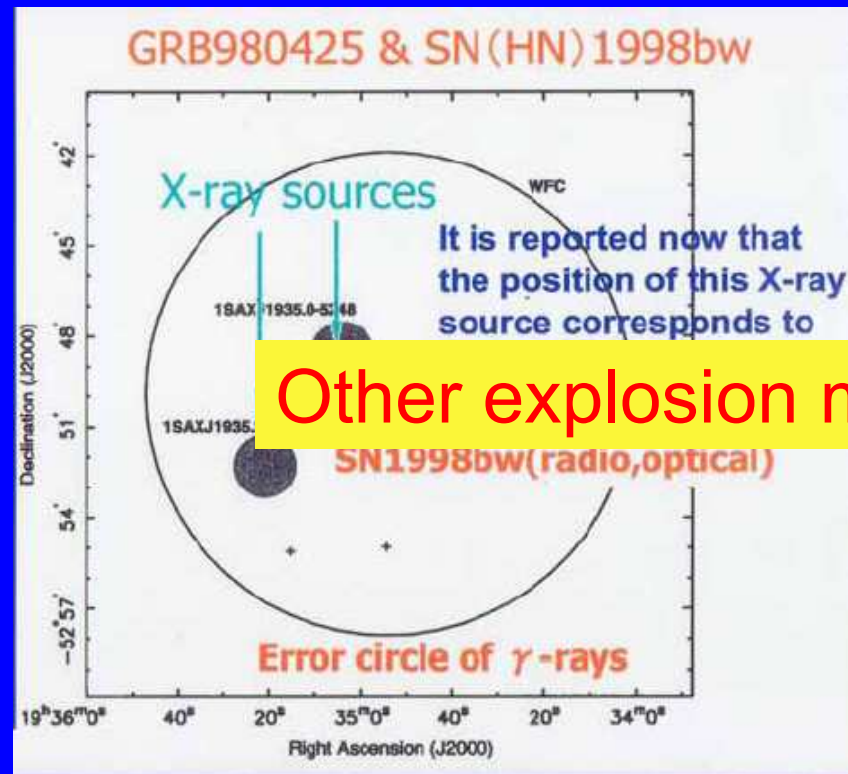
with Realistic Equation of State, Rotation, Self-gravity, Photo-disintegration, and Neutrino Cooling, (+Neutrino Heating, Magnetic Field)

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§ Introduction

Some GRBs are Accompanied by Hypernovae



Other explosion mechanism is required.

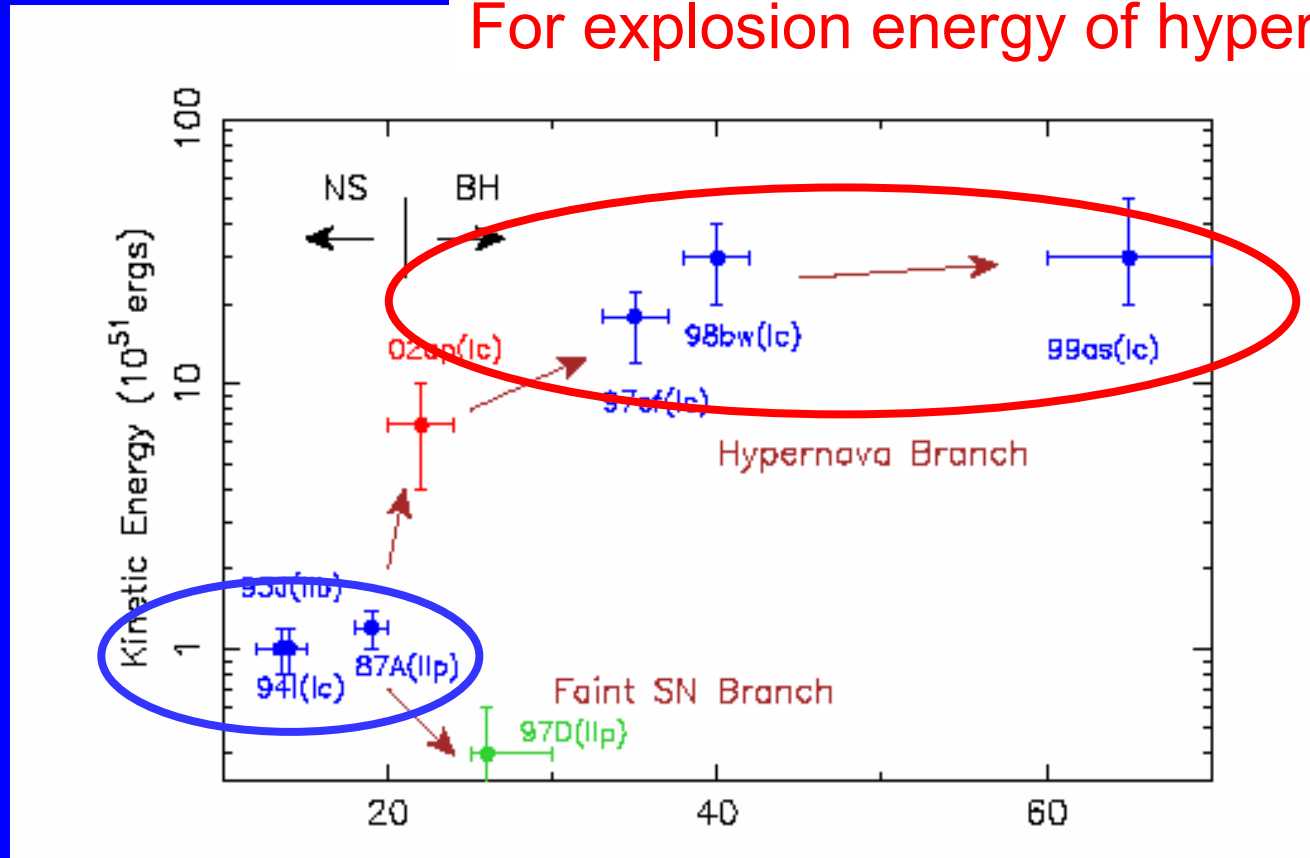
GRB030329/SN2003dh

GRB980425/SN1998bw

Explosion energy of a hypernova is estimated to be $\sim 1E+52$ ergs, which cannot be explained by the standard scenario of a normal collapse-driven supernova.

