



The LZ Dark Matter Experiment



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(for the LZ Collaboration)

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The LZ Collaboration

March 2017

**36 institutions;
about 250 scientists,
engineers, and
technicians**

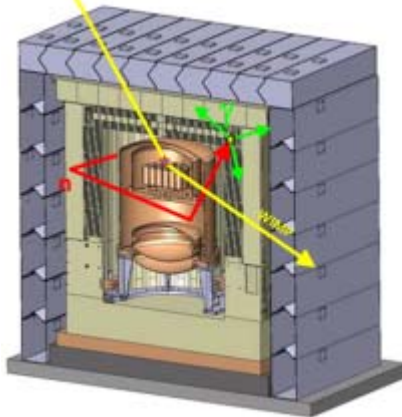


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- 3) MEPhi (Russia)
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LZ = LUX + ZEPLIN

ZEPLIN-III

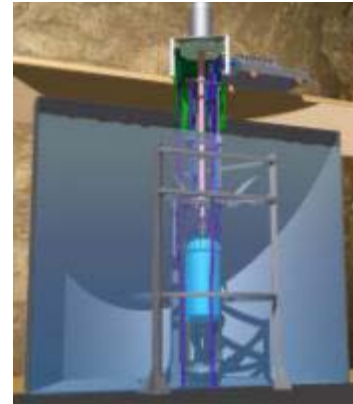


6 kg LXe fid

ZEPLIN pioneered WIMP-search with 2-phase Xe
 $3.9 \times 10^{-44} \text{ cm}^2$

+

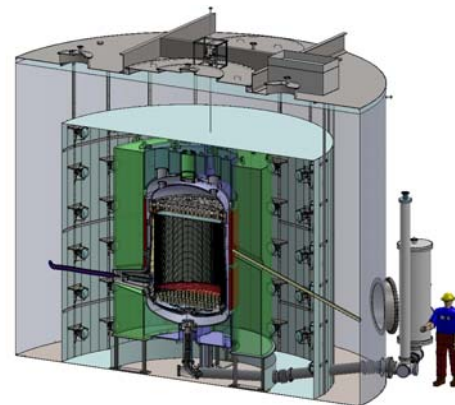
LUX



100 kg

$1.1 \times 10^{-46} \text{ cm}^2$
at 50 GeV/c²
(decommissioned in early 2017)

LZ



5,600 kg

Scale-up using demonstrated technology and experience for low-risk but aggressive program:

- internal background-free strategy
- some infrastructure inherited from LUX
- **LZ expected sensitivity: $2.3 \times 10^{-48} \text{ cm}^2$ with 3-yr run**



