

HW 6: Selections. *Abbers*

The first order Stark effect shift for these levels is

$$\Delta^{(1)} = 3eaE = \frac{3e^2}{\beta^2 a} = \frac{3\alpha^2 m}{\beta^2} \quad (S7.86)$$

Write the second-order shift from parts (a) and (b) as

$$\Delta^{(2)} = -\gamma \frac{1}{\alpha^3 m^3} E^2 = -\gamma \frac{\alpha^2 m}{\beta^4} \quad (S7.87)$$

Evidently for the ground state,  $\gamma$  is about 2. The ratio of the second order ground state splitting to the first order  $n = 2$  splitting is about

$$\left| \frac{\Delta^{(2)}}{\Delta^{(1)}} \right| = \frac{\gamma}{3\beta^2} \quad (S7.88)$$

When  $\beta = 10$ , this ratio is about

$$\left| \frac{\Delta^{(2)}}{\Delta^{(1)}} \right| \approx 0.007 \quad (S7.89)$$

Part d) When  $\beta = 2$

$$\left| \frac{\Delta^{(2)}}{\Delta^{(1)}} \right| \approx 0.16 \quad (S7.90)$$

The two effects become comparable when  $\beta \approx \sqrt{\gamma/3}$ .

### Problem 7.10 Some Magnetic Field Effects

Part a) The fine-structure splitting between the  $n = 2$  levels is

$$\Delta_{FS} = \frac{\alpha^4 m}{32} \quad (S7.91)$$

which is about  $4.523 \times 10^{-5} \text{ eV}$ .

The shift due to an external magnetic field  $B$  is of the order  $\mu_0 B$ , where  $\mu_0 = e/2m$  which is about  $5.79 \times 10^{-9}$  electron volts per gauss.<sup>2</sup> When  $B = .5$  gauss, this is much smaller than the fine structure splitting. But when  $B = 10^5$  gauss, the magnetic energy is larger than the fine-structure splitting, comparable to the spacing between the gross-structure energy levels.

Part b) The  $A^2$  term is of the order  $(e^2/m)B^2 a^2$ , since  $\tau$  is of the order of the Bohr radius  $a$ . Therefore it is comparable to the term linear in  $B$  when  $\mu_0 B = B^2 a^2 e^2/m$ , or

$$B = \frac{m\mu_0}{e^2 a^2} = \frac{m^2 \alpha^2}{e} \quad (S7.92)$$

or

$$B^2 = m^4 \alpha^3 \quad (S7.93)$$

A gauss is a cgs unit equivalent to energy<sup>2</sup> in natural units. But this isn't enough to tell you the cgs units. There are some conventions, and you have to know something else. Remember that in cgs units,  $B$  and  $E$  have the same units. You can figure it out from the Lorentz force law, equation (1.20), or from any number of books that tell you that  $\mu_0 = e\hbar/2mc$ . Remember that in cgs,  $e^2$  has the same dimensions as  $\hbar c$ . The upshot is that gauss<sup>2</sup> is a cgs unit equivalent to erg/cm<sup>3</sup>. Thus

$$\begin{aligned} B^2 &= \alpha^3 \times (511,000 \text{ eV})^4 \times 1.602189 \times 10^{-12} \frac{\text{erg}}{\text{eV}} \times \left( \frac{1}{1.97 \times 10^{-5} \text{ eVcm}} \right)^3 \\ &= 5.552 \times 10^{18} \text{ gauss}^2 \end{aligned} \quad (S7.94)$$

and the desired magnetic field is of the order  $2.3 \times 10^9$  gauss. For some other atoms and molecules, the size of  $B$  for which the  $A^2$  term becomes important turns out to be rather lower.

<sup>2</sup>See Equation (6.39).

Part c) In the strong field limit, label the states by  $|n, l, m_l, m_s\rangle$ . The energies in the field along the  $z$ -axis are (ignoring the fine structure effects)

$$E_{n,l,m_l,m_s}^0 = -\frac{\alpha^2 m}{2n^2} - \mu_0 B(m_l + 2m_s) \quad (\text{S7.95})$$

so the difference between the energies of the highest and lowest states is  $4\mu_0 B$ . This is comparable to the fine-structure splitting when

$$4\mu_0 B = \Delta_{FS} = \frac{\alpha^4 m}{32} \quad (\text{S7.96})$$

or

$$B = \frac{\Delta_{FS}}{4\mu_0} = \frac{4.523 \times 10^{-5} \text{ eV}}{4 \times 5.79 \times 10^{-9} \text{ eV/gauss}} = 1953 \text{ gauss} \quad (\text{S7.97})$$

Part d) We want to compute the first order perturbed energy.

$$\Delta_{FS} = \langle \Psi_{m_l m_s}^{n l} | H_{LS} + H_{kin} | \Psi_{m_l m_s}^{n l} \rangle \quad (\text{S7.98})$$

