## 1. Problem 10.10a

10 Points

Winter 2004

We consider the Smythe-Kirchhoff integral,

$$E_{diff} = \frac{1}{2\pi} \nabla \times \int_{hole} \hat{\mathbf{n}} \times \mathbf{E} \frac{\exp(\mathbf{i}kR)}{R} da'$$

where  $\hat{\mathbf{n}}$  is the normal of the conducting plane pointing into the volume of interest, and  $\mathbf{R} = \mathbf{x} - \mathbf{x}'$ . It is  $\hat{\mathbf{n}} \times \mathbf{E} = \hat{\mathbf{n}} \times \mathbf{E}_{tan}$ , where  $E_{tan}$  is the total electric field tangential with the conducting plane. Also, in the radiation zone

$$\frac{\exp(\mathrm{i}kR)}{R} = \frac{\exp(\mathrm{i}kr)}{r}\exp(-\mathrm{i}\mathbf{k}\mathbf{x}')$$

where  $\mathbf{k} = k\hat{\mathbf{r}}$  is the k-vector pointing to the observation point.

Thus, in the radiation zone

$$E_{diff} = \frac{1}{2\pi} \nabla \times \frac{\exp(ikr)}{r} \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \exp(-i\mathbf{kx}') da'$$

Also, since for expressions of the kind " $E = \nabla \times \frac{\exp(ikr)}{r} \mathbf{F}(\theta, \phi)$ " in the radiation zone the usual replacement " $\nabla \times = i\mathbf{k} \times$ " applies, we obtain Eq. 10.109:

$$E_{diff} = \frac{\mathrm{i}}{2\pi} \mathbf{k} \times \frac{\exp(\mathrm{i}kr)}{r} \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \exp(-\mathrm{i}\mathbf{k}\mathbf{x}') da'$$

Since the aperture is small, we can make the small-source approximation for the fields emanating from the hole,  $\exp(-i\mathbf{kx'}) = 1 - i\mathbf{k} \cdot \mathbf{x'}$ , and get:

$$E_{diff} = \frac{\mathrm{i}}{2\pi} \frac{\exp(\mathrm{i}kr)}{r} \mathbf{k} \times \left[ \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) da' - \mathrm{i} \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \mathbf{k} \cdot \mathbf{x}' da' \right]$$

As advertised in class, we employ the vector identity Eq. 9.31 with  $\hat{\mathbf{n}} \times \mathbf{E}_{tan}$  in place of  $\mathbf{J}$ ,

$$(\hat{\mathbf{n}} \times \mathbf{E}_{tan})\mathbf{k} \cdot \mathbf{x}' = \frac{1}{2} \left[ \mathbf{x}' \times (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \right] \times \mathbf{k} + \frac{1}{2} \left[ (\mathbf{k} \cdot \mathbf{x}')(\hat{\mathbf{n}} \times \mathbf{E}_{tan}) + (\mathbf{k} \cdot (\hat{\mathbf{n}} \times \mathbf{E}_{tan})) \mathbf{x}' \right]$$

to get

$$E_{diff} = \frac{i}{2\pi} \frac{\exp(ikr)}{r} \mathbf{k} \times \left[ \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \left\{ 1 - \frac{1}{2} \mathbf{i} \mathbf{k} \cdot \mathbf{x}' \right\} da' - \frac{i}{2} \int_{hole} \mathbf{x}' \times (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) da' \times \mathbf{k} - \frac{i}{2} \int_{hole} \mathbf{k} \cdot (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \mathbf{x}' da' \right]$$

