## Opportunities with polarized protons at Fermilab

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Dark energy

$$
\left|f_{1 T}^{\perp}\right|_{D I S}=-\left.f_{1 T}^{\perp}\right|_{D Y}
$$

## Current Facilities

- $\mathrm{T} \& \mathrm{~L}$ polarized $p$ beams $\left(V_{s}=200,500 \mathrm{GeV}\right)$
- L program:
$\rightarrow \mathrm{A}_{\mathrm{LL}}{ }^{\pi 0}$ (PHENIX) \& $\mathrm{A}_{\mathrm{LL}}{ }^{\text {jet }}($ STAR $) \rightarrow \Delta \mathbf{g}(\mathbf{x})$
- first significant non-zero results on $\Delta g(x)$
$\rightarrow A_{L}{ }^{W \pm}$ at $V_{s}=500 \mathrm{GeV} \rightarrow \Delta \mathbf{q}_{\mathrm{bar}}(\mathrm{x})$
- surprise: $\Delta \bar{u}-\Delta \bar{d}>0$
- T program:
$\rightarrow \mathrm{A}_{N}{ }^{\pi 0, \eta, \text {,jet }, \ldots} \rightarrow$ Sivers/Collins/Twist-3



120 GeV p from Main Injector on $\mathrm{LH}_{2}, \mathrm{LD}_{2}$, $\mathrm{C}, \mathrm{Ca}, \mathrm{W}$ targets $\rightarrow$ high-x Drell-Yan

- Science data started in March 2014
$\rightarrow$ run for 2 yrs


## COMPASS-II

- $190 \mathrm{GeV} \pi^{-}$beam on T-pol H target $\rightarrow$ polarized Drell-Yan
- Data collection started in May 2015
$\rightarrow$ run for 100 days


## Future Spin Measurements

New instrumentation in forward direction
$\rightarrow$ higher $\eta$ : higher $x_{\text {beam, }}$ lower $x_{\text {target }}$

- STAR Forward Calorimeter System: EMCal + HCal
- fsPHENIX: forward spectometer w/ EMCal, HCal, RICH, tracking
$\rightarrow$ planned spin program in $\Delta \mathbf{g}\left(\mathbf{x}, \mathrm{Q}^{2}\right)$ at low-x
(longitudinal) as well as Jets, Drell-Yan (transverse), ...

Polarized Beam and/or Target wl SeaQuest detector
$\rightarrow$ high-Iuminosity facility for polarized Drell-Yan

- E-1039: SeaQuest w/ pol $\mathrm{NH}_{3}$ target
$\rightarrow$ probe sea quark distributions
- E-1027: pol p beam on unpol tgt
$\rightarrow$ Sivers sign change (valence quark)


## TMDs: Sivers Function



## cannot exist w/o quark OAM

- describes transverse-momentum distribution of unpolarized quarks inside transversely polarized proton
- captures non-perturbative spin-orbit coupling effects inside a polarized proton
- Sivers function is naïve time-reversal odd
- leads to
$\rightarrow \sin \left(\phi-\phi_{\mathrm{S}}\right)$ asymmetry in SIDIS
$\rightarrow \sin \phi_{\mathrm{b}}$ asymmetry in Drell-Yan
- measured in SIDIS (HERMES, COMPASS)
- future measurements at Jlab@12 GeV planned


First moment of Sivers functions:
$\rightarrow$ u- and d- Sivers have opposite signs, of roughly equal magnitude

## Sivers Asymmetry in SIDIS



- Global fit to $\sin \left(\phi_{\mathrm{h}}-\phi_{S}\right)$ asymmetry in SIDIS (HERMES (p), COMPASS (p), COMPASS (d))

COMPASS (p)


COMPASS (d)


## QCD Evolution of Sivers Function




- Initial global fits by Anselmino group included DGLAP evolution only in collinear part of TMDs (not entirely correct for TMD-factorization)
- Using TMD Q ${ }^{2}$ evolution:
$\longrightarrow$ agreement with data improves



## TMD Evolution of Sivers Asymmetry (W)



- much stronger than any other known evolution effects
- but needs input from data to constrain nonpertubative part in evolution
- Can only be done at RHIC (plans for 2\% measurement in 2016)