Sanford Underground Research Facility



Davis Cavern 1480 m
(4200 mwe)
LZ in LUX Water Tank
South Dakota, USA



LZ Here



Scale up ≈ 50 in Fiducial Mass

LZ

Total mass – 10 T

WIMP Active Mass – 7 T

WIMP Fiducial Mass – 5.6 T

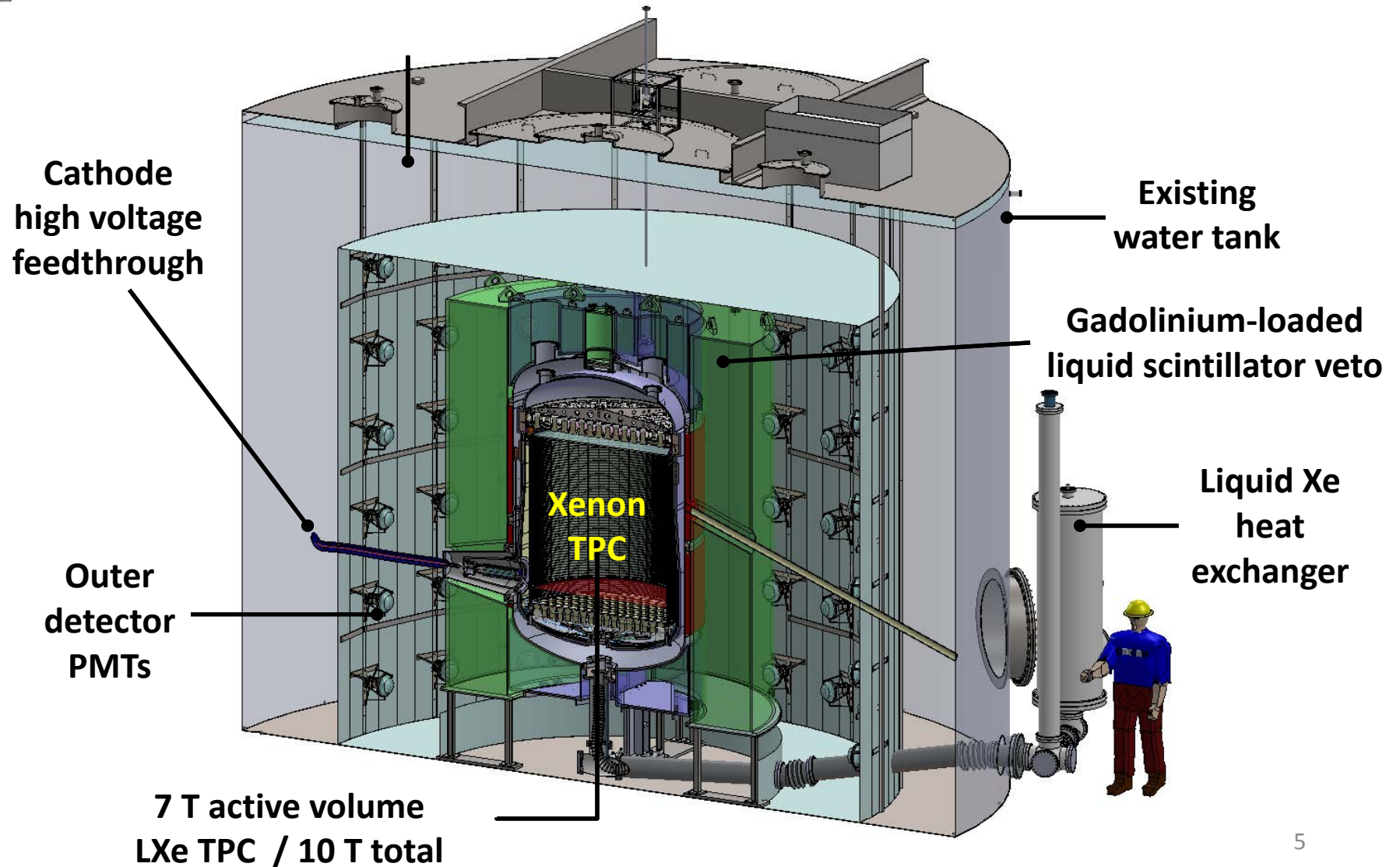


LUX

+ maintain background-free, low-energy response

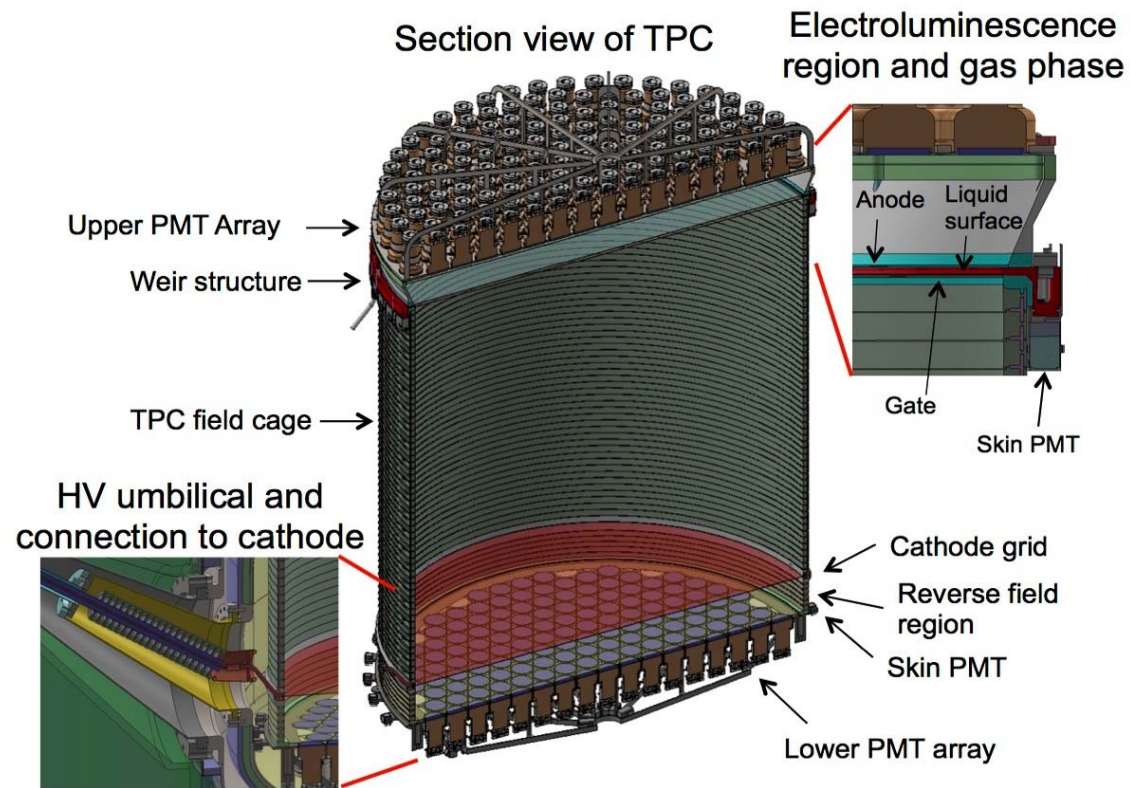


LZ Detector Overview





Dual-phase liquid xenon TPC



- 7 T active LXe mass, 146 cm diameter, 146 cm length
- 488 PMTs (247 top, 241 bot) 3" R11410 PMTs (activity \sim mBq; high QE)
- TPC lined with high-reflectivity PTFE ($R_{\text{PTFE}} \geq 95\%$)
- instrumented "Skin" region optically separated from TPC (180 PMT)