Since the hole is small, in the first integral we may set  $1 - \frac{1}{2}\mathbf{i}\mathbf{k}\cdot\mathbf{x}' = 1$ . In the second integral,  $\mathbf{x}' \times (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) = (\mathbf{x}' \cdot \mathbf{E}_{tan})\hat{\mathbf{n}} - (\mathbf{x}' \cdot \hat{\mathbf{n}})\mathbf{E}_{tan} = (\mathbf{x}' \cdot \mathbf{E}_{tan})\hat{\mathbf{n}}$ , because  $\mathbf{x}' \cdot \hat{\mathbf{n}} = 0$ . Thus,

$$E_{diff} = \frac{\mathrm{i}}{2\pi} \frac{\exp(\mathrm{i}kr)}{r} \mathbf{k} \times \left[ \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) da' - \frac{\mathrm{i}}{2} \left( \hat{\mathbf{n}} \int_{hole} \mathbf{x}' \cdot \mathbf{E}_{tan} da' \right) \times \mathbf{k} - \frac{\mathrm{i}}{2} \int_{hole} \mathbf{k} \cdot (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \mathbf{x}' da' \right]$$

The first term can be re-written as

$$E_{diff,1} = -\frac{Z_0 k^2}{4\pi} \frac{\exp(ikr)}{r} \left( \hat{\mathbf{k}} \times \left[ \frac{-2i}{Z_0 k} \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) da' \right] \right)$$

This field can be compared with Eq. 9.36. Thereby, the term in the square bracket can be identified with an effective magnetic dipole

$$\mathbf{m} = \frac{-2\mathrm{i}}{Z_0 k} \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) da' = \frac{2}{\mathrm{i}\omega\mu} \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) da'$$

(note in vacuum  $k = \omega/c = \omega \mu/Z_0$ ). The second term,

$$E_{diff,2} = \frac{Z_0 ck^2}{4\pi} \frac{\exp(ikr)}{r} \hat{\mathbf{k}} \times \left( \left[ \frac{1}{Z_0 c} \hat{\mathbf{n}} \int_{hole} \mathbf{x}' \cdot \mathbf{E}_{tan} da' \right] \times \mathbf{k} \right)$$

can be, by comparison with Eq. 9.19, identified with the electric field of an electric dipole

$$\mathbf{p} = \epsilon \hat{\mathbf{n}} \int_{hole} \mathbf{x}' \cdot \mathbf{E}_{tan} da'$$

Both the first and the third term have an  $\hat{\mathbf{n}} \times \mathbf{E}_{tan}$  under the integral. Further, the third term is of the order of the first term times  $kx' \ll 1$ . Thus, the third term can be neglected.

## 2. Problem 10.12

We start with the Smythe-Kirchhoff formula in the radiation zone,

$$E_{diff} = \frac{\mathrm{i}}{2\pi} \frac{\exp(\mathrm{i}kr)}{r} \mathbf{k} \times \int_{hole} (\hat{\mathbf{n}} \times \mathbf{E}_{tan}) \exp(-\mathrm{i}\mathbf{k}\mathbf{x}') da'$$

The plane normal  $\hat{\mathbf{n}} = \hat{\mathbf{z}}$ , the incident wavevector  $\mathbf{k}_0 = k(\cos \alpha \hat{\mathbf{z}} + \sin \alpha \hat{\mathbf{x}})$ , and the wavevector pointing to the observation point,  $\mathbf{k} = k(\sin \theta \cos \phi \hat{\mathbf{x}} + \sin \theta \sin \phi \hat{\mathbf{y}} + \cos \theta \hat{\mathbf{z}})$ . The incident electric field is linearly polarized transverse to the plane of incidence (the *xz*-plane), i.e.  $\mathbf{E}_0 = E_0 \hat{\mathbf{y}}$ . The circular hole over which we integrate extends in the x'y'-plane. Thus, using 2-dimensional cylindrical coordinates  $\rho'$  and  $\beta'$  in the x'y'-plane,

$$E_{diff} = \frac{iE_0}{2\pi} \frac{\exp(ikr)}{r} \mathbf{k} \times \int_{\rho'=0}^a \int_{\beta=0}^{2\pi} (\hat{\mathbf{z}} \times \hat{\mathbf{y}}) \exp(i(\mathbf{k_0} - \mathbf{k})\mathbf{x}')\rho' d\rho' d\beta'$$
$$= \frac{-iE_0}{2\pi} \frac{\exp(ikr)}{r} (\mathbf{k} \times \hat{\mathbf{x}}) \int_{\rho'=0}^a \left\{ \int_{\beta'=0}^{2\pi} \exp(ik\rho'(\sin\alpha\cos\beta' - \sin\theta\cos(\phi - \beta'))) d\beta' \right\} \rho' d\rho'$$

