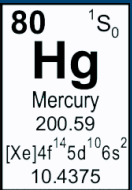




Precision Measurements using Atomic Systems



<http://www.gocomics.com/nonsequitur/1994/07/21>

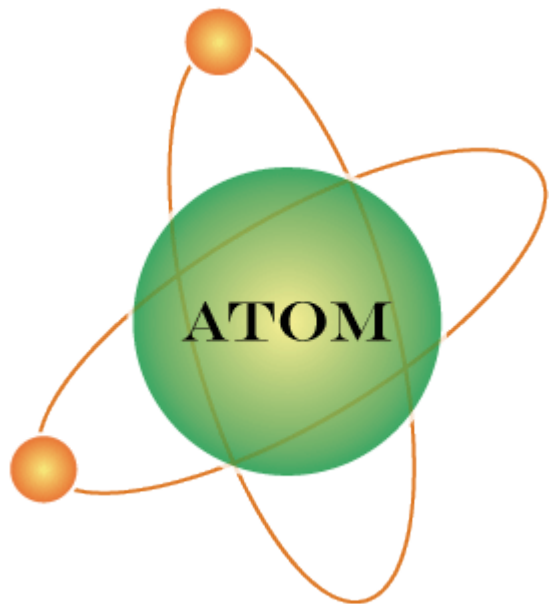
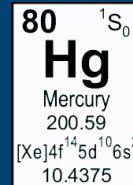
E.A.ALDEN

Leanhardt Lab

Department of Physics, University of Michigan



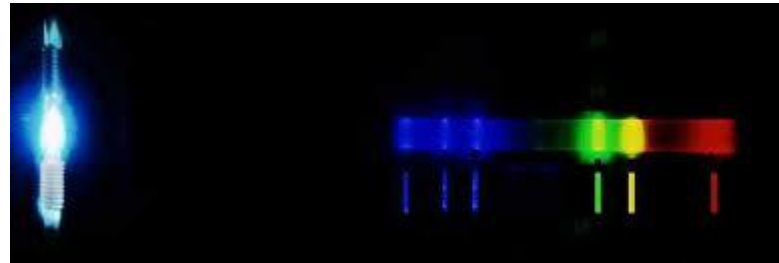
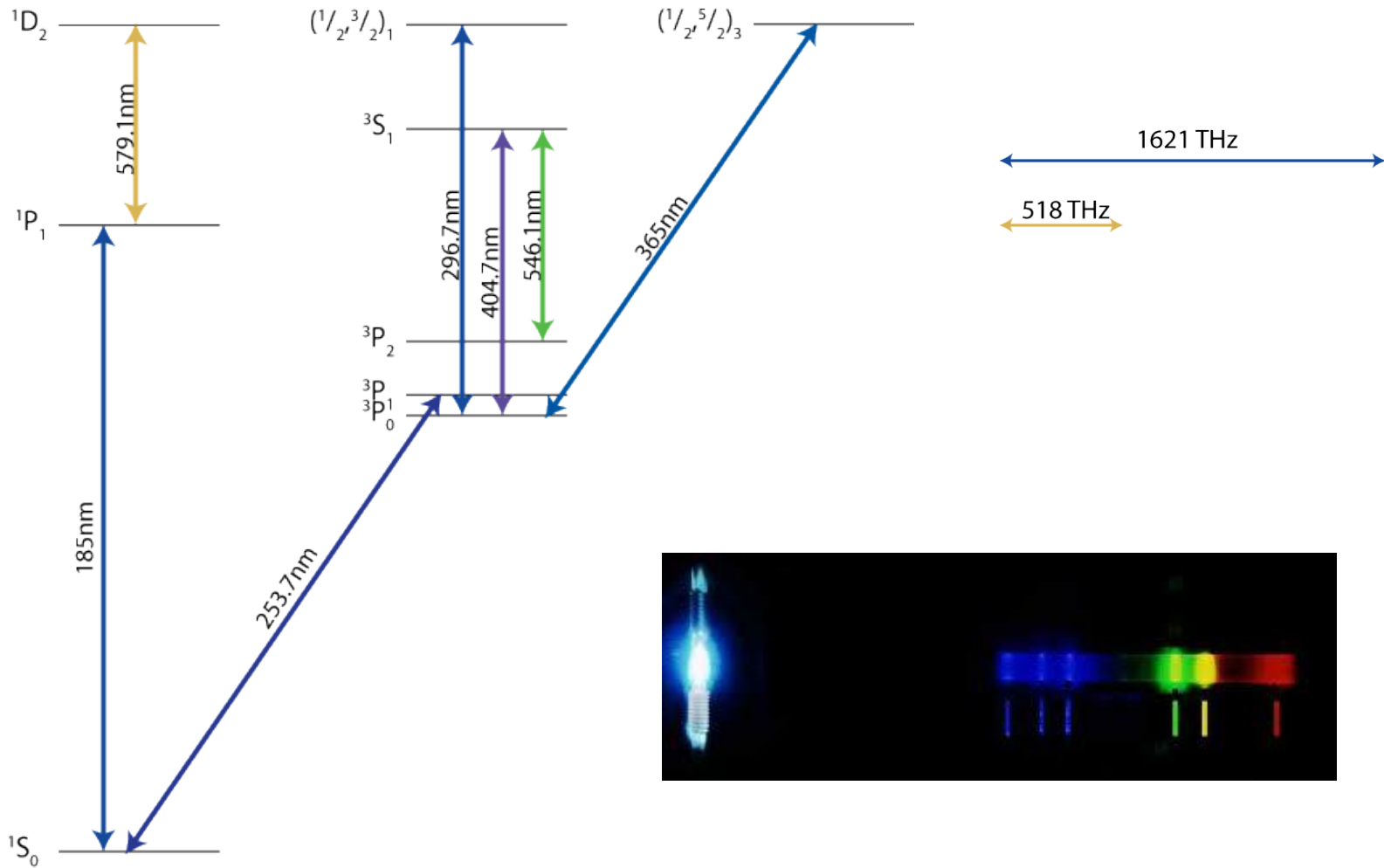
So what do you do?





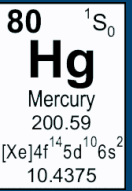
Light Interacts with Matter

80 1S_0
Hg
Mercury
200.59
[Xe]4f¹⁴5d¹⁰6s²
10.4375



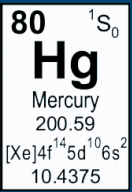


What is a clock?





Some Examples



Mechanical Clocks

- Pendulum Clock
- Quartz Clock

Atomic clocks

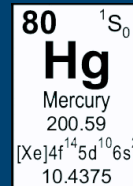
- Optical clocks



Sr Optical Atomic Clock Credit: Jun Ye Group



Why Atomic Clocks?

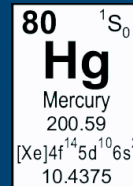


EM frequencies

Atoms are stable resonators



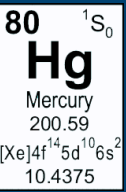
The Internet Told Me So...



Accessed 06-2011



The Internet also said...