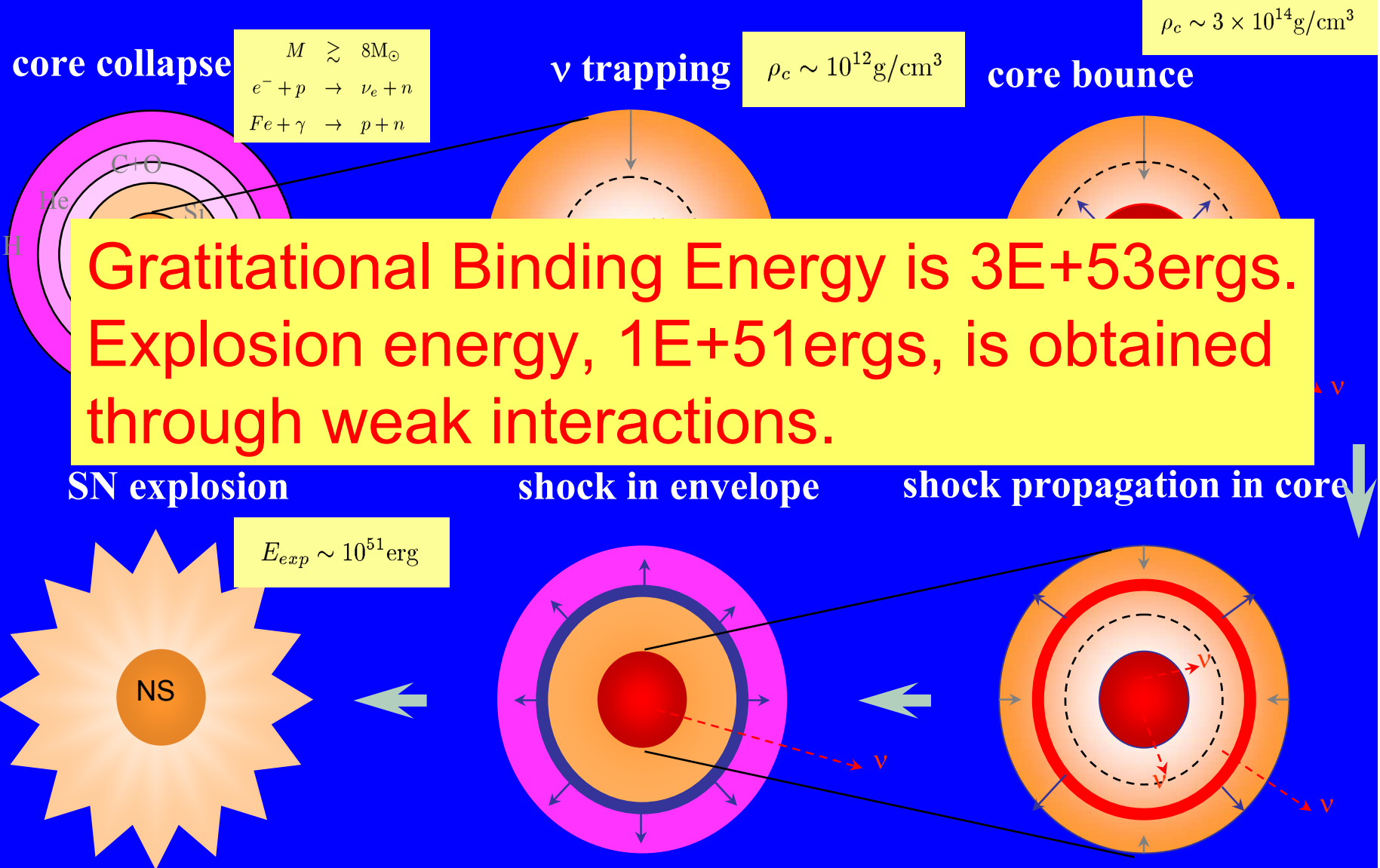
What determines the explosion energy of hypernovae?

Is there an energy scale/upper-limit
For explosion energy of hypernovae?



Nomoto et al. '02

Scenario of Normal Collapse-driven Supernovae



Some possible energy scales

$$E_{\text{pot}} \sim \frac{G M_{\text{BH}} M_{\text{acc}}}{r_{\text{H}}} \sim \frac{G M_{\text{BH}} M_{\text{acc}}}{\frac{G M_{\text{BH}}}{c^2}}$$
$$\sim M_{\text{acc}} c^2 = 2 \times 10^{54} \left(\frac{M_{\text{acc}}}{1 M_{\odot}} \right) \text{ [erg]}$$

(a) Gravitational Potential
Energy of Accreted Matter.
Duration is ~ 10 sec.
Collapsar Model

$$E_{\text{rot}} = f(a) M_{\text{BH}} c^2 \leq f(a=1) M_{\text{BH}} c^2$$
$$= 1.6 \times 10^{54} \left(\frac{M_{\text{BH}}}{3 M_{\odot}} \right) \text{ [erg]}$$

$$f(a) = 1 - \sqrt{\frac{1}{2} [1 + \sqrt{1 - a^2}]}$$

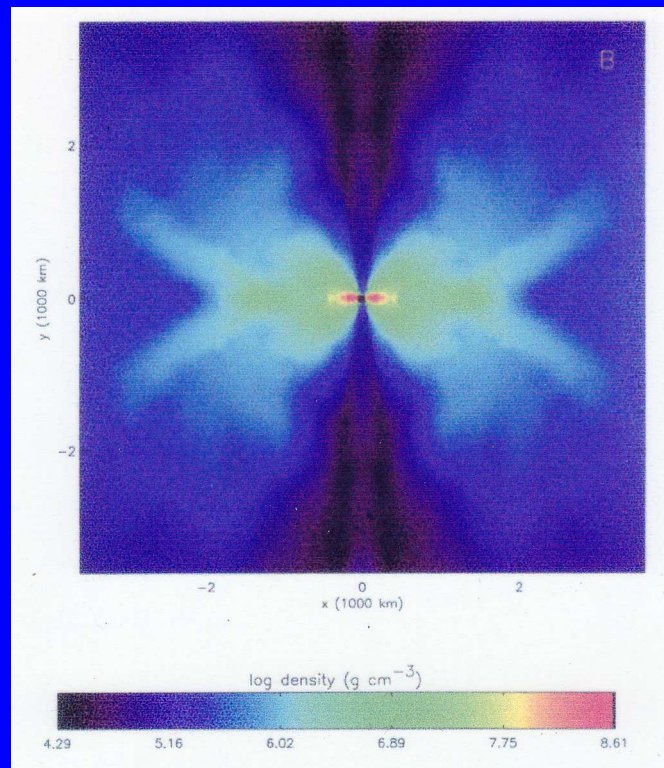
$$a = \frac{Jc}{GM^2} \leq 1$$

(b) Rotation Energy of
a central BH.
Duration is $\sim (1-10)$ sec?
Blandford-Znajek Process