The angular function in the exponent can be rewritten,

$$\sin \alpha \cos \beta' - \sin \theta \cos(\phi - \beta') = \cos \beta \left[\sin \alpha - \sin \theta \cos \phi\right] + \sin \beta \left[-\sin \theta \sin \phi\right]$$
$$= \xi \cos(\beta' + \delta)$$

where the amplitude  $\xi$  is the square-root of the sum of the squares of the terms in square-brackets, and  $\delta$  is a constant phase shift. Thus,

$$\xi = \sqrt{\left[\sin\alpha - \sin\theta\cos\phi\right]^2 + \left[-\sin\theta\sin\phi\right]^2} = \sqrt{\sin^2\theta + \sin^2\alpha - 2\sin\alpha\sin\theta\cos\phi}$$

In the angular integral the phase shift  $\delta$  is irrelevant, because the angular integral is over a full circle:

$$\int_{\beta'=0}^{2\pi} \exp(ik\rho'(\sin\alpha\cos\beta' - \sin\theta\cos(\phi - \beta')))d\beta' = \int_{0}^{2\pi} \exp(ik\rho'\xi\cos(\beta' + \delta))d\beta'$$
$$= \int_{0}^{2\pi} \exp(ik\rho'\xi\cos\beta')d\beta' = \int_{0}^{2\pi} \exp(ik\rho'\xi\sin\beta')d\beta' = 2\pi J_0(k\rho'\xi)$$

and the diffracted fields

$$\begin{aligned} \mathbf{E}_{diff}(r,\alpha,\theta,\phi) &= -\mathrm{i}E_0 \frac{\exp(\mathrm{i}kr)}{r} (\mathbf{k} \times \hat{\mathbf{x}}) \int_{\rho'=0}^{a} J_0(k\rho'\xi) \rho' d\rho' \\ &= -\mathrm{i}E_0 a^2 \frac{\exp(\mathrm{i}kr)}{r} (\mathbf{k} \times \hat{\mathbf{x}}) \frac{J_1(k\xi a)}{ak\xi} \\ \mathbf{H}_{diff}(r,\alpha,\theta,\phi) &= \frac{1}{Z_0} \hat{\mathbf{k}} \times \mathbf{E}_{diff}(\mathbf{x}) \end{aligned}$$

The diffracted power per solid angle

$$\begin{aligned} \frac{dP}{d\Omega} &= r^2 \frac{1}{2Z_0} \mathbf{E}_{diff} \cdot \mathbf{E}^*_{diff} \\ &= \frac{|E_0|^2}{2Z_0} a^4 \left( \frac{J_1(k\xi a)}{ak\xi} \right)^2 |(\mathbf{k} \times \hat{\mathbf{x}})|^2 \\ &= \frac{|E_0|^2}{2Z_0} a^4 k^2 \left( \frac{J_1(k\xi a)}{ak\xi} \right)^2 (\cos^2 \theta + \sin^2 \theta \sin^2 \phi) \end{aligned}$$

This can be normalized with the power incident on the hole,

$$P_{in} = \frac{1}{2Z_0} \left| E_0 \right|^2 a^2 \pi \cos \alpha$$

yielding

$$\frac{dP}{d\Omega}/P_{in} = \frac{a^2k^2}{\pi\cos\alpha} \left(\frac{J_1(k\xi a)}{ak\xi}\right)^2 (\cos^2\theta + \sin^2\theta\sin^2\phi)$$

**b**): The result we have obtained equals that of Eq. 10.114 (case of polarization in plane of incidence) times a factor

$$\frac{1}{\cos^2 \alpha} \frac{(\cos^2 \theta + \sin^2 \theta \sin^2 \phi)}{(\cos^2 \theta + \sin^2 \theta \cos^2 \phi)}$$

It is also somewhat similar with the result of the scalar calculation, given in Eq. 10.119. In fact, all three results share the essential dependence

$$\propto k^2 a^2 \left( \frac{J_1(k\xi a)}{ak\xi} \right)^2$$

It is also noted that for the case of normal incidence  $\alpha = 0$  the two vectorial results are identical, as required. To see this, take the polarization directions into account. Then, note that in the case of normal incidence in both calculations - polarization perpendicular to and in the plane of incidence - the respective terms  $\sin \phi$ and  $\cos \phi$  are equal to the sine of the angle between the laser polarization and the projection of **k** into the *xy*-plane.