Background Reduction: key design points

- Photomultipliers of ultra-low natural radioactivity
- Low background titanium cryostat
- LUX water shield and an added Gadolinium-loaded liquid scintillator active veto
- Instrumented “skin” region of peripheral xenon as another veto system
- Radon suppression during construction, assembly and operations
- Ultra-low levels of Kr in Xe



Performance Drivers

Detector Parameter	Reduced	Baseline	Goal
Light collection (PDE)	0.05	0.075	0.12
Drift field (V/cm)	160	310	650
Electron lifetime (μs)	850	850	2800
PMT phe detection	0.8	0.9	1.0
N-fold trigger coincidence	4	3	2
^{222}Rn (mBq in active region)	13.4	13.4	0.67
Live days	1000	1000	1000

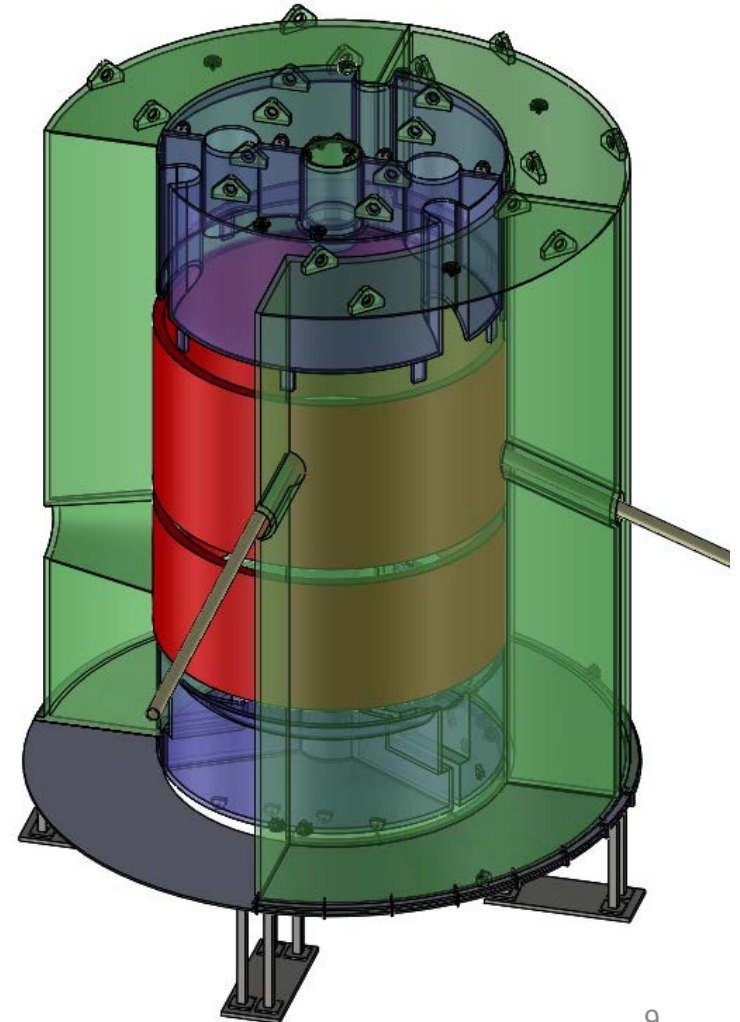
- $5.8 \text{ keV}_{\text{nr}}$ S1 threshold ($4.5 \text{ keV}_{\text{nr}}$ LUX)
- 0.31 kV/cm drift field, 99.5% ER/NR disc.
(already surpassed in LUX at 0.2 kV/cm)



The Outer Detector (OD)

- Essential to utilize most Xe, maximize fiducial volume
- Hermetic measurement of penetrating backgrounds
- Segmented tanks – installation constraints (shaft, water tank)
- 60 cm thick, 21.5 T of Gadolinium-loaded LAB* liquid scintillator, OK underground
- 97% efficiency for neutrons
- Daya Bay legacy, scintillator & tanks (and people)