CESIUM ATOMIC CLOCKS - Mozilla Firefox

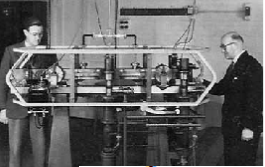
File Edit View History Bookmarks Tools Help

CESIUM ATOMIC CLOCKS

http://tycho.usno.navy.mil/cesium.html

Cesium Atoms at Work

"...till like a clock worn out with eating time."
 John Dryden (1631-1701)



The 1955 Cesium Atomic Clock at the National Physical Laboratory, UK. It kept time to a second in 300 years.

A "cesium(-beam) atomic clock" (or "cesium-beam frequency standard") is a device that uses as a reference the exact frequency of the microwave spectral line emitted by atoms of the metallic element cesium, in particular its isotope of atomic weight 133 (¹³³Cs-133). The integral of frequency is time, so this frequency, **9,192,631,770 hertz** (Hz = cycles/second), provides the fundamental unit of time, which may thus be measured by cesium clocks.

Today, cesium clocks measure frequency with an accuracy of from 2 to 3 parts in 10 to the 14th, i.e. 0.00000000000002 Hz; this corresponds to a time measurement accuracy of 2 nanoseconds per day or one second in 1,400,000 years. *It is the most accurate realization of a unit that mankind has yet achieved.* A cesium clock operates by exposing cesium atoms to microwaves until they vibrate at one of their resonant frequencies and then counting the corresponding cycles as a measure of time. The


It is the most accurate realization of a unit that mankind has yet achieved.

According to question refer to a change in the electron and nuclear spin ("hyperfine") energy level of the lowest set of orbits called the "ground state." Cesium is the best choice of atom for such a measurement because all of its 55 electrons but the outermost are confined to orbits in stable shells of electromagnetic force. Thus, the outermost electron is not disturbed much by the others. The cesium atoms are kept in a very good vacuum of about 10 trillionths of an atmosphere so that the cesium atoms are little affected by other particles. All this means that they radiate in a narrow spectral line whose wavelength or frequency can be accurately determined.

Kinds of Cesium Clocks

Cesium clocks are of two general kinds: a "laboratory (or primary) standard" about as large as a railroad flatcar and a "commercial (or secondary) standard" about as large as a suitcase. Only a few laboratory standards exist; they are used at research labs for frequency measurements of the highest accuracy. Examples are the NIST-7 standard at the [National Institute of Standards and Technology \(NIST\)](#) in Boulder, CO and the atomic fountains at [NIST](#), [Physikalisch-Technische Bundesanstalt \(PTB\)](#), Germany, the [Paris Observatory](#), France, and [USNO](#). Commercial standards, being industrially produced, are cheaper, but still provide state-of-the-art measurement of precise time and time interval. A timing center maintaining an ensemble of such clocks can average their readings to produce a "mean timescale" for scientific and public use.

The U.S. Naval Observatory operates about 70 such cesium clocks, as well as other precision clocks like hydrogen masers, in 18 vaults whose temperature and, usually, humidity are closely controlled in order to minimize perturbations by their environment. The time measurements are made by devices called **time-interval counters** that compare each clock's time against that of one "Master Clock," whose frequency is steered to match its time to the average of the other clocks. This time is the Observatory's measure of the atomic time called **Coordinated Universal Time (UTC)**. Some cesium clocks are transported to remote locations in order to synchronize other clocks.



USNO Cesium Clocks

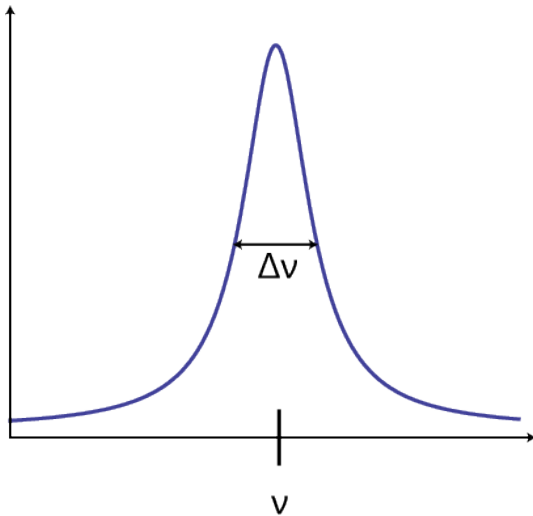
Most of the Observatory's cesium clocks are model HP5071A, made by Agilent



How good is a clock?

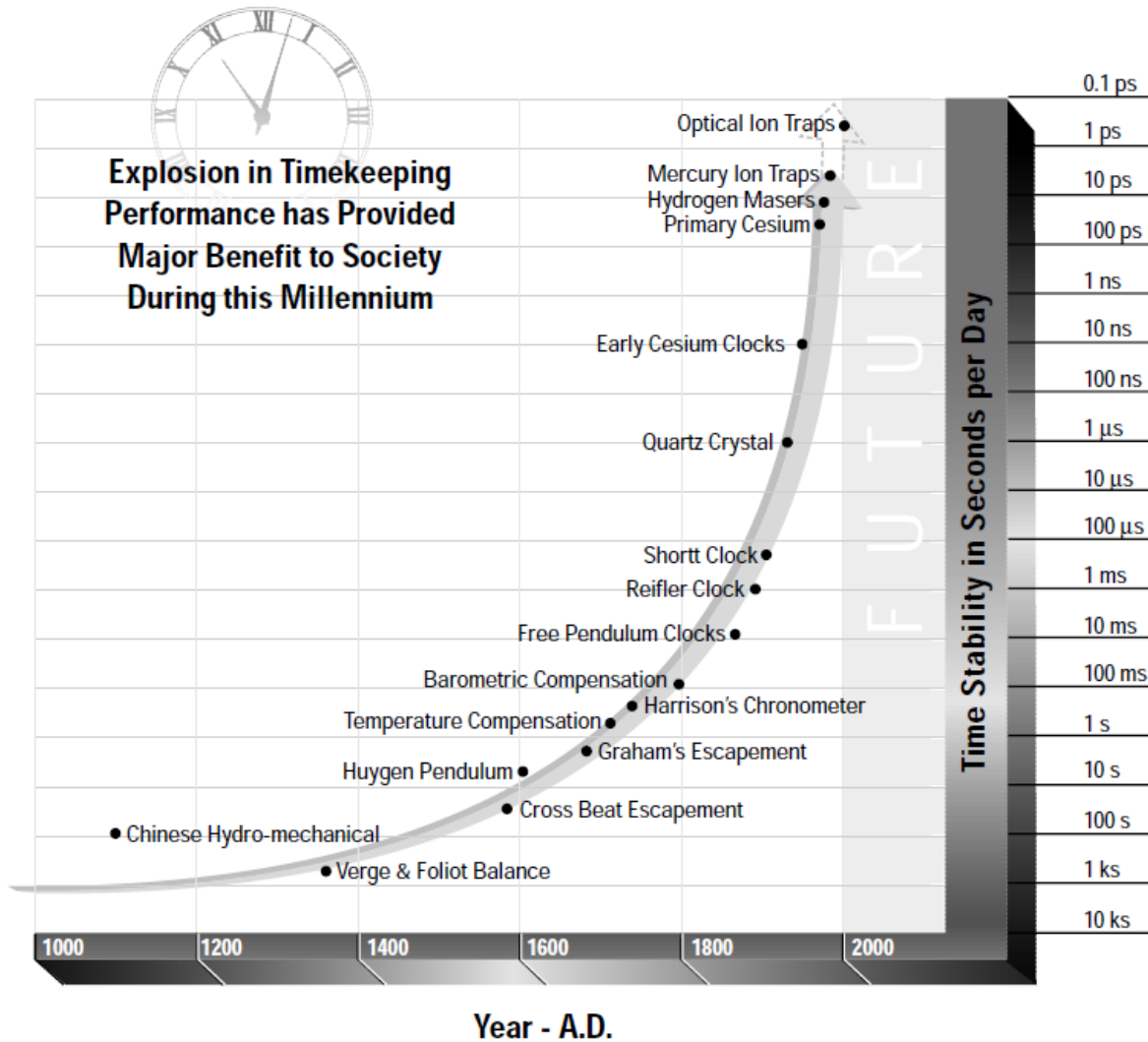
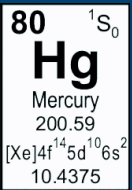
80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	