## 3. Problem 10.16

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= \sum_{i} \frac{(\epsilon_{i}^{*} \cdot F_{sh})(\epsilon_{i} \cdot F_{sh}^{*})}{E_{0}E_{0}^{*}} \\ &= \frac{k^{2}}{4\pi^{2}} \sum_{i} |\epsilon_{i}^{*} \cdot \epsilon_{0}|^{2} \left( \int_{shadow} \exp(-i\mathbf{k}_{\perp} \cdot \mathbf{x}_{\perp}) d^{2}x_{\perp} \right) \left( \int_{shadow} \exp(i\mathbf{k}_{\perp} \cdot \mathbf{x}_{\perp}') d^{2}x_{\perp}' \right) \end{aligned}$$

where the integrals go over the shadow of the object in the xy-plane. As orthonormal basis for the exit polarizations we can use

$$\epsilon_1 = \hat{\phi}_k = \begin{pmatrix} -\sin\phi_k \\ \cos\phi_k \\ 0 \end{pmatrix} \quad \text{and} \quad \epsilon_2 = \hat{\theta}_k = \begin{pmatrix} \cos\theta_k\cos\phi_k \\ \cos\theta_k\sin\phi_k \\ -\sin\theta_k \end{pmatrix}$$

To cover the case of arbitrary incident polarization, we use  $\epsilon_0 = c_1 \hat{\mathbf{x}} + c_2 \hat{\mathbf{y}}$  with complex numbers  $c_1 c_1^* + c_2 c_2^* = 1$ . Then,

$$\sum_{i} |\epsilon_{i}^{*} \cdot \epsilon_{0}|^{2} = |c_{1}|^{2} (\sin^{2} \phi_{k} + \cos^{2} \theta_{k} \cos^{2} \phi_{k}) + |c_{2}|^{2} (\cos^{2} \phi_{k} + \cos^{2} \theta_{k} \sin^{2} \phi_{k}) =: A(\theta_{k}, \phi_{k})$$

Then,

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= \frac{k^2}{4\pi^2} \int_{sh} \int_{sh} \exp(-\mathbf{i}\mathbf{k}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{x}'_{\perp})) A(\theta_k, \phi_k) d^2 x_{\perp} d^2 x'_{\perp} \\ \sigma &= \frac{k^2}{4\pi^2} \int_{\theta_k, \phi_k} \int_{sh} \int_{sh} \exp(-\mathbf{i}\mathbf{k}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{x}'_{\perp})) A(\theta_k, \phi_k) d^2 x_{\perp} d^2 x'_{\perp} \sin \theta_k d\theta_k d\phi_k \end{aligned}$$

Since  $\hat{\mathbf{x}} \cdot \mathbf{k} = \hat{\mathbf{x}} \cdot \mathbf{k}_{\perp} = k_x = k \sin \theta_k \cos \phi_k$  and  $\hat{\mathbf{y}} \cdot \mathbf{k} = \hat{\mathbf{y}} \cdot \mathbf{k}_{\perp} = k_y = k \sin \theta_k \sin \phi_k$ , in the angular integration we can substitute

$$d\theta_k d\phi_k = \left| \frac{\partial(\theta_k, \phi_k)}{\partial(k_x, k_y)} \right| dk_x dk_y = \left| \frac{\partial(k_x, k_y)}{\partial(\theta_k, \phi_k)} \right|^{-1} d^2 k_\perp = \frac{1}{k^2 \sin \theta_k \cos \theta_k} d^2 k_\perp$$

and

$$\sigma = \frac{1}{4\pi^2} \int_{|\mathbf{k}_{\perp}| < k} \int_{sh} \int_{sh} \exp(-\mathbf{i}\mathbf{k}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{x}_{\perp}')) \frac{A(\theta_k, \phi_k)}{\cos \theta_k} d^2 x_{\perp} d^2 x_{\perp}' d^2 k_{\perp}$$