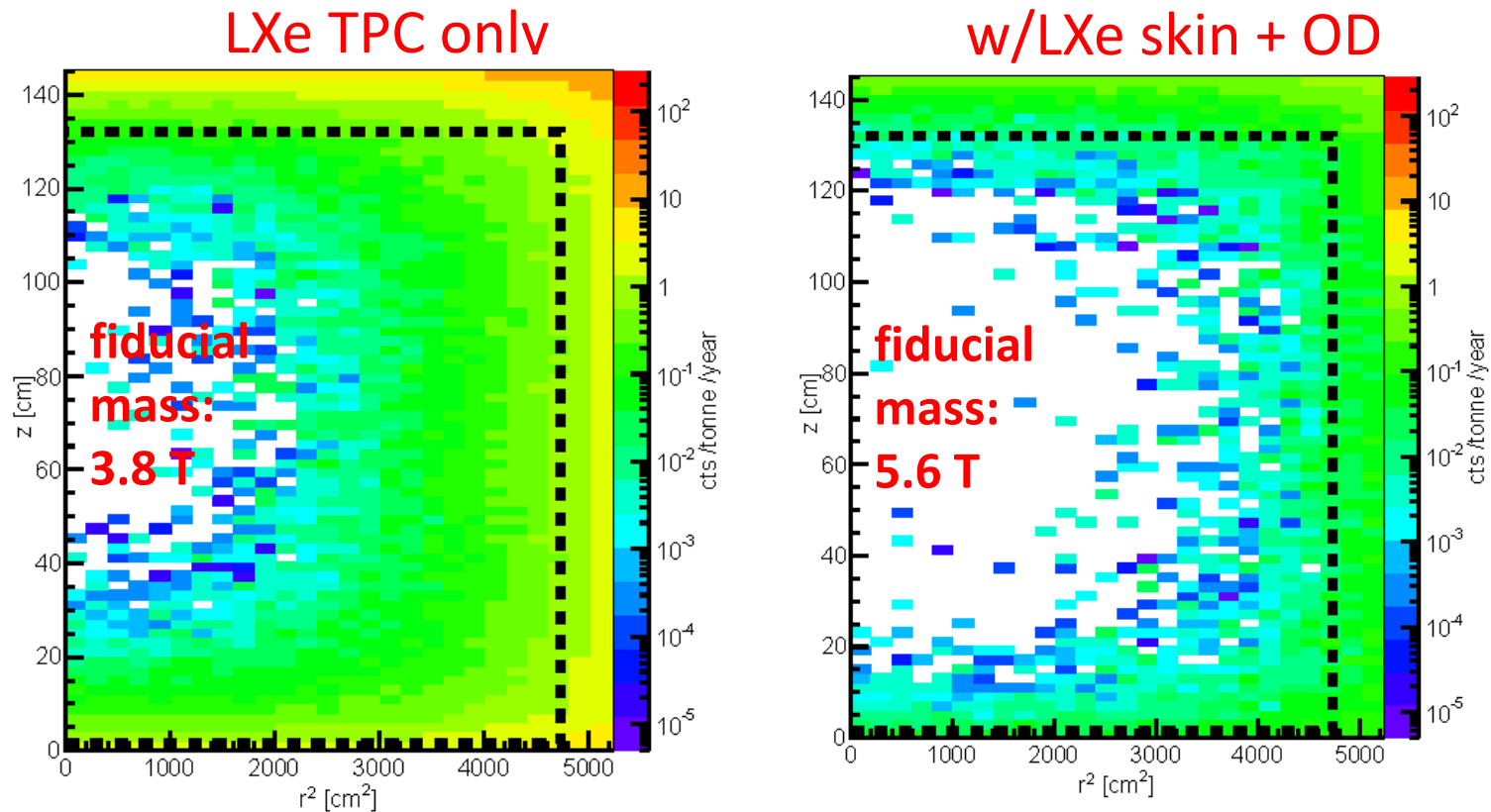
* Linear AlkylBenzene





Powerful Background Rejection

Simulated single NR scatter in TPC before/after Skin+OD vetoes



- Increases effective fiducial mass from 3.8 T \rightarrow 5.6 T
- Internal backgrounds now dominate