$$\sigma(\tau) = \left\langle \frac{\Delta \nu}{\nu} \right\rangle_{\tau}$$





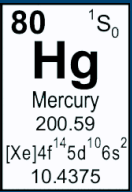
History of Clocks



D.W. Allan, N. Ashby, and C.C. Hodge, *The Science of timekeeping*, Hewlett-Packard Application Note 1289

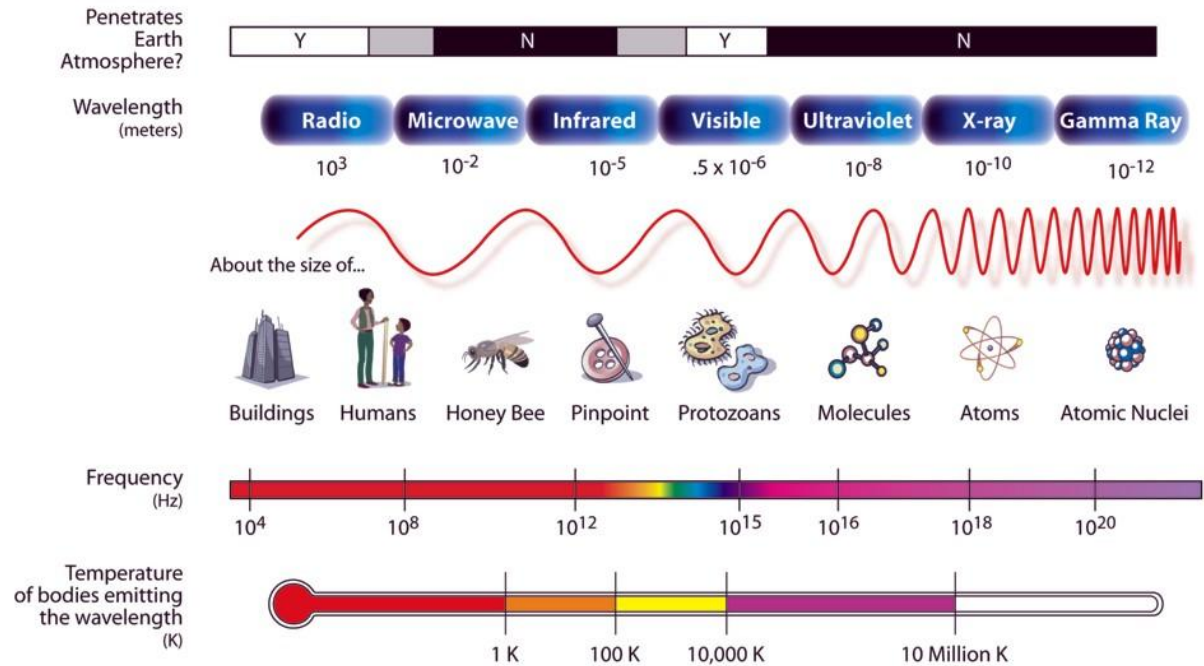


Make ν big



$$\sigma(\tau) = \left\langle \frac{\Delta \nu}{\nu} \right\rangle_{\tau}$$

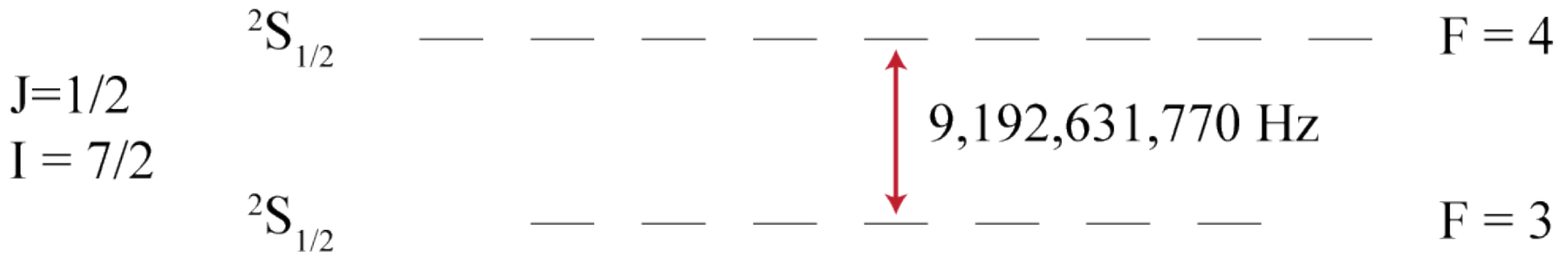
THE ELECTROMAGNETIC SPECTRUM





Where is ^{133}Cs ?

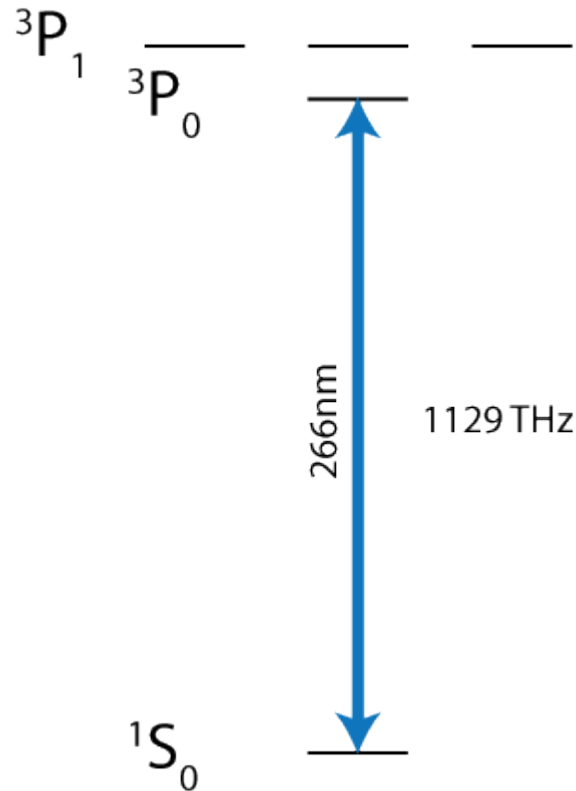
80 $^1\text{S}_0$
Hg
Mercury
200.59
[Xe]4f¹⁴5d¹⁰6s²
10.4375