Since the shadow region is much larger than the wavelength, in the double-integration over the area the phase term is rapidly oscillating unless  $k_{\perp} \ll k$ , that is unless  $\theta_k \approx 0$ . Angles  $\theta_k$  substantially different form 0 will not significantly contribute to the integral. We are, essentially, restating the fact that short-wavelength

shadow scattering mostly occurs into the forward directions. Thus, in the angle-dependent term  $\frac{A(\theta_k, \phi_k)}{\cos \theta_k}$  we may set  $\theta_k = 0$ , and we may extend the integration range over  $k_{\perp}$  to infinity:

$$\begin{split} \sigma &= \frac{1}{4\pi^2} \int_{|\mathbf{k}_{\perp}| < \infty} \int_{sh} \int_{sh} \exp(-\mathbf{i}\mathbf{k}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{x}'_{\perp})) \frac{A(0, \phi_k)}{\cos(0)} d^2 x_{\perp} d^2 x'_{\perp} d^2 k_{\perp} \\ &= \frac{1}{4\pi^2} \int_{|\mathbf{k}_{\perp}| < \infty} \int_{sh} \int_{sh} \exp(-\mathbf{i}\mathbf{k}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{x}'_{\perp}))(|c_1|^2 + |c_2|^2) d^2 x_{\perp} d^2 x'_{\perp} d^2 k_{\perp} \\ &= \frac{1}{4\pi^2} \int_{sh} \int_{sh} \left\{ \int_{|\mathbf{k}_{\perp}| < \infty} \exp(-\mathbf{i}\mathbf{k}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{x}'_{\perp})) d^2 k_{\perp} \right\} d^2 x_{\perp} d^2 x'_{\perp} \\ &= \frac{1}{4\pi^2} \int_{sh} \int_{sh} \left\{ \int_{sh} (2\pi)^2 \delta^2 (\mathbf{x}_{\perp} - \mathbf{x}'_{\perp}) d^2 x_{\perp} d^2 x'_{\perp} \right. \end{split}$$

**b**): According to the optical theorem, the total cross section (= the sum of scattering and absorption cross section) is

$$\sigma_{t} = \sigma + \sigma_{abs} = \frac{4\pi}{k} \operatorname{Im} \left[ \epsilon_{0}^{*} \cdot \frac{\mathbf{F}(\mathbf{k}_{0} \cdot \mathbf{k}_{0})}{E_{0}} \right]$$

$$\approx \frac{4\pi}{k} \operatorname{Im} \left[ \epsilon_{0}^{*} \cdot \frac{\mathbf{F}_{sh}(\mathbf{k}_{0} \cdot \mathbf{k}_{0})}{E_{0}} \right]$$

$$= \frac{4\pi}{k} \operatorname{Im} \left[ \frac{\mathrm{i}k}{2\pi} (\epsilon_{0}^{*} \cdot \epsilon_{0}) \frac{E_{0}}{E_{0}} \int_{shadow} \exp(-\mathrm{i}\mathbf{k}_{\perp} \cdot \mathbf{x}_{\perp}) d^{2}x_{\perp} \right]_{\mathbf{k}_{\perp} = 0}$$

$$= 2A_{shadow}$$
(1)

This result makes sense because of the following. As seen in part a), small-angle shadow scattering has a cross section of  $A_{shadow}$ , independent of what happens to the radiation that actually hits the target. Since the radiation that hits the target either gets absorbed or re-scattered into directions  $\mathbf{k} \neq \mathbf{k}_0$ , absorption and scattering of the illuminated portion of the target also have a cross section of  $A_{shadow}$ . The total cross section thus is  $2A_{shadow}$ .