Comparison of extracted TMD (Sivers) will provide strong constraint on TMD evolution

## The Sign Change

$$
\left.f_{1 T}^{\perp}\left(x, k_{T}\right)\right|_{\text {SIDIS }}=-\left.f_{1 T}^{\perp}\left(x, k_{T}\right)\right|_{D Y, W}
$$

- fundamental prediction of QCD (in non-perturbative regime)
$\rightarrow$ goes to heart of gauge formulation of field theory
- "Smoking gun" prediction of TMD formalism
- Universality test includes not only the sign-reversal character of the TMDs but also the comparison of the amplitude as well as the shape of the corresponding TMDs
- NSAC Milestone HP13 (2015):
"Test unique QCD predictions for relations between single-transverse spin phenomena in p-p scattering and those observed in deep-inelastic lepton scattering"


## Planned Polarized Drell-Yan Experiments

| Experiment | Particles | $\begin{gathered} \text { Energy } \\ \text { (GeV) } \end{gathered}$ | $\mathrm{x}_{\mathrm{b}}$ or $\mathrm{x}_{\mathrm{t}}$ | $\underset{\left(\mathrm{cm}^{-2} \mathrm{~s}^{-1}\right)}{\text { Luminosity }}$ | $A_{T}^{\sin \phi_{s}}$ | $\mathrm{P}_{\mathrm{b}}$ or $\mathrm{P}_{\mathrm{t}}(\mathrm{f})$ | rFOM ${ }^{*}$ | Timeline |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| COMPASS (CERN) | $\pi^{ \pm}+p^{\uparrow}$ | $\begin{aligned} & 160 \mathrm{GeV} \\ & \sqrt{s}=17 \end{aligned}$ | $\mathrm{x}_{\mathrm{t}}=0.1-0.3$ | $2 \times 10^{33}$ | 0.14 | $\begin{gathered} P_{t}=90 \% \\ f=0.22 \end{gathered}$ | $1.1 \times 10^{-3}$ | 2015, 2018 |
| PANDA (GSI) | $\overline{\mathbf{p}}+\mathbf{p}^{\uparrow}$ | $\begin{aligned} & 15 \mathrm{GeV} \\ & V_{\mathrm{s}}=5.5 \end{aligned}$ | $\mathrm{x}_{\mathrm{t}}=0.2-0.4$ | $2 \times 10^{32}$ | 0.07 | $\begin{aligned} P_{t} & =90 \% \\ f & =0.22 \end{aligned}$ | $1.1 \times 10^{-4}$ | >2018 |
| $\begin{aligned} & \text { PAX } \\ & \text { (GSI) } \end{aligned}$ | $\mathbf{p}^{\uparrow}+\bar{p}$ | collider $V_{s}=14$ | $\mathrm{x}_{\mathrm{b}}=0.1-0.9$ | $2 \times 10^{30}$ | 0.06 | $\mathrm{P}_{\mathrm{b}}=90 \%$ | $2.3 \times 10^{-5}$ | >2020? |
| NICA <br> (JINR) | $p^{\uparrow}+\mathbf{p}$ | collider $V_{s}=26$ | $\mathrm{x}_{\mathrm{b}}=0.1-0.8$ | $1 \times 10^{31}$ | 0.04 | $\mathrm{P}_{\mathrm{b}}=70 \%$ | $6.8 \times 10^{-5}$ | >2018 |
| PHENIX/STAR (RHIC) | $\mathbf{p}^{\uparrow}+\mathbf{p}^{\uparrow}$ | collider $V_{s}=510$ | $\mathrm{x}_{\mathrm{b}}=0.05-0.1$ | $2 \times 10^{32}$ | 0.08 | $\mathrm{P}_{\mathrm{b}}=60 \%$ | $1.0 \times 10^{-3}$ | >2018 |
| fsPHENIX (RHIC) | $\mathbf{p}^{\uparrow}+\mathbf{p}^{\uparrow}$ | $\begin{aligned} & V_{s}=200 \\ & V_{s}=510 \end{aligned}$ | $\begin{gathered} x_{b}=0.1-0.5 \\ x_{b}=0.05-0.6 \end{gathered}$ | $\begin{aligned} & 8 \times 10^{31} \\ & 6 \times 10^{32} \end{aligned}$ | 0.08 | $\begin{aligned} & P_{b}=60 \% \\ & P_{b}=50 \% \end{aligned}$ | $\begin{aligned} & 4.0 \times 10^{-4} \\ & 2.1 \times 10^{-3} \end{aligned}$ | >2021 |
| SeaQuest <br> (FNAL: E-906) | $p+p$ | $\begin{aligned} & 120 \mathrm{GeV} \\ & \sqrt{s}=15 \end{aligned}$ | $\begin{aligned} x_{b} & =0.35-0.9 \\ x_{t} & =0.1-0.45 \end{aligned}$ | $3.4 \times 10^{35}$ | --- | --- | --- | 2012-2016 |
| Pol tgt DY ${ }^{\ddagger}$ <br> (FNAL: E-1039) | $p+p^{\uparrow}$ | $\begin{aligned} & 120 \mathrm{GeV} \\ & V_{\mathrm{s}}=15 \end{aligned}$ | $\mathrm{x}_{\mathrm{t}}=0.1-0.45$ | $4.4 \times 10^{35}$ | $\begin{gathered} 0- \\ 0.2^{*} \end{gathered}$ | $\begin{gathered} P_{t}=85 \% \\ f=0.176 \end{gathered}$ | 0.15 | 2017-2018 |
| Pol beam DY ${ }^{\S}$ <br> (FNAL: E-1027) | $p^{\uparrow}+p$ | $\begin{aligned} & 120 \mathrm{GeV} \\ & \sqrt{\mathrm{~s}}=15 \end{aligned}$ | $\mathrm{x}_{\mathrm{b}}=0.35-0.9$ | $2 \times 10^{35}$ | 0.04 | $P_{\text {b }}=60 \%$ | 1 | >2018 |