Neutral Hg

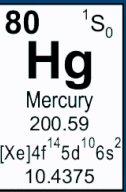
80 1S_0
Hg
Mercury
200.59
[Xe]4f¹⁴5d¹⁰6s²
10.4375



Isotope	Spin
^{196}Hg	0
^{198}Hg	0
^{199}Hg	1/2
^{200}Hg	0
^{201}Hg	3/2
^{202}Hg	0
^{204}Hg	0



Make $\Delta\nu$ small

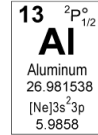


$$\sigma(\tau) = \left\langle \frac{\Delta\nu}{\nu} \right\rangle_{\tau}$$

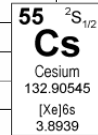
- Narrow linewidth transition
- Narrow linewidth laser
- Large number of addressed resonators
- Long coherence times
- Long measurement times
- Laser excitation geometry (transit broadening)
- So on....

TABLE I. Systematic effects that shift the clock from its ideal unperturbed frequency. Shifts and uncertainties given are in fractional frequency units ($\Delta\nu/\nu$). See text for discussion.

Effect	Shift (10^{-18})	Uncertainty (10^{-18})
Excess micromotion	-9	6
Secular motion	-16.3	5
Blackbody radiation shift	-9	3
Cooling laser Stark shift	-3.6	1.5
Quad. Zeeman shift	-1079.9	0.7
Linear Doppler shift	0	0.3
Clock laser Stark shift	0	0.2
Background-gas collisions	0	0.5
AOM freq. error	0	0.2
Total	-1117.8	8.6



Physical Effect	Bias	Type B Uncertainty
Gravitational Red shift	+179.95	0.03
Second-Order Zeeman	+180.91	0.025
Blackbody	-22.84	0.28
Microwave Amplitude Shift	-0.05	0.15
Spin Exchange (low density)	(-0.32)*	(0.17)*
AC Zeeman (heaters)	0.05	0.05
Cavity Pulling	0.02	0.02
Rabi Pulling	10^{-4}	10^{-4}
Ramsey Pulling	10^{-4}	10^{-4}
Majorana Transitions	0.02	0.02
Fluorescence Light Shift	10^{-5}	10^{-5}
Cavity Phase (distributed)	0.02	0.02
Second-Order Doppler	0.02	0.02
DC Stark Effect	0.02	0.02
Background Gas Collisions	10^{-3}	10^{-3}
Bloch-Siegert	10^{-4}	10^{-4}
RF Spectral purity	3×10^{-3}	3×10^{-3}
Integrator offset	0	0.01
Total Type B Standard Uncertainty		0.33



*For information purposes only. Not used in total. see section 1-B for details

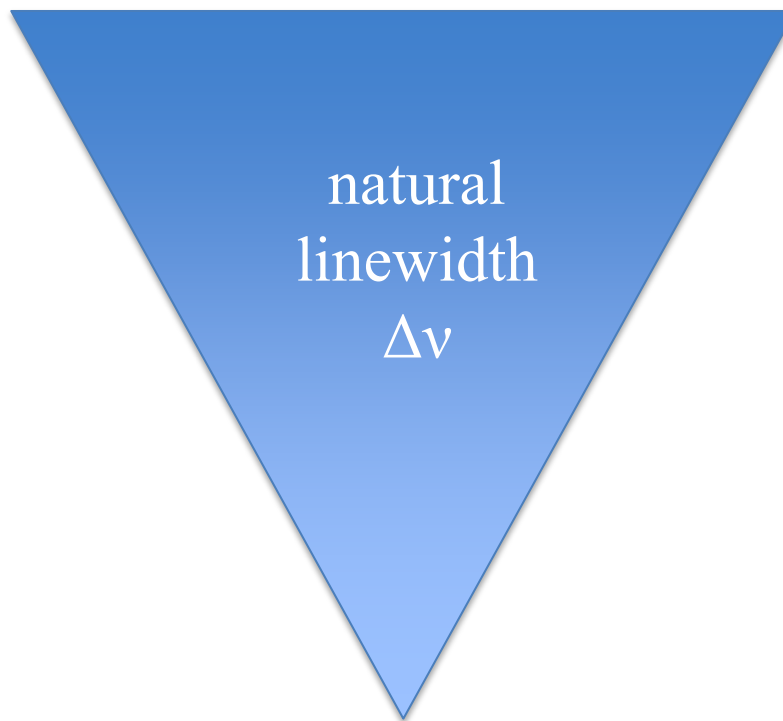
Table 1 – The list of known frequency biases for NIST-F1. This includes both the magnitude of the biases as well as the uncertainty of each individual contribution to the final uncertainty.



Choose Transition Wisely

80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	

“forbiddenness”