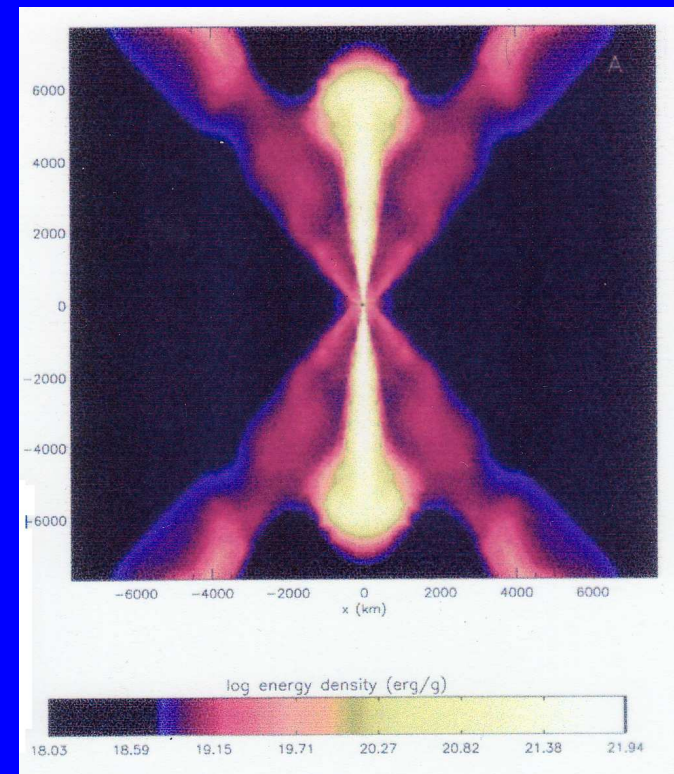
(a) Gravitational Potential Energy Model

MacFadyen and Woosley (1998)

Jet off



Jet on



BH with 3solar mass is put at the center as an initial condition.

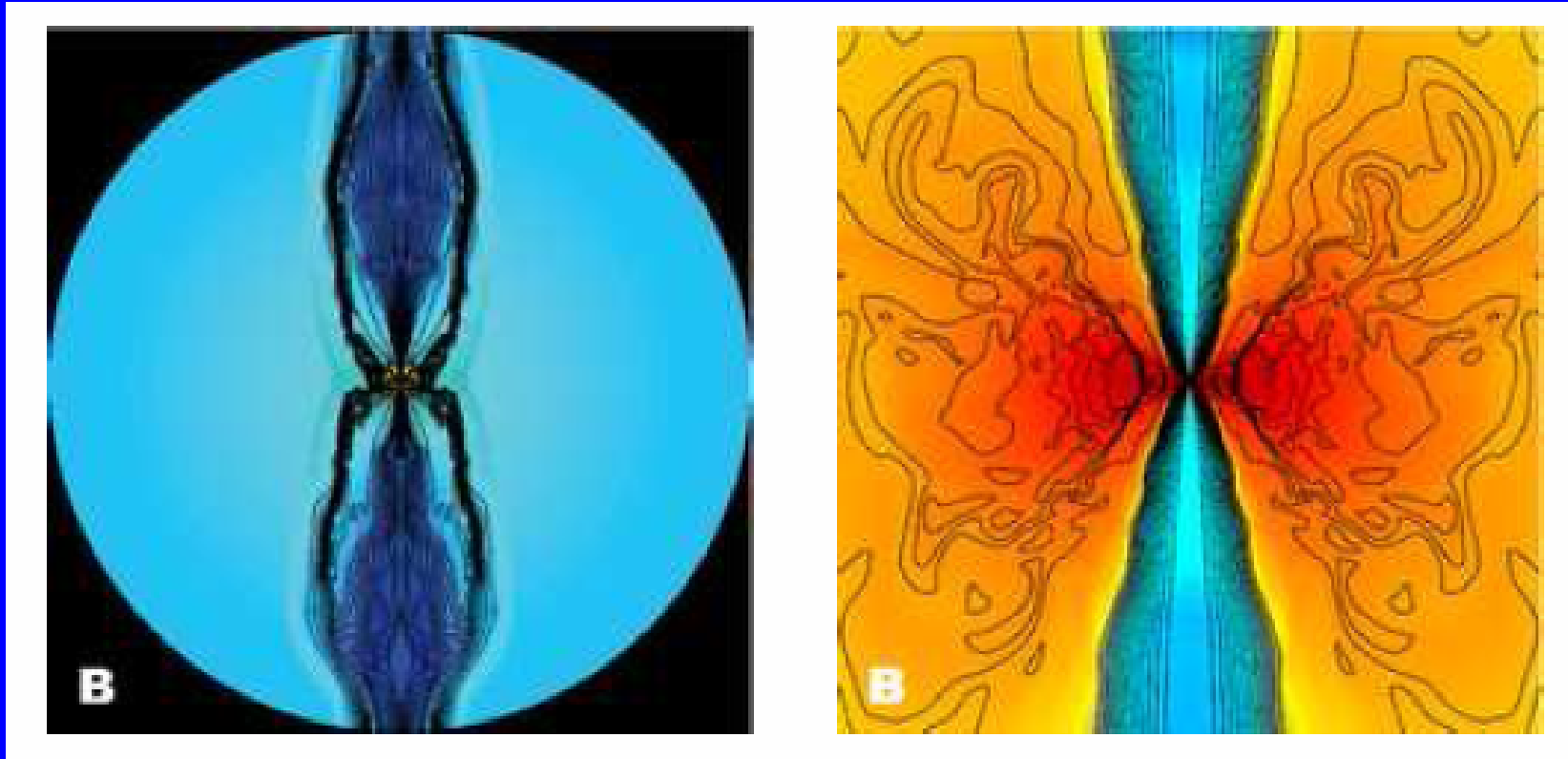
Rotation is introduced so that the model mimics Heger et al. (00).

Thermal energy is deposited at the inner most region, which might be realized when effect of neutrino annihilation is included. Newtonian gravity.