[^0]
## Sivers Asymmetry at Fermilab Main Injector

- Experimental Sensitivity
$\rightarrow$ luminosity: $\mathrm{L}_{\mathrm{av}}=2 \times 10^{35}$ ( $10 \%$ of available beam time: $\mathrm{I}_{\mathrm{av}}=15 \mathrm{nA}$ )
$\rightarrow 3.2 \times 10^{18}$ total protons for $5 \times 10^{5} \mathrm{~min}:\left(=2 \mathrm{yrs}\right.$ at $50 \%$ efficiency) with $\mathrm{P}_{\mathrm{b}}=60 \%$


Note:

$$
A_{N}=\frac{2}{\pi} A_{T U}^{\sin \phi_{b}}
$$

$\rightarrow$ Can measure not only sign, but also the size \& maybe shape of the Sivers function!

## Polarized Beam Drell-Yan at Fermilab (E-1027)

- Extraordinary opportunity at Fermilab (best place for polarized DY) :
$\longrightarrow$ high luminosity, large x-coverage
$\rightarrow$ (SeaQuest) spectrometer already setup and running
$\rightarrow$ run alongside neutrino program (w/ 10\% of beam)
$\longrightarrow$ experimental sensitivity:
> 2 yrs at $50 \%$ eff, $\mathrm{P}_{\mathrm{b}}=60 \%, \mathrm{I}_{\mathrm{av}}=15 \mathrm{nA}$
$>$ luminosity: $\mathrm{L}_{\mathrm{av}}=2 \times 10^{35} / \mathrm{cm}^{2} / \mathrm{s}$
> measure sign, size \& shape of Sivers function



## A Novel, Compact Siberian Snake for the Main Injector

Single snake design (6.4m long):