Optical Clock Level Structure

80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	

3P_0 _____

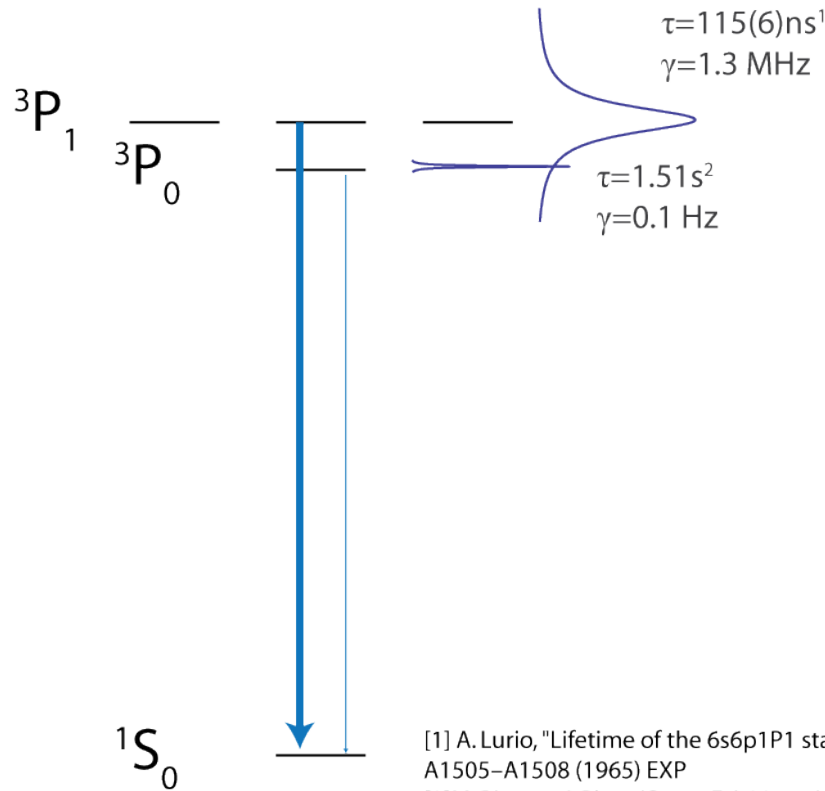
1S_0 _____



Narrow Linewidth Transition

80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	

¹⁹⁹Hg I=1/2

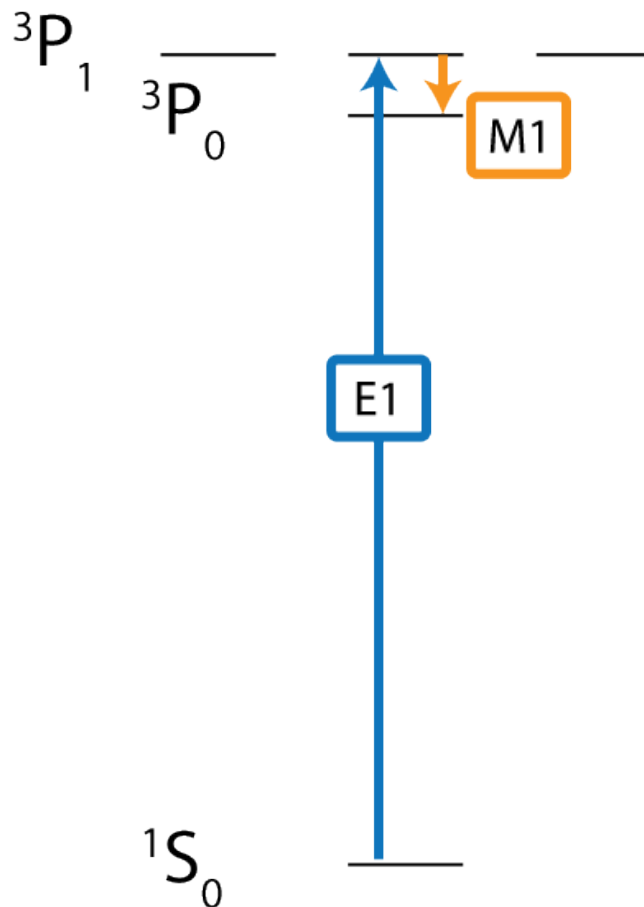


- [1] A. Lurio, "Lifetime of the 6s6p1P1 state of mercury," Phys. Rev. 140, A1505–A1508 (1965) EXP
- [2] M. Bignon, J. Phys. (Orsay, Fr.) 28, 51 (1967). THEORY



Even Narrower Transition?

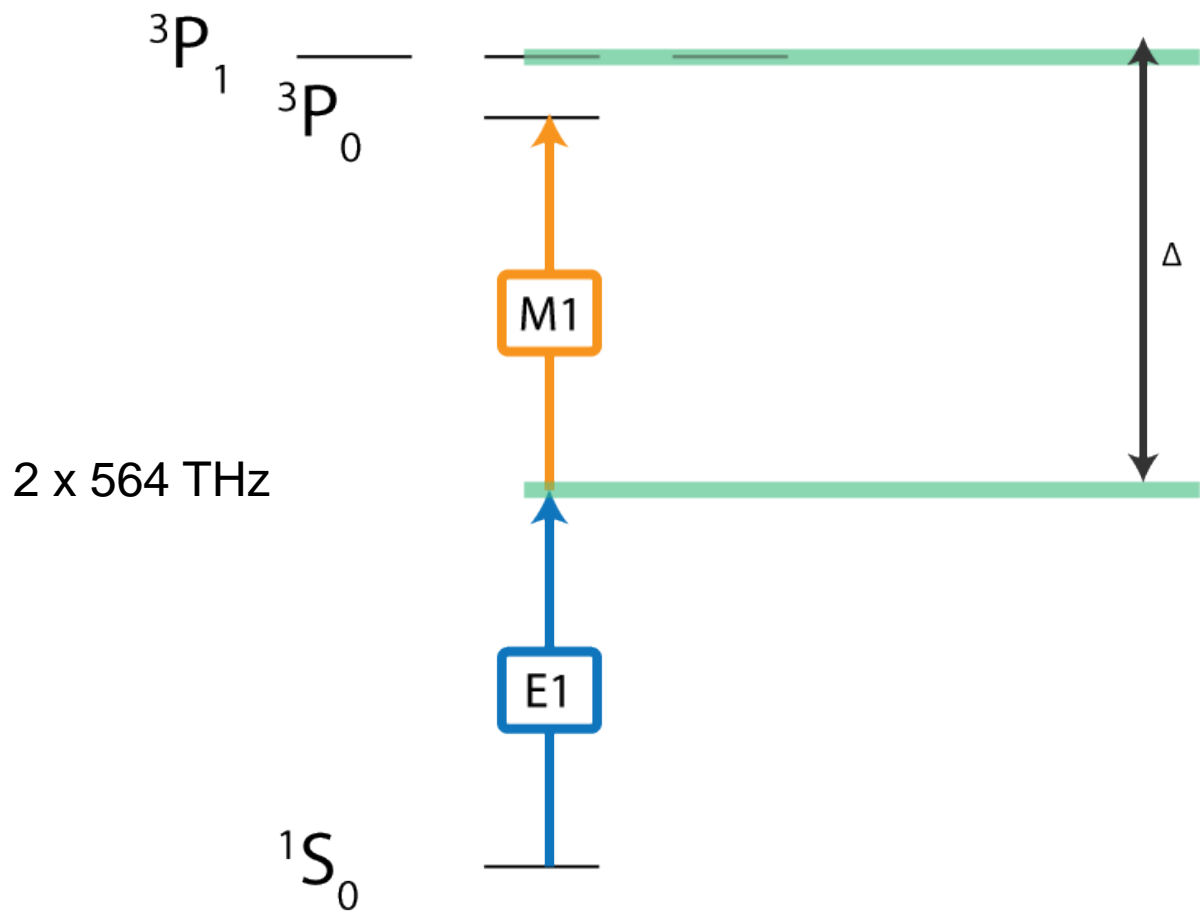
80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	





Even Narrower Transition?