Magnetic Fields are not included. Inner most boundary is set to be 50km.

(b) Rotation Energy of a Central BH Model

McKinney 05



Blandford-Znajek Process can be seen in numerical simulations.

§ Method of Calculation

Input Physics

- Progenitor: Model E25 in Heger et al. (2000).
- 2-D Ideal MHD with Rotation $(r,\theta)=(150,30)$.
- Inner boundary is set to be 10km.
- Realistic EOS (Blinnikov et al. 1996).
- Nuclear Reaction (N, P, He, O, Fe).
- Newton Gravity (Including self gravity; $M_{\text{BH}}=2.2M_{\odot}$).
- Neutrino Cooling.
- Neutrino Heating (ON-OFF).
- Magnetic Field (ON-OFF).

Neutrino Processes (1)

Cooling Process

Locally Determined

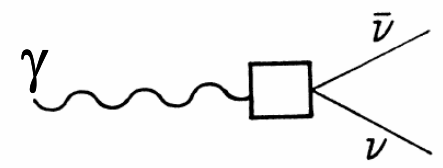
Electron and positron Capture on free nucleons
(Epstein and Pethick 1981)

$$\begin{cases} p + e^- \rightarrow n + \nu_e \\ n + e^+ \rightarrow p + \bar{\nu}_e \end{cases}$$

Pair annihilations
(Itoh et al. 1999)

$$e^- + e^+ \rightarrow \nu_l + \bar{\nu}_l$$

Plasmon decays
(Itoh et al. 1989)



Heating Process

Globally Determined

Electron-type neutrino capture on free nucleons
(Epstein and Pethick 1981)

$$\begin{cases} n + \nu_e \rightarrow p + e^- \\ p + \bar{\nu}_e \rightarrow n + e^+ \end{cases}$$

Neutrino pair annihilations
(Goodman et al. 1987)

$$\nu_l + \bar{\nu}_l \rightarrow e^- + e^+$$

Neutrino Processes (2)

Example of Heating Processes: Neutrino Pair Annihilation

$$\dot{Q}(C) = \frac{KG_F^2}{2\pi c} \int \int dx_A^3 dx_B^3 \frac{(1 - \cos \psi)^2}{d_A^2 d_B^2} \int \int d\epsilon_\nu d\epsilon_{\bar{\nu}} \epsilon_\nu \epsilon_{\bar{\nu}} (\epsilon_\nu + \epsilon_{\bar{\nu}}) F_\nu F_{\bar{\nu}} (1 - f_e^-)(1 - f_e^+)$$

K is $(1 - 4 \sin^2 \theta_W + 8 \sin^4 \theta_W)/(6\pi)$ for $\nu_\mu \nu_{\bar{\mu}}$ and $\nu_\tau \nu_{\bar{\tau}}$ annihilations [erg/cc/s]

$(1 + 4 \sin^2 \theta_W + 8 \sin^4 \theta_W)/(6\pi)$ for $\nu_e \nu_{\bar{e}}$ annihilation

$F_{\nu, \bar{\nu}}$ are energy spectrum [$\text{cm}^{-3} \text{s}^{-1} \text{GeV}^{-1}$] of (anti-)neutrinos

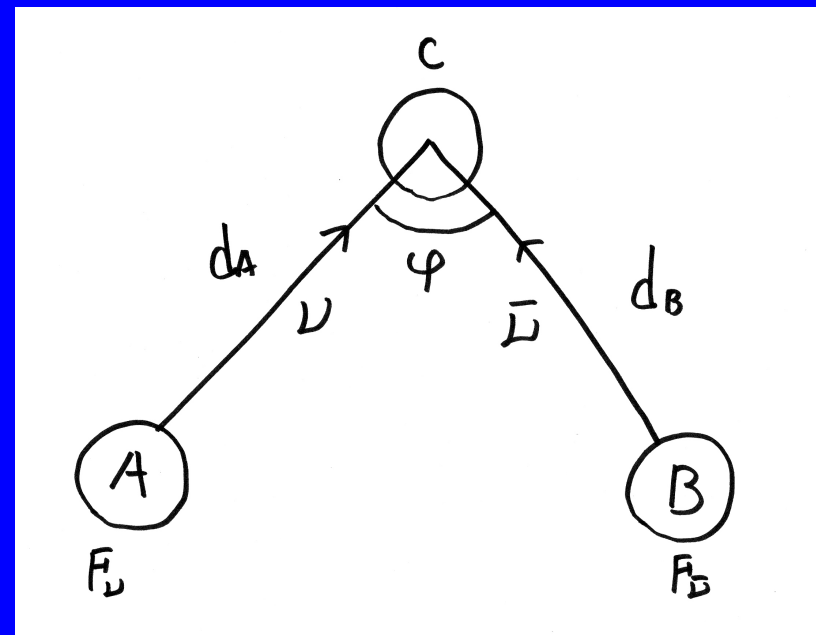
f_e^\pm are fermi blocking factors of electrons/positrons at point C.

8 dimensional integration!



We approximated as $(1 - \cos \psi)^2 \sim 1$

This approximation sets upper limit for the Effects of neutrino heating through neutrino Pair annihilation.



Magnetic Fields

Initial Condition

$$\begin{aligned}\vec{B}(\vec{r}) &= \frac{1}{3}B_0 \left(\frac{r_0}{r}\right)^3 (2 \cos \theta \vec{e}_r + \sin \theta \vec{e}_\theta) \quad \text{for } r \geq r_0 \\ &= \frac{2}{3}B_0(\cos \theta \vec{e}_r - \sin \theta \vec{e}_\theta) \quad \text{for } r < r_0.\end{aligned}$$