Background Control

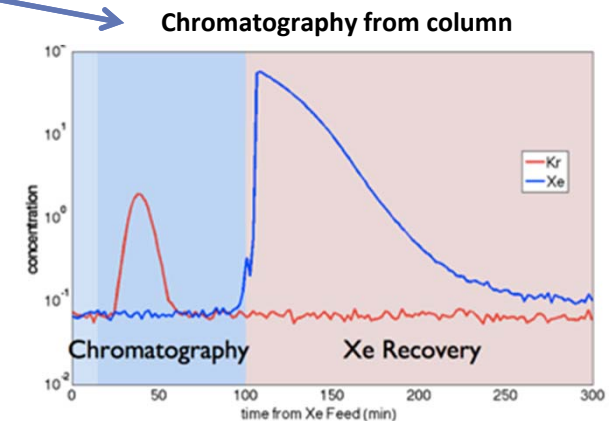
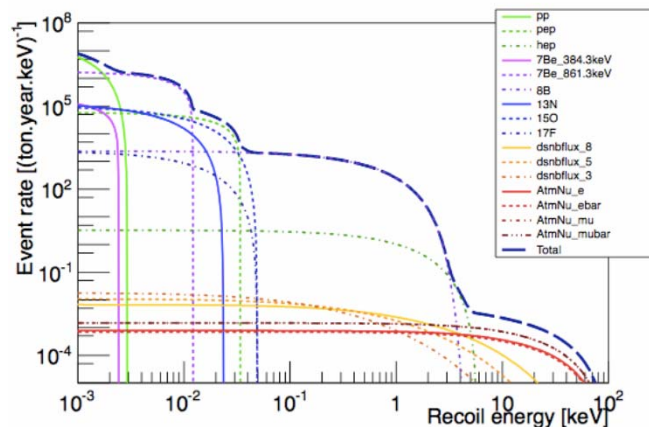
- Assay and assess all candidate detector materials and components with many dedicated screening facilities prior to adoption
- Assay techniques:
 - gamma spectroscopy
 - mass spectroscopy
 - neutron activation analysis(NAA)
 - radon emanation counting
 - alpha spectroscopy

	Intrinsic Contamination Backgrounds	Mass (kg)	U early (mBq/kg)	U late (mBq/kg)	Th early (mBq/kg)	Th late (mBq/kg)	Co60 (mBq/kg)	K40 (mBq/kg)	ER (cts)	NR (cts) (w SF rej.)
Detector components	Upper PMT Structure	46.7	5.25	0.80	1.07	0.72	0.03	3.77	0.14	0.001
	Lower PMT Structure	71.7	2.69	0.24	0.42	0.30	0.00	1.36	0.08	0.001
	R11410 3" PMTs	91.9	71.63	3.20	3.12	2.99	2.82	15.41	1.46	0.013
	R11410 PMT Bases	2.8	287.74	75.80	28.36	27.93	1.43	69.39	0.36	0.004
	R8778 2" PMTs	6.1	137.50	59.38	16.88	16.88	16.25	412.50	0.13	0.008
	R8520 Skin 1" PMTs	2.2	60.50	5.19	4.75	4.75	24.20	332.76	0.02	0.001
	R8520 Skin PMT Bases	0.2	212.95	108.46	42.19	37.62	2.23	123.61	0.00	0.000
	PMT Cabling	104.2	30.13	1.55	3.32	3.15	0.65	33.12	1.45	0.001
	TPC PTFE	184.0	0.02	0.02	0.03	0.03	0.00	0.12	0.06	0.008
	Grid Wires	0.18	1.20	0.27	0.33	0.49	1.60	0.40	0.00	0.000
	Grid Holders	92.3	2.86	0.83	0.94	0.82	1.42	2.82	0.97	0.008
	Field Shaping Rings	92.5	5.49	0.13	0.32	0.26	0.00	0.71	0.27	0.004
	TPC Sensors	1.32	22.40	8.94	11.38	9.57	0.35	19.44	0.01	0.002
	TPC Thermometers	0.08	335.50	90.46	38.48	25.02	7.26	3,359	0.06	0.000
	Xe Recirculation Tubing	15.1	0.79	0.18	0.23	0.33	1.05	0.30	0.00	0.000
	HV Conduits and Cables	137.7	2.0	2.0	0.4	0.6	1.4	1.2	0.04	0.001
	HX and PMT Conduits	199.6	3.36	0.48	0.48	0.58	1.24	1.47	0.05	0.001
	Cryostat Vessel	2409.6	1.70	0.14	0.30	0.25	0.10	0.64	0.72	0.014
	Cryostat Seals	33.7	73.91	26.22	3.22	4.24	10.03	69.12	0.45	0.002
	Cryostat Insulation	23.8	18.91	18.91	3.45	3.45	1.97	51.65	0.43	0.007
Cryostat Teflon Liner	26.0	0.02	0.02	0.03	0.03	0.00	0.12	0.00	0.000	
Outer Detector Tanks	3199.3	0.16	0.39	0.02	0.06	0.04	5.36	0.45	0.001	
Liquid Scintillator	17640.3	0.01	0.01	0.01	0.01	0.00	0.00	0.03	0.000	
Outer Detector PMTs	204.7	570	470	395	388	0.00	534	0.01	0.000	
Outer Detector PMT Supports	770.0	12.35	12.35	4.07	4.07	9.62	9.29	0.00	0.000	
Subtotal (Detector Components)									7.18	0.077
Xenon contaminants	222Rn (1.65 μBq/kg)								597	-
	220Rn (0.08 μBq/kg)								101	-
	natKr (0.015 ppt g/g)								24.5	-
	natAr (0.45 ppb g/g)								2.47	-
	210Bi (0.1 μBq/kg)								40.0	-
Environment, cosmogenic, surface contamination	Laboratory and Cosmogenics								4.3	0.06
	Fixed Surface Contamination								0.19	0.37
Subtotal (Non-v counts)									776	0.50
Physics backgrounds	Physics Backgrounds									
	136Xe 2vββ								67	0
	Astrophysical ν counts (pp+7Be+13N)								255	0
	Astrophysical ν counts (8B)								0	0
	Astrophysical ν counts (Hep)								0	0.21
	Astrophysical ν counts (diffuse supernova)								0	0.05
	Astrophysical ν counts (atmospheric)								0	0.46
Subtotal (Physics backgrounds)									322	0.72
Total									1,100	1.22
Total (with 99.5% ER discrimination, 50% NR efficiency)									6.10	0.61