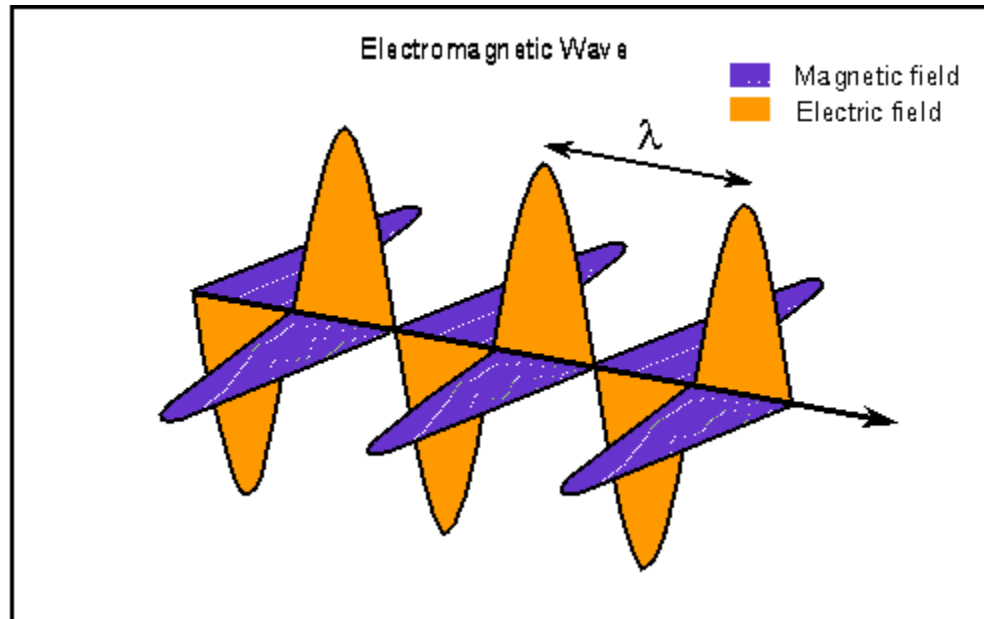
80 1S_0
Hg
Mercury
200.59
[Xe]4f¹⁴5d¹⁰6s²
10.4375





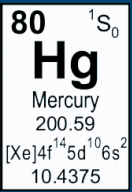
EM waves

80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	





Advantages

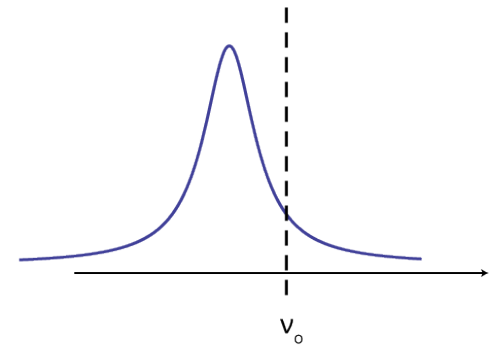
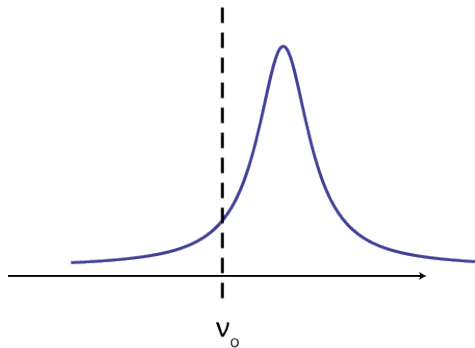
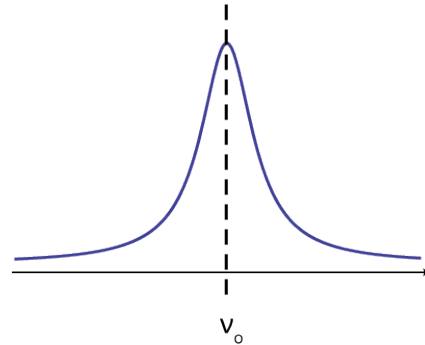


- Linewidth preserving access to 3P_0 level
- Doppler-Free Spectroscopy with room temperature atoms.
- Vapor cells have high densities of atoms



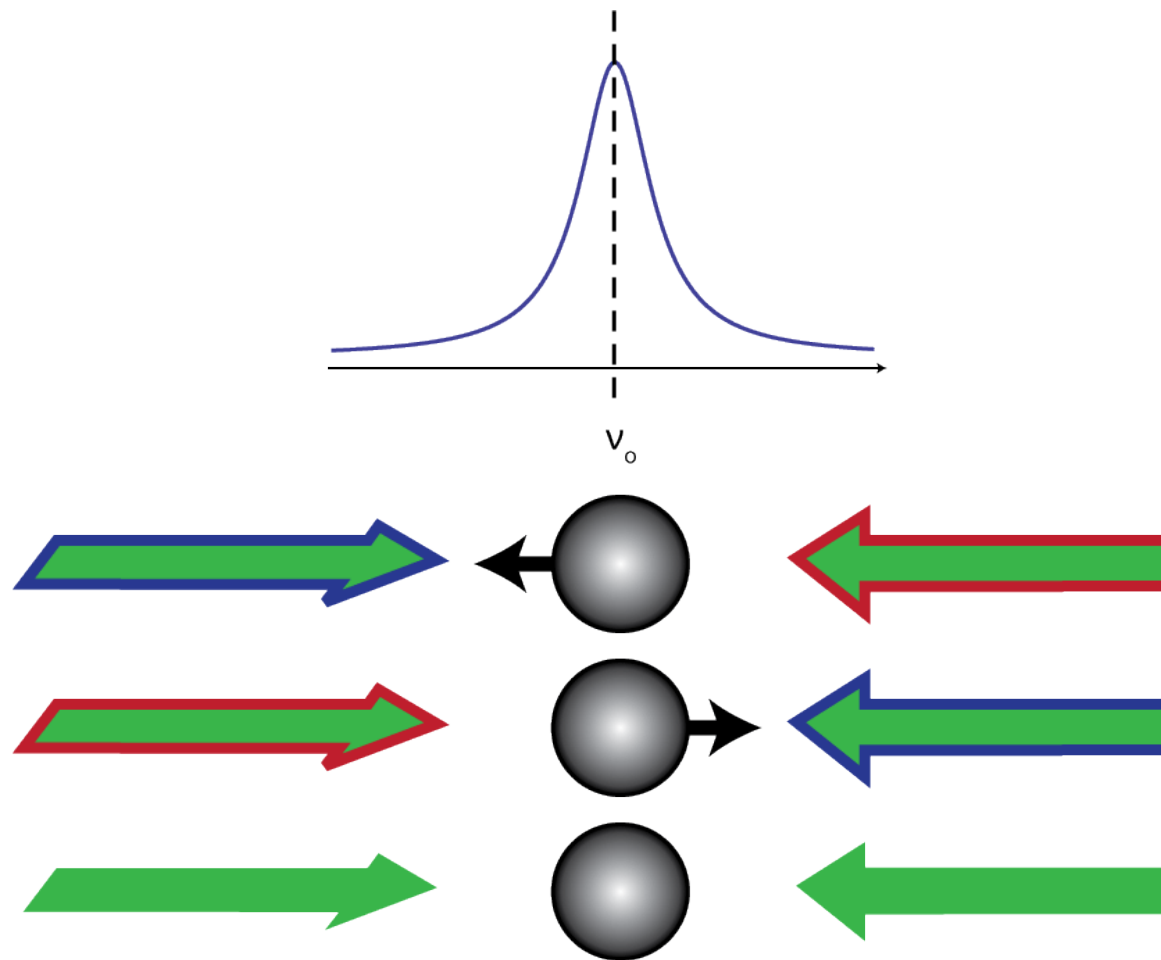
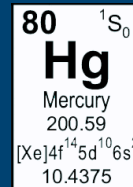
Doppler Broadening