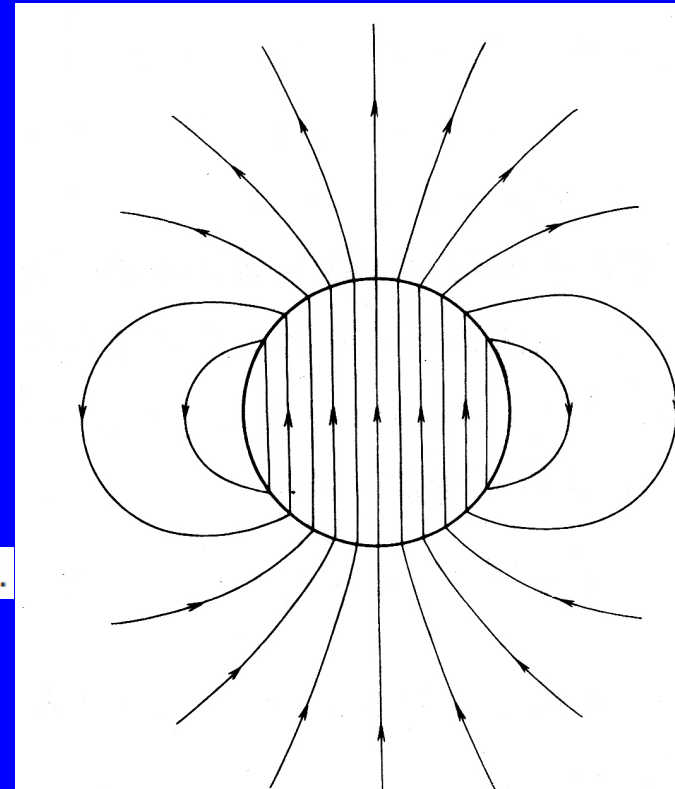
We set r_0 to be the boundary between CO core/He layer
($= 3.6 \times 10^9$ cm)

B_0 corresponds to the strength of the magnetic field in the sphere.

B_0 to be 0, 10^8 G, 10^9 G, and 10^{10} G.

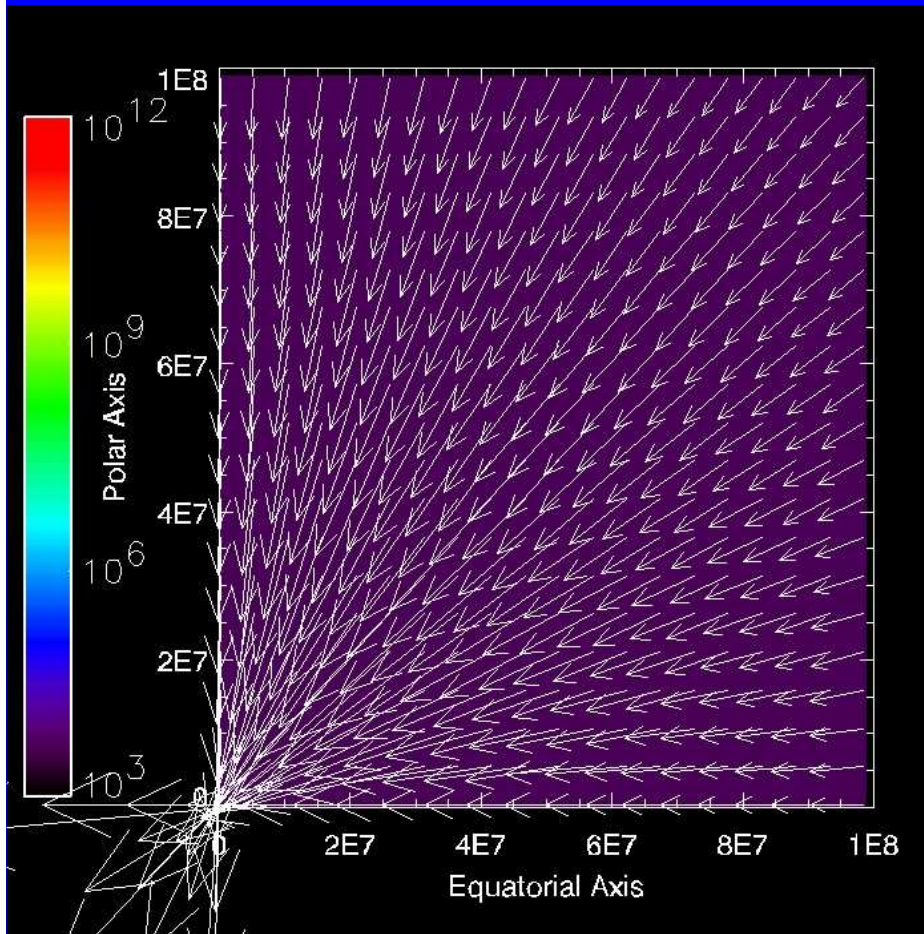
Initial |Em/W| are 0, 5E-8, 5E-6, 5E-4.

Initial |T/W| is 8E-3.

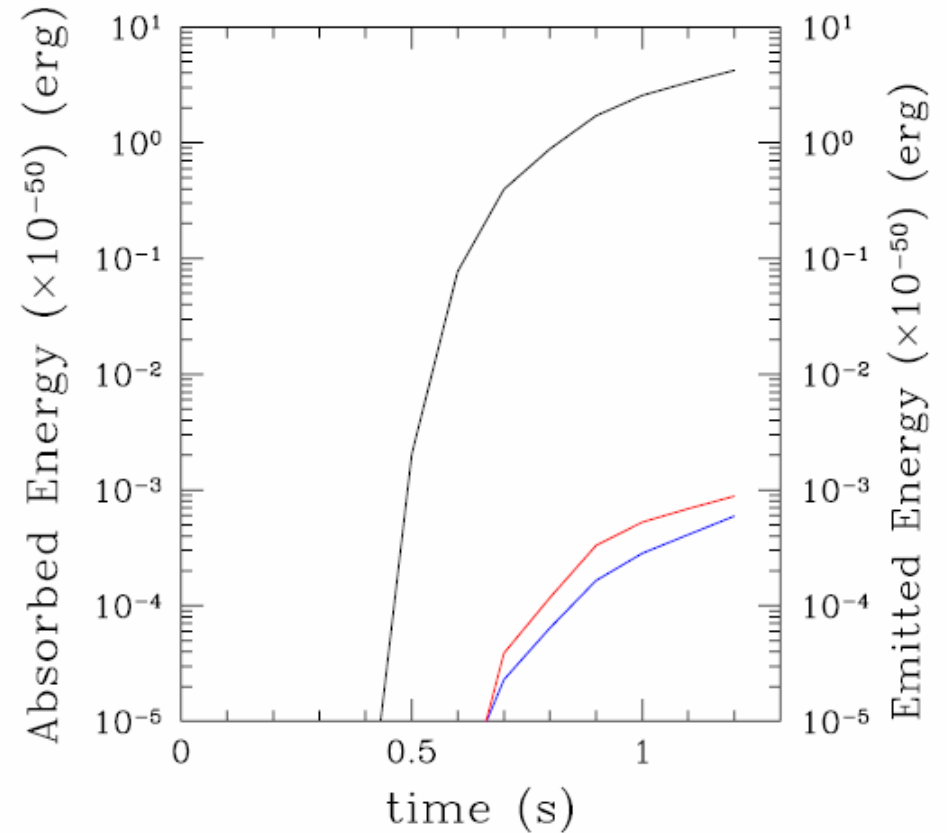


§ Results

Case of Neutrino Heating (no B-Field)