- 1 helical dipole +2 conv. dipoles
- helix: 4T / $5.6 \mathrm{~m} / 4$ " ID
- dipoles: 4T / 0.2 m/4" ID
- use 4-twist magnets
- $8 \pi$ rotation of $B$ field
- never done before in a high energy ring
- RHIC uses snake pairs
- 4 single-twist magnets ( $2 \pi$ rotation)

initial design studies
Siberian Snake B-Fields


Excursion vs. Length
 beam energy

## Differences compared to RHIC

- Most significant difference: Ramp time of Main Injector < 0.7 s , at RHIC 1-2 min
$\rightarrow$ warm magnets at MI vs. superconducting at RHIC
$\rightarrow$ pass through all depolarizing resonances much more quickly
- Beam remains in $\mathrm{MI} \sim 5 \mathrm{~s}$, in RHIC $\sim 8$ hours
$\rightarrow$ extracted beam vs. storage ring
$\rightarrow$ much less time for cumulative depolarization
- Disadvantage compared to RHIC — no institutional history of accelerating polarized proton beams
$\rightarrow$ Fermilab E704 had polarized beams through hyperon decays



## The Path to a polarized Main Injector

Stage 1 approval from Fermilab: 14-November-2012

- PAC request: detailed machine design and costing using 1 snake in MI
$\rightarrow$ Spin@Fermi collaboration provide design
$\rightarrow$ Fermilab (AD) does verification \& costing
- Collaboration with A.S. Belov at INR and Dubna to develop polarized source
- Initial simulations in 2013-2014:
$\rightarrow$ set up Zgoubi spin-tracking package (M. Bai, F. Meot, BNL, M. Syphers, MSU)
$\longrightarrow$ single particle tracking, emittance, momentum spread of particles
$\rightarrow$ conceptual design that works at least for a perfect machine - perfect magnet alignment, perfect orbits, no momentum spread, etc.
$\longrightarrow$ but slow and limited support:
difficulties implementing orbit errors, quadrupole mis-alignments/rolls, ramp rates


## Effect of emittance on final polarization vs Energy



Point-like snake in correct location, perfect orbit, no momentum smearing.

Average polarization for 8 particles

Only small difference seen at final energy of 120 GeV

## The Path to a polarized Main Injector - II

Breakthrough: AD support from Fermilab: July-2015

- Fermilab AD support in 2015
$\rightarrow$ S. Nagaitsev pledges support for simulations (April 2015)
$\rightarrow$ Meiqin Xiao from AD set up PTC (Etienne Forest, KEK)
$\rightarrow$ repeated Zgoubi work in 1 month
$\rightarrow$ "easy" to include orbit errors, quadrupole mis-alignments/rolls, ramp rates
$\rightarrow$ support for one year
$\rightarrow$ plan to complete simulations
$\rightarrow$ go back to PAC



## Polarized Traget Drell-Yan at Fermilab (E-1039)

- Probe Sea-quark Sivers Asymmetry with a polarized proton target at SeaQuest

- Statistics shown for one calendar year of running:
$-\mathrm{L}=7.2 * 10^{42} / \mathrm{cm}^{2} \leftrightarrow \mathrm{POT}=2.8^{*} 10^{18}$
- Running will be two calendar years of beam time
- existing SIDIS data poorly constrain sea-quark Sivers function
- significant Sivers asymmetry expected from meson-cloud model
- first Sea Quark Sivers Measurement
- determine sign and value of ū Sivers distribution

If $A_{N} \neq 0$, major discovery:
"Smoking Gun" evidence for $L_{\bar{u}} \neq 0$

## Status and Plans (E-1039)

## Target

Polarization: 85\%
Packing fraction 0.6
Dilution factor: 0.176
Density: $0.89 \mathrm{~g} / \mathrm{cm}^{3}$


- use current SeaQuest setup, a polarized proton target, unpolarized beam
- add third magnet SMO $\sim 5 \mathrm{~m}$ upstream
$\rightarrow$ improves dump-target separation
$\rightarrow$ reduces overall acceptance
- Current status
- magnet system is finished and working
- refrigerator is finished and tested (at 1 K )
- NMR system reached final design
- mechanical design laid out
ahead of schedule, ready for installation in Fall 2016


## Exploring the Dark Side of the Universe



- Dark sector could interact with the standard model sector via a hidden gauge boson (A' or "dark photon" or "para photon" or "hidden photon")
- Dark photons can provide a portal into the dark sector
- Dark photons could couple to standard model matter with $\alpha^{\prime}=\alpha \varepsilon^{2}$
B. Holdom, PLB 166 (1986) 196
J. D. Bjorken et al, PRD 80 (2009) 075018
$\varepsilon \sim 10^{-2}$ to $10^{-8}$ from loops of heavy particles