80 ¹S₀
Hg
Mercury
200.59
[Xe]4f¹⁴5d¹⁰6s²
10.4375





Doppler-Free Spectroscopy





Drawbacks: Transition Rates

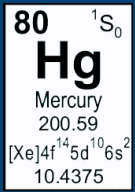
80	¹ S ₀
Hg	
Mercury	
200.59	
[Xe]4f ¹⁴ 5d ¹⁰ 6s ²	
10.4375	

$$R_{1\gamma} \propto \left\langle {}^3P_1 | \mu | {}^1S_0 \right\rangle_{E1}^2 I$$

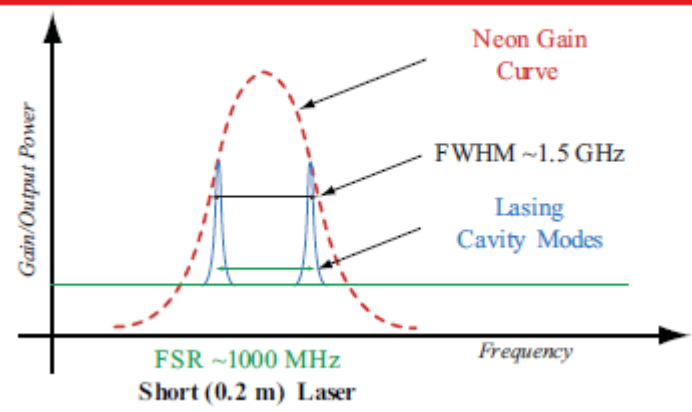
$$R_{2\gamma} \propto \frac{\left\langle {}^3P_0 | \mu | {}^3P_1 \right\rangle_{M1}^2 \left\langle {}^3P_1 | \mu | {}^1S_0 \right\rangle_{E1}^2 I^2}{\Delta^2}$$



Off-the-shelf Laser

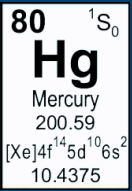


HNL008R(L) Red HeNe Laser System





Narrow Linewidth Laser



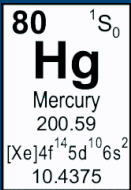
- <10 kHz linewidth 1062nm @ ~10mW
- Fiber amplifier 10 mW -> 50W
- Single pass PPMgO:SLT SHG



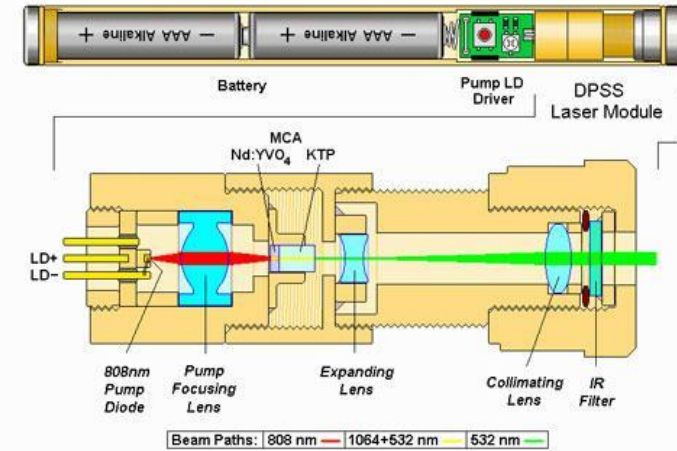
10-25W @ 531nm



Green Laser Pointers



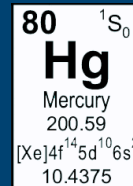
- 5mW w/ 1.1mm dia beam
->0.58 W/cm²
- Human eye blink time
– 200ms
- Damage threshold
-> 0.005 W/cm²
- We're building a laser ->77 W/cm²



Wikipedia 2011



And the point of all this...



when ↔ where



Recalculating



How do we fit into this puzzle?



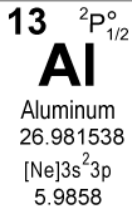
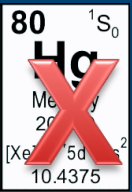
- Explore this level structure transition
- Make clocks portable
- Map the geoid w/ 30m gravitational redshift resolution

$$\frac{\Delta \nu}{\nu} \approx 10^{-16}$$

ESTIMATE



The Best in the World



PRL 104, 070802 (2010)

PHYSICAL REVIEW LETTERS

week ending
19 FEBRUARY 2010

Frequency Comparison of Two High-Accuracy Al^+ Optical Clocks

C. W. Chou,^{*} D. B. Hume, J. C. J. Koelemeij,[†] D. J. Wineland, and T. Rosenband

Time and Frequency Division, National Institute of Standards and Technology, Boulder, Colorado 80305, USA

(Received 23 November 2009; published 17 February 2010)

We have constructed an optical clock with a fractional frequency inaccuracy of 8.6×10^{-18} , based on

$$\frac{\Delta \nu}{\nu} = 8.6 \times 10^{-18}$$



They can see 33cm

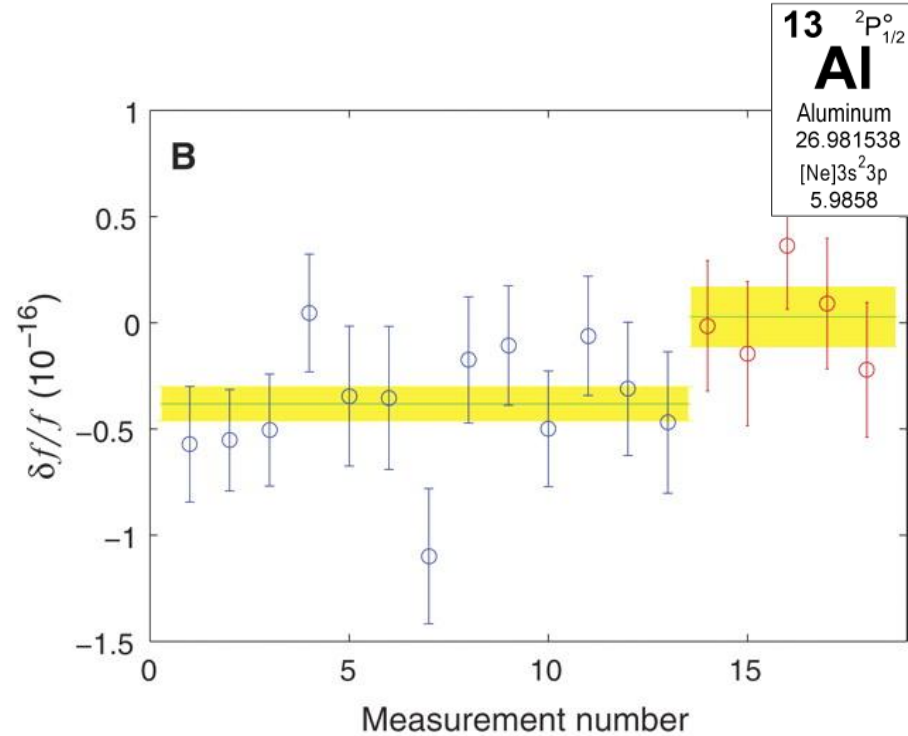
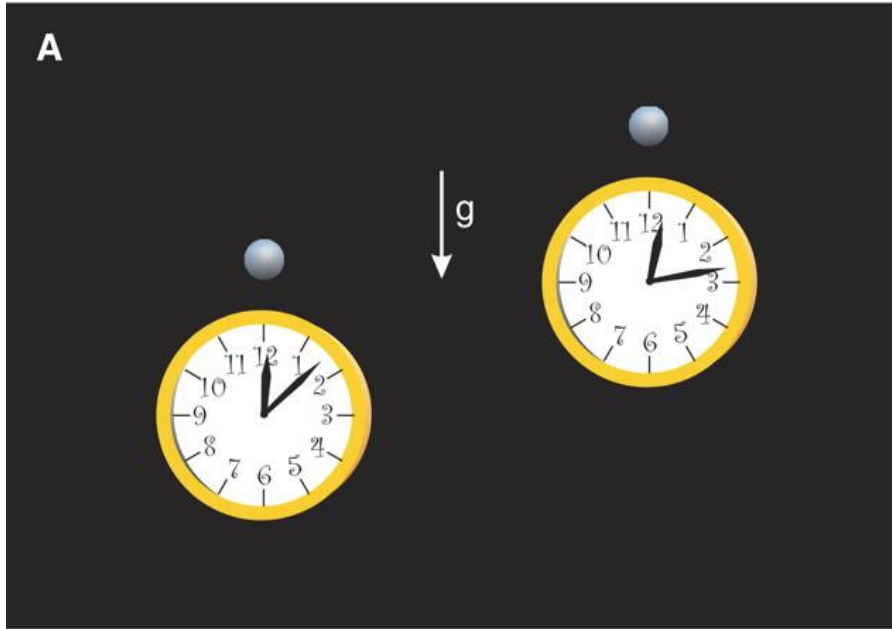
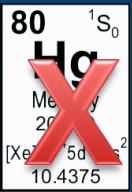