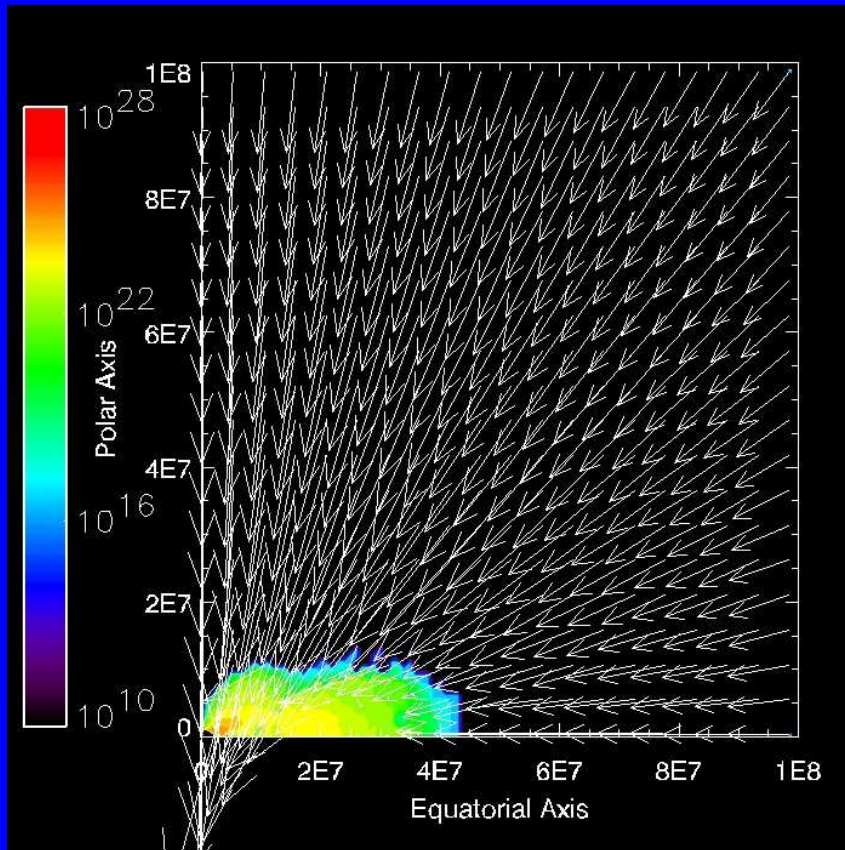
Density contour, Final time is 1.3sec.



Black: Neutrino Cooling
Red: Neutrino Pair Annihilation
Blue: Neutrino Absorption

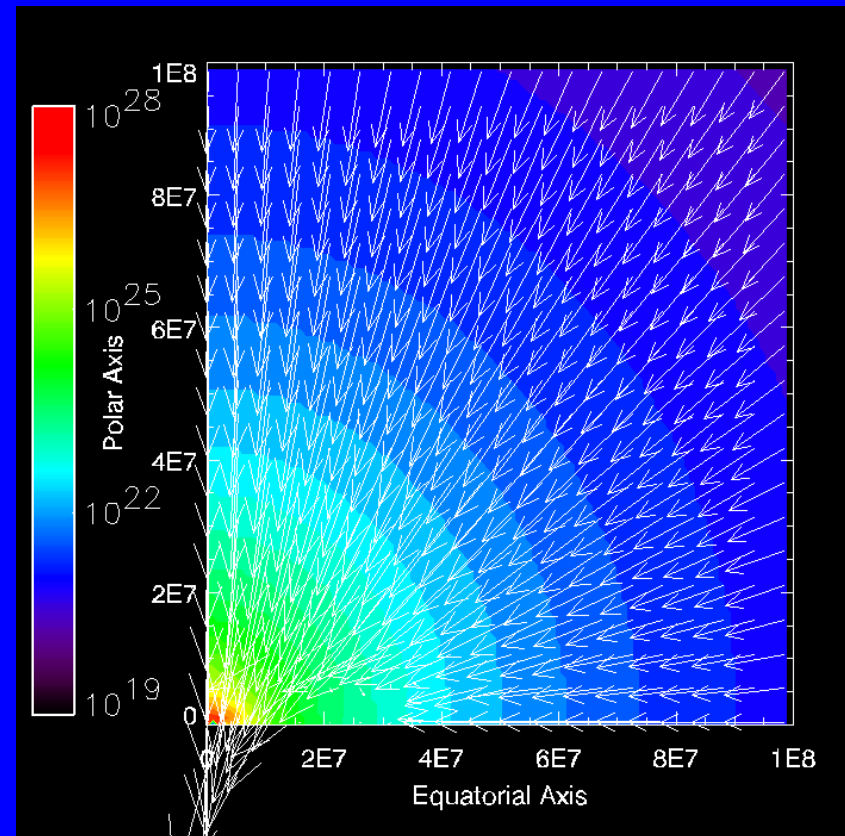
Case of Neutrino Heating (no B-Field)