Internal Backgrounds

- Rn (and Kr) dominant internal background sources
- Rn:
 - Emanates from most materials
 - 13.4 mBq requirement, 0.67 mBq goal (active region)
 - Rn removal system at UMich
 - Four measurement systems with ~ 0.1 mBq sensitivity
 - Main assembly laboratory at SURF will have reduced radon air system
- Kr:
 - Remove ^{85}Kr to <15 ppq using gas chromatography
 - Neutrino-induced backgrounds

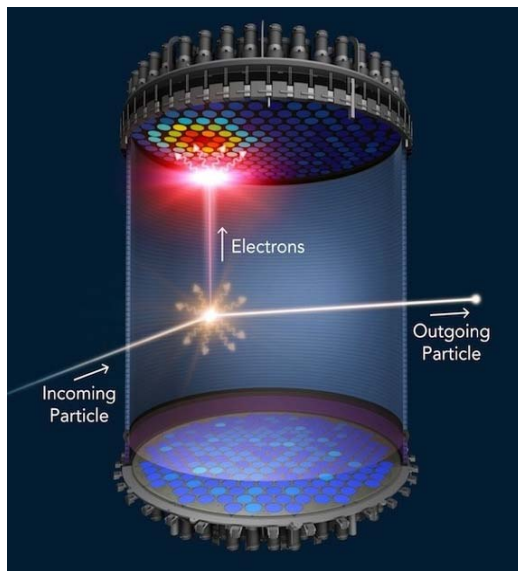




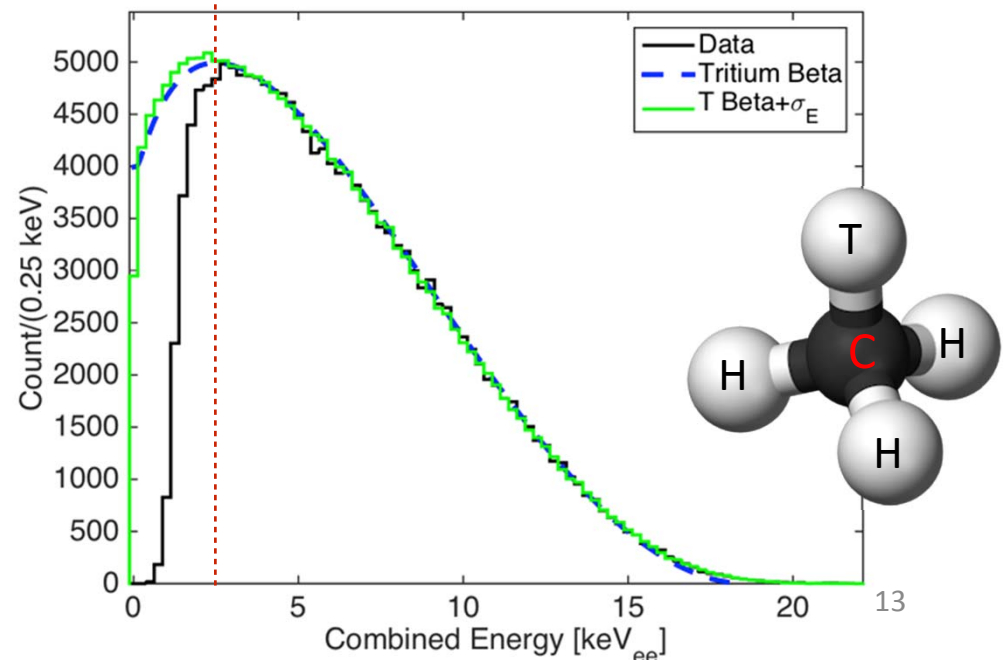
Calibrations

- Expand upon successful LUX program (and other experience)
- DD Neutron Generator (Nuclear Recoils)
- Tritiated Methane (Electron Recoils)
- Movable photon sources e.g. tubes penetrating cryostat
- Additional sources e.g. YBe source for low energy (Nuclear Recoils)