Fig. 3 Gravitational time dilation at the scale of daily life.

C W Chou et al. Science 2010;329:1630-1633

Published by AAAS





80 ¹S₀
Hg
Mercury
200.59
[Xe]4f¹⁴5d¹⁰6s²
10.4375

Leanhardt Research Group

Aaron Leanhardt
(PI)

Jinhai Chen
(post-doc)

Yisa Rumala
(grad)

Jeongwon Lee
(grad)

Emily Alden (+1)
(grad)

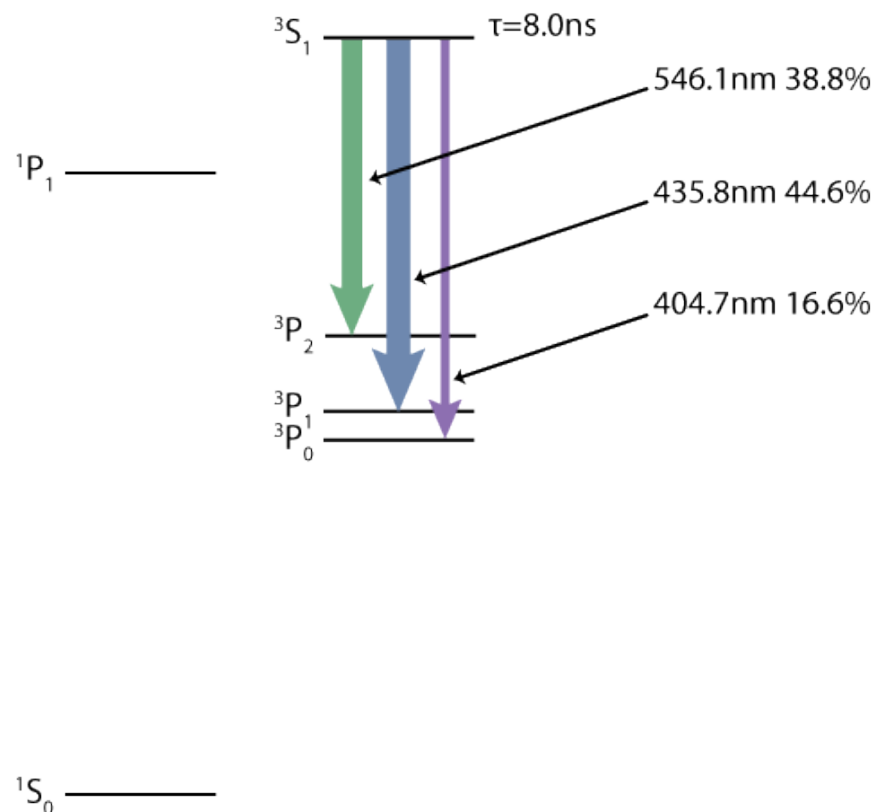
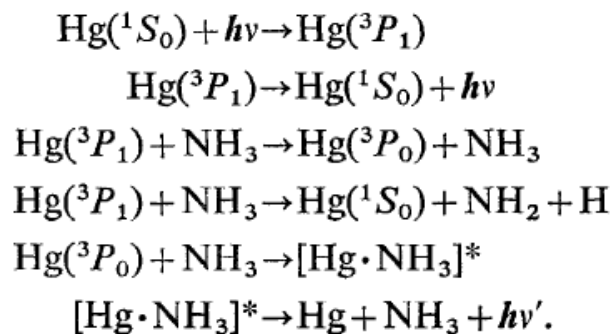
Kaitlin Moore
(post-bac)

Not Shown:

- Charlie Steiner (undergrad)

Hg Level Structure & ³S₁ Branching Ratios

Collision NH₃



Mercury-Sensitized Luminescence of NH₃ and ND₃

BY R. H. NEWMAN, C. G. FREEMAN, M. J. McEWAN, R. F. C. CLARIDGE
 AND L. F. PHILLIPS

Chemistry Dept., University of Canterbury, Christchurch, New Zealand

Received 11th May, 1970

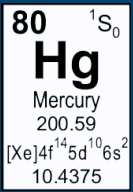


Mike Mosedale

- Power broadening
- Stark shifts
- Rayleigh scattering

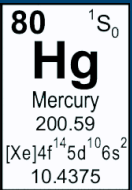


Light Interacts with Matter





Atomic Clock's Figure of Merit

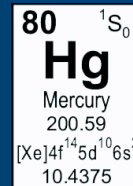


Allan Deviation

$$\frac{\Delta \nu}{\nu} \quad \sigma_y(\tau) \approx \frac{\Delta \nu}{\nu} \frac{1}{\sqrt{N\tau}}$$



Clocks



Technique	$\sigma_{\langle\mu\rangle}$ @ 24 hours	Date
Pendulum	10 s	1656
Chronometer	0.4 s	1761
Quartz	10 ⁻⁴ s	1927
Cs	4.3 x 10 ⁻¹¹ s	1952
Rb	10 ⁻⁶ s	1958
Al ⁺	7.4 x 10 ⁻¹³ s	2010