Local, High Energy Deposition Rate



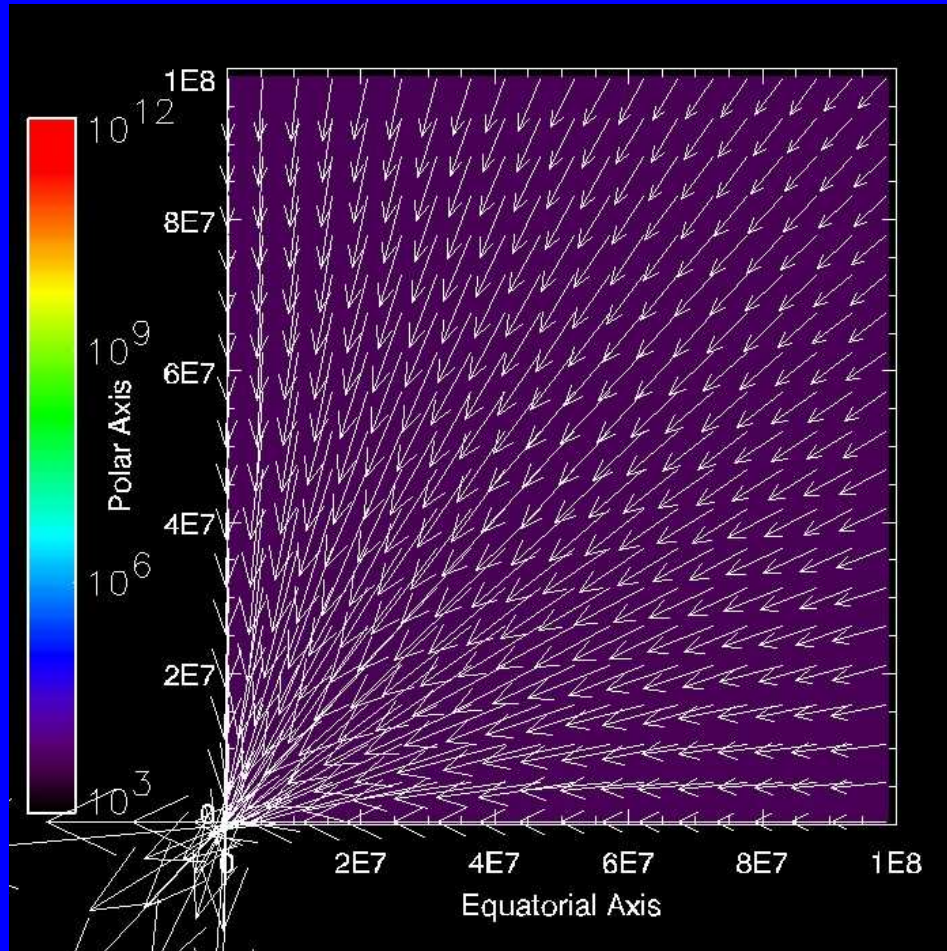
Energy Deposition Rate [erg/s/cc]
By neutrino absorption on free nucleons
At $t = 1.3$ sec.

Global, Low Energy Deposition Rate



Energy Deposition Rate [erg/s/cc]
By neutrino pair annihilation
at $t = 1.3$ sec.

Case of $B=10^9$ G (no neutrino heating)



Density contour, Final time is 1.63sec.

Jet energy at $t=1.63$ sec is $1E+47$ erg.
Def: Total energy (Kin+th+mag-Grav)
For $\theta < 10$ degree.
Mass of Jet is $2E+28$ g.



This jet is non-relativistic jet at Present.

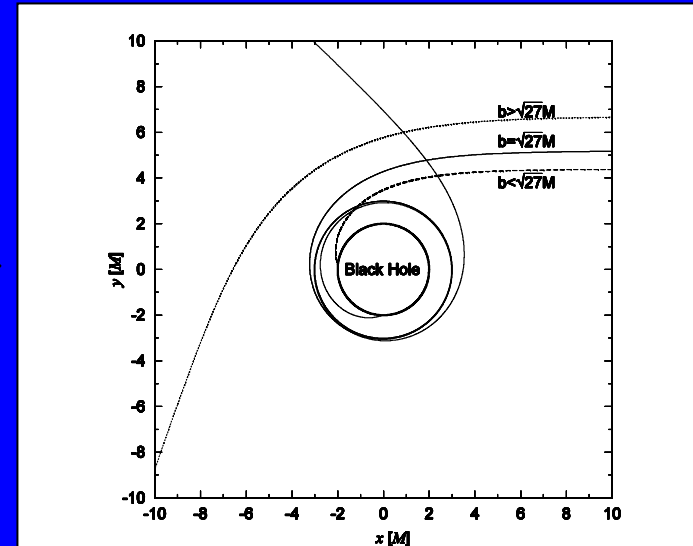
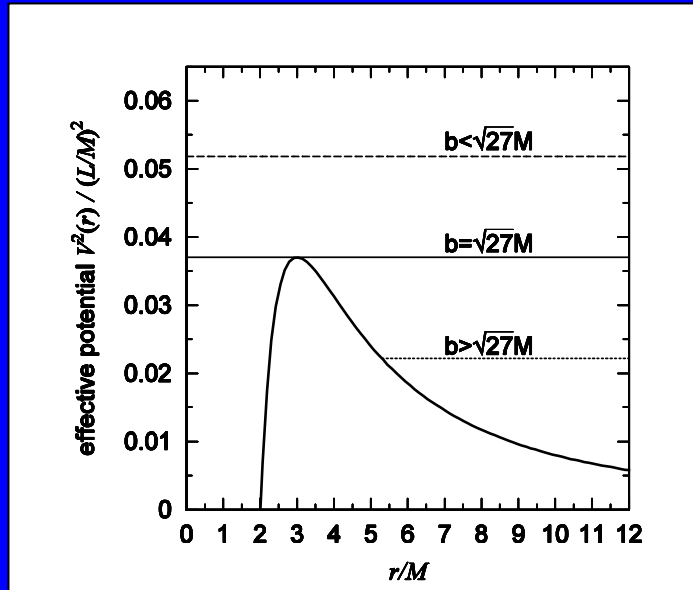
§ Discussion