DD neutron calibration



Tritium Beta Spectrum Measured in LUX





Detector Prototyping

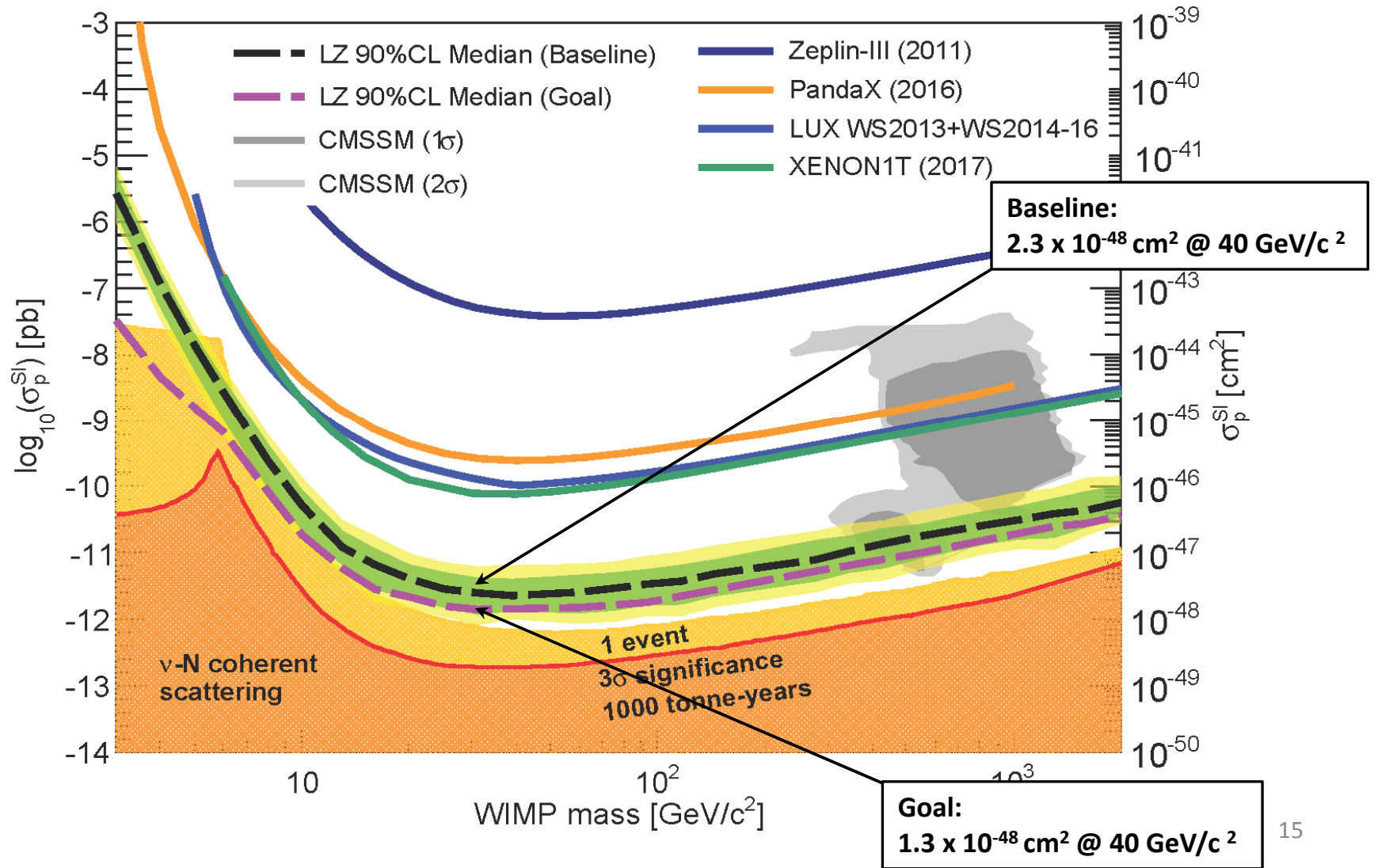
Extensive program of prototype development undertaken, with three general approaches:

- Testing in liquid argon, primarily of HV elements at LBNL
- Design choice and validation in small (few kg) LXe test chambers in many locations: LLNL, UC Berkeley, LBNL, U Michigan, UC Davis, Imperial College, MEPHI, LIP ([arXiv:1507.01310](#), [JINST 10, (2015) P10040], [arXiv:1608.01717](#) [NIMA 856 (2017) 86])
- System test platform at SLAC,
 - Phase I about 100 kg of LXe
 - Phase II GXe only: full scale grids (preproduction and final LZ grids)