We know the matrix elements in the  $|\Phi_{jm}^{nl}\rangle$  basis, where  $H_{LS} + H_{kin}$  is diagonal. So in the  $m_l, m_s$  basis

$$\begin{aligned} & \sum_j \langle \Psi_{m_l m_s}^{n l} | \Phi_{jm}^{n l} \rangle \langle \Phi_{jm}^{n l} | H_{FS} | \Phi_{jm}^{n l} \rangle \langle \Phi_{jm}^{n l} | \Psi_{m_l m_s}^{n l} \rangle \\ &= \frac{l \pm m + 1/2}{2l + 1} \langle \Phi_{l+\frac{1}{2}, m_l \pm \frac{1}{2}}^{n l} | H_{FS} | \Phi_{l+\frac{1}{2}, m_l \pm \frac{1}{2}}^{n l} \rangle \\ &+ \frac{l \mp m + 1/2}{2l + 1} \langle \Phi_{l-\frac{1}{2}, m_l \pm \frac{1}{2}}^{n l} | H_{FS} | \Phi_{l-\frac{1}{2}, m_l \pm \frac{1}{2}}^{n l} \rangle \end{aligned} \quad (\text{S7.99})$$

$$= \frac{\alpha^4 m}{2n^4} \left[ \frac{l \pm m + 1/2}{2l + 1} \left( \frac{3}{4} - \frac{2n}{2l + 2} \right) + \frac{l \mp m + 1/2}{2l + 1} \left( \frac{3}{4} - \frac{2n}{2l} \right) \right]$$

When  $l = 0$ , there is no  $j = l - 1/2$  state, and the second term is absent.

In the  $n = 2$  level, the energies of the three states with  $m = \frac{1}{2}$  in this limit are:

$$\begin{aligned} E(2S) &= -\frac{\alpha^2 m}{8} - \frac{5}{4} \Delta - \mu_0 B \\ E(2P)_{m_s = +\frac{1}{2}} &= -\frac{\alpha^2 m}{8} - \frac{7}{12} \Delta - \mu_0 B \quad (\text{S7.100}) \\ E(2P)_{m_s = -\frac{1}{2}} &= -\frac{\alpha^2 m}{8} - \frac{11}{12} \Delta \end{aligned}$$

### Problem 7.11 The van der Waals Effect

The problem is simplified if you exploit the remaining axial symmetry. Write

$$\mathbf{r}_1 \cdot \mathbf{r}_2 - 3z_1 z_2 = x_1 x_2 + y_1 y_2 - 2z_1 z_2 \quad (\text{S7.101})$$

In spherical-tensor notation.

$$r_{\pm} = \mp \frac{x \pm iy}{\sqrt{2}} \quad r_0 = z \quad (\text{S7.102})$$

so that

$$H' = -\frac{e^2}{R^3} (\tau_1^+ \tau_2^- + \tau_1^- \tau_2^+ + 2\tau_1^0 \tau_2^0) \quad (\text{S7.103})$$

Now let us look at the  $m = 2, 1$ , and 0 separately:

Since the variational estimate of the ground state energy is always an upper bound,

$$\Delta < -2E^2a^3 \quad (\text{S7.203})$$

In problem 7.9, with three terms in the perturbation series we found

$$\Delta < -1.746a^3E^2 \quad (\text{S7.204})$$

So the variational method does better than that. It is still above the lower bound from equation (7.47):

$$\Delta > ge - \frac{8}{3}a^3E^2 = -2.6666\dots a^3E^2 \quad (\text{S7.205})$$

Our estimate of course does not contradict the exact second order answer (7.59):

$$\Delta = -\frac{9}{4}a^3E^2 = -2.25a^3E^2 \quad (\text{S7.206})$$

### Problem 7.17 Born-Oppenheimer Approximation—A Simple Model

Part a)

$$V = \frac{\beta}{2} [(x_2 - x_3 - a)^2 + (x_2 - x_1 + a)^2] = \beta(x_2 - R)^2 + \beta\left(\frac{r}{2} - a\right)^2 \quad (\text{S7.207})$$

where

$$R = \frac{x_1 + x_3}{2} \quad \text{and} \quad r = x_1 - x_3 \quad (\text{S7.208})$$

are the nuclear cm and relative coordinates. If we fix  $R$  and  $r$ , this is just a harmonic oscillator with the zero of the energy shifted: The electron energies are

$$E_n(r) = \omega' \left( n + \frac{1}{2} \right) + \beta \left( \frac{r}{2} - a \right)^2 \quad (\text{S7.209})$$

with  $\omega'^2 = 2\beta/m$ .

Part b) Now use  $E_n(r)$  as the effective potential for the nuclear motion. The Hamiltonian is

$$H = \frac{P^2}{2M_{\text{tot}}} + \frac{p^2}{2\mu} + E_n(r) \quad (\text{S7.210})$$

where  $P$  and  $p$  are the total and center of mass frame momentum,  $M_{\text{tot}}$  is the total mass  $2M$ , and  $\mu$  is the reduced mass  $M/2$ . The energy has a constant term  $\omega' \left( n + \frac{1}{2} \right)$ , plus

$$\beta \left( \frac{r}{2} - a \right)^2 = \frac{\mu\omega'^2}{2} (r - 4a)^2 \quad (\text{S7.211})$$

In the center of mass frame,  $P = 0$ , and the energy is the constant term plus the allowed energies of a harmonic oscillator with frequency  $\omega''$ :

$$E_{nm} = \omega' \left( n + \frac{1}{2} \right) + \omega'' \left( m + \frac{1}{2} \right) \quad (\text{S7.212})$$

where

$$\omega''^2 = \frac{\beta}{2\mu} = \frac{\beta}{M} \quad (\text{S7.213})$$

Part c) Now compare with the exact solution from problem 3.7. That is

$$\omega''^2 = \frac{\beta}{2\mu} = \frac{\beta}{M} \quad \omega'^2 = 2\frac{\beta}{m} \left( 1 + \frac{m}{2M} \right) \quad (\text{S7.214})$$

The error is of order  $m/M$ , as expected.