Effects of Neutrino Pair Annihilation

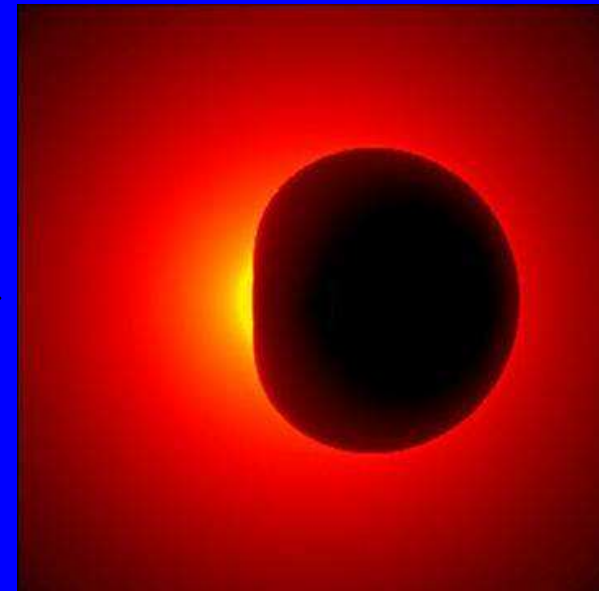
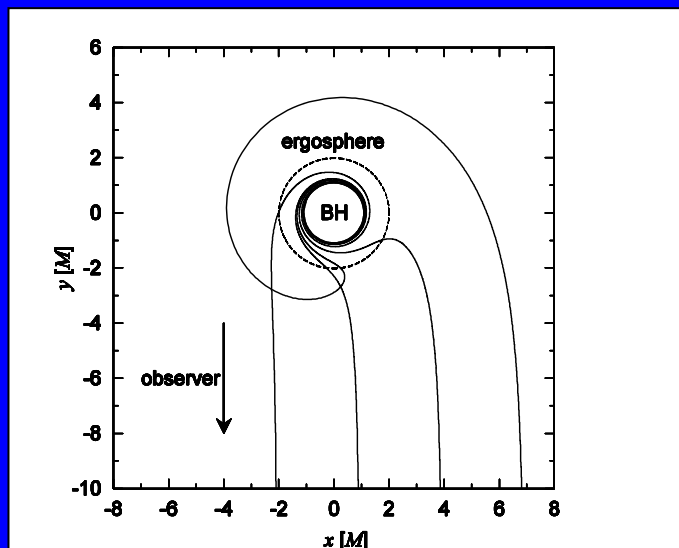
with **General Relativity**

Takahashi and Nagataki in prep.

Effective Potential

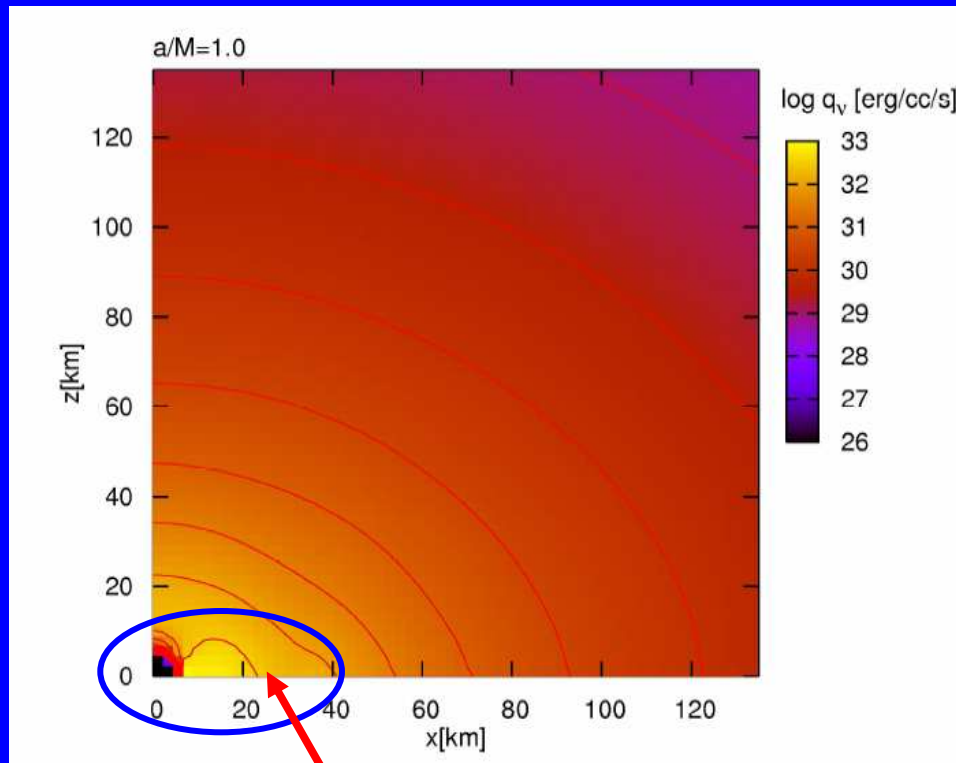


Geodesic of Neutrinos

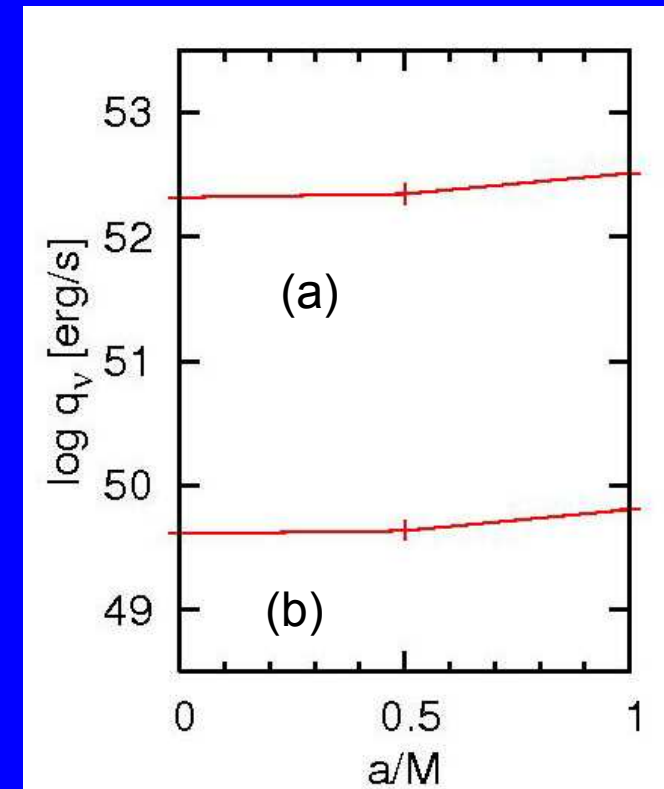


Preliminary results of Takahashi and Nagataki

Energy deposition rate due to neutrino
Pair annihilation



Energy deposition rate is enhanced because neutrinos are trapped in the Gravitational potential well.



Curve (a): Total energy depositon rate.
Curve (b): The same with curve (a),
But temperature of the accretion disk is
Artificially set to be lower by a factor of 2.

BZ process should be also important.

§ Summary

Summary and Future Work

- Neutrino heating processes and magnetic fields are included in the collapsar model.
- It is found that neutrino heating processes are insufficient to launch a jet in this study.
- A Jet is launched by magnetic fields, although this jet is non-relativistic at present.
- Effects of general relativity should be important for the formation of relativistic jet (Efficiency of neutrino heating, Blandford-Znajek effect), which we are planning to include in the near future.