## Possible Mechanisms for producing A' at SeaQuest

- Proton Bremsstrahlung
- $\pi^{0}, \eta \ldots$ decay



## SeaQuest A' search strategy

- $A^{\prime}$ generated by $\eta$ decay and/or proton
Bremsstrahlung in the Iron beam dump
- A' could travel a distance $I_{0}$ without interacting
- A' decays into di-leptons
- Reconstructed di-lepton vertex is displaced, downstream of the target in the beam dump



## A' sensitivity region for SeaQuest

$l_{o} \approx \frac{0.8 \mathrm{~cm}}{N_{e f f}}\left(\frac{E_{o}}{10 \mathrm{GeV}}\right)\left(\frac{10^{-4}}{\varepsilon}\right)^{2}\left(\frac{100 \mathrm{MeV}}{m_{A}}\right)^{2}$
J.D. Bjorken et al, PRD 80 (2009) 075018

- $E_{0}=$ energy of the $\mathrm{A}^{\prime}$
$\rightarrow E_{0}=5-20 \mathrm{GeV}$ for $\eta$ decay
$\rightarrow \mathrm{E}_{0}=5-110 \mathrm{GeV}$ for p bremsstrahlung
- $\quad N_{\text {eff }}=$ no. of avail. decay products
$\rightarrow \mathrm{N}_{\text {eff }}=2$
- $I_{0}=$ distance that $A^{\prime}$ travels before decaying

$$
\rightarrow I_{0}=0.17 \mathrm{~m}-5.95 \mathrm{~m}
$$

- $\varepsilon=$ coupling constant between standard model and dark sector
- $m_{A^{\prime}}=$ mass of $\mathrm{A}^{\prime}$


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$\rightarrow N_{\text {eff }}=2$
- $I_{0}=$ distance that $A^{\prime}$ travels before decaying

$$
\rightarrow I_{0}=0.17 m-5.95 m
$$

- $\varepsilon=$ coupling constant between standard model and dark sector
- $m_{A^{\prime}}=$ mass of $\mathrm{A}^{\prime}$


DY-like: can access A' with larger mass

## Polarized Proton Beams

## and Searches for Dark Forces

Searches for a dark photon also limit other possibilities

## Parity violation studies could prove key

$$
\mathcal{L}_{\text {darkZ }}=-\left(\varepsilon e J_{\mathrm{em}}^{\mu}+\varepsilon_{Z} \frac{g}{2 \cos \theta_{W}} J_{\mathrm{NC}}^{\mu}\right) Z_{d \mu}
$$

[Davoudiasl, Lee, Marciano, 2OI4]
If the $A^{\prime}$ is a dark $Z$, then ...


The dilepton yield can change with proton polarization: the asymmetry can be $\mathrm{O}(\mathrm{r})$ !
[SG, Holt, Tadepalli, 2015]

## Summary

- There are many exiting opportunities with polarized hadron beams in the coming decade
- RHIC, Fermilab, COMPASS offer complementary probes and processes to study hadronic landscape
$\longrightarrow$ a complete spin program requires multiple hadron species
- Hope to answer some of the burning questions
$\longrightarrow$ how much do the gluons contribute to the nucleon spin?
$\longrightarrow$ is there significant orbital angular momentum?
$\longrightarrow$ does TMD formalism work? Does Sivers function change sign?
- Explore the Dark Sector
$\longrightarrow$ SeaQuest is nearly ideal beam-stop experiment
$\longrightarrow$ underway for at least the next year
$\longrightarrow$ probe not only dark photons, but also $Z_{d}$ with a polarized beam


## Thank You

## COMPASS, E-1027, E-1039 (and Beyond)

|  | Beam Pol. | Target Pol. | Favored Quarks | Physics Goals |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | (Sivers Function) |  |  | $L_{\text {sea }}$ | $A^{\prime}, Z_{\text {d }}$ |
|  |  |  |  | sign change | size | shape |  |  |
| COMPASS <br> $\pi^{-} p^{\uparrow} \rightarrow \mu^{+} \mu^{-} X$ | $X$ | $\checkmark$ | valence | $\checkmark$ | $X$ | $X$ | $X$ | $X$ |
| $\begin{gathered} \mathrm{E}-1027 \\ p^{\uparrow} p \rightarrow \mu^{+} \mu^{-} X \end{gathered}$ | $\checkmark$ | X | valence | $\checkmark$ | $\checkmark$ | $\checkmark$ | $X$ | $\checkmark$ |
| $\begin{gathered} \mathrm{E}-1039 \\ p^{\uparrow} \rightarrow \mu^{+} \mu^{-} X \end{gathered}$ | X | $\checkmark$ | sea | X | $\checkmark$ | $(\sqrt{ })$ | $\checkmark$ | $(\sqrt{ })$ |
| E-10XX | $\sqrt{ }$ | $\sqrt{ }$ | sea \& valence | Transversity, Helicity, Other TMDs ... |  |  |  |  |
| $\vec{p} \vec{p} \rightarrow \mu^{+} \mu^{-} X$ |  |  |  |  |  |  |  |  |  |  |

## Polarized Beam at Fermilab Main Injector

- Polarized Beam in Main Injector
$\rightarrow$ use SeaQuest target
$\sqrt{ }$ liquid $\mathrm{H}_{2}$ target can take about $\mathrm{I}_{\mathrm{av}}=5 \times 10^{11} \mathrm{p} / \mathrm{s}(=80 \mathrm{nA})$
$\rightarrow 1 \mathrm{~mA}$ at polarized source can deliver about $\mathrm{t}_{\mathrm{av}}=1 \times 10^{12} \mathrm{p} / \mathrm{s}(=150 \mathrm{nA})$ for 100\% of available beam time (A. Krisch: Spin@Fermi report in (Aug 2011): arXiv:1110.3042 [physics.acc-ph])
$26 \mu$ s linac pulses, 15 Hz rep rate, 12 turn injection into booster, 6 booster pulses into Recycler Ring, followed by 6 more pulses using slip stacking in MI

1 MI pulse $=1.9 \times 10^{12} \mathrm{p}$
using three 2 -sec cycles $/ \mathrm{min}$ ( $\sim 10 \%$ of beam time):
$\rightarrow 2.8 \times 10^{12} \mathrm{p} / \mathrm{s}(=450 \mathrm{nA})$ instantaneous beam current , and $\mathrm{I}_{\mathrm{av}}=0.95 \times 10^{11} \mathrm{p} / \mathrm{s}(=15 \mathrm{nA})$
$\rightarrow$ possible scenarios:

$$
\begin{array}{ll}
\mathrm{L}_{\mathrm{av}}=2.0 \times 10^{35} / \mathrm{cm}^{2} / \mathrm{s} & \left(10 \% \text { of available beam time: } \mathrm{I}_{\mathrm{av}}=15 \mathrm{nA}\right) \\
\mathrm{L}_{\mathrm{av}}=1 \times 10^{36} / \mathrm{cm}^{2} / \mathrm{s} & \left(50 \% \text { of available beam time: } \mathrm{I}_{\mathrm{av}}=75 \mathrm{nA}\right)
\end{array}
$$

$\rightarrow$ Systematic uncertainty in beam polarization measurement (scale uncertainty)

$$
\Delta \mathrm{P}_{\mathrm{b}} / \mathrm{P}_{\mathrm{b}}<5 \%
$$

## Dark Photons at SeaQuest (FNAL)

[SG, Holt, Tadepalli, arXiv:1509.00050]


## Dark Photons: SeaQuest vs. SHiPS

 "apples \& oranges"

## 5 yr exposure 400 GeV beam opt. detectors VS.

I yr exposure
120 GeV beam SeaQuest spect.
Sharper constraints are possible!


[^0]:    

    * not constrained by SIDIS data / \#rFOM = relative lumi * $\mathbf{p}^{2}{ }^{*} \mathbf{f}^{2}$ wrt $\mathrm{E}-1027$ ( $\mathrm{f}=1$ for pol p beams, $\mathrm{f}=0.22$ for $\pi^{-}$beam on $\mathrm{NH}_{3}$ )