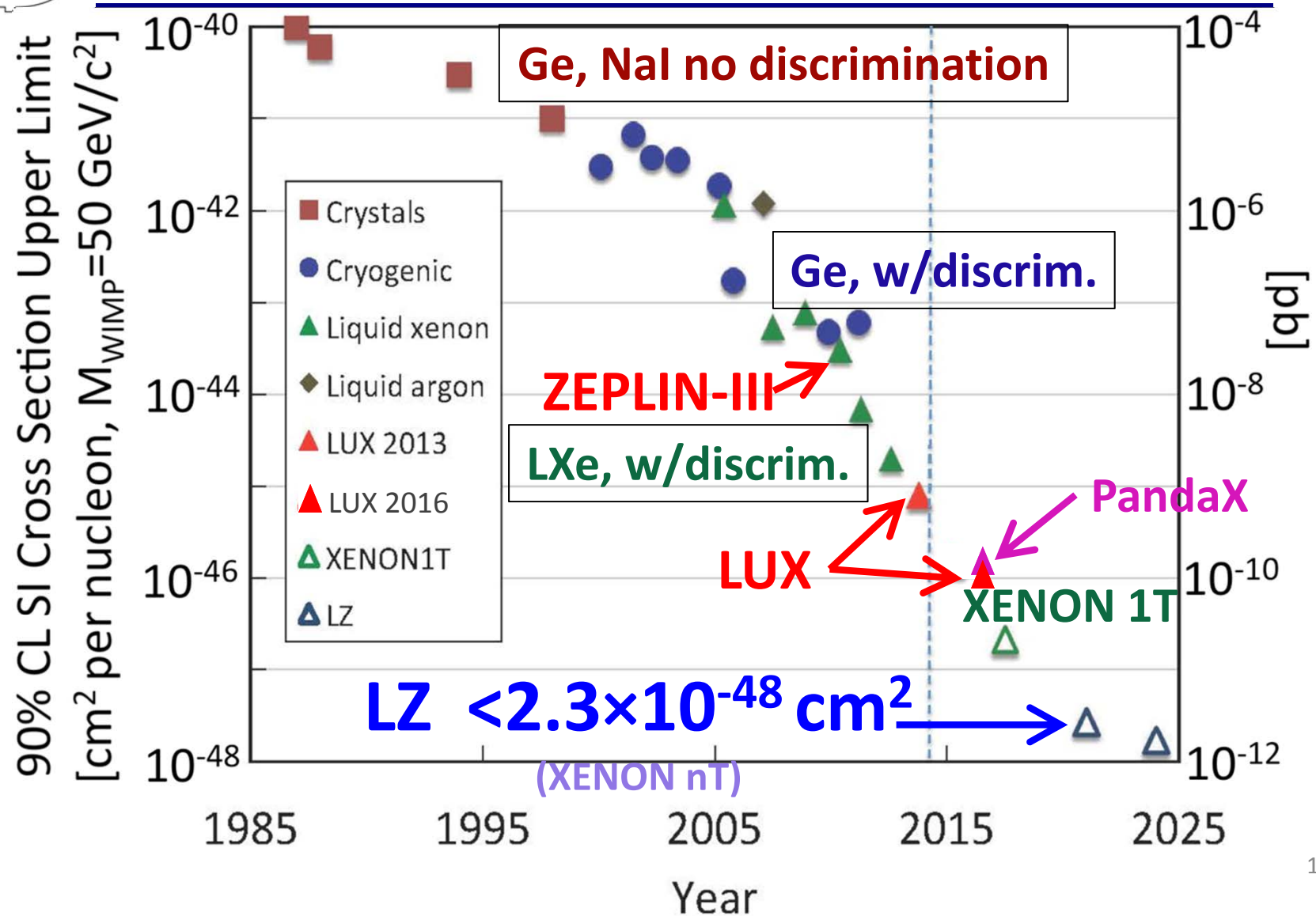
Projected LZ Sensitivity – Spin Independent

(5.6 T fiducial, 1000 live-days)





Time Evolution





Timeline

Year	Month	Activity
2012	March	LZ (LUX-ZEPLIN) collaboration formed
2014	July	LZ Project selected in US and UK
2015	April	DOE CD-1/3a approval, similar in UK Conceptual Design Report arXiv: 1509.02910
2017	February	DOE CD-2/3b approval Technical Design Report arXiv: 1703.09144
2017	March	LUX removed from underground
2017	June	Begin preparations for surface assembly at SURF
2018	July	Begin underground installation
2019	late	Begin commissioning
2024+		Planning on 5+ years of operations



Summary

- LZ Project well underway, on schedule
- Long lead-time item procurement underway: Xenon, PMTs, Cryostat, vessel, etc.
- Materials screening program well underway
- LZ benefits from the excellent LUX calibration techniques and understanding of background
- Will explore significant fraction of available phase space:
 - WIMP sensitivity $2.3 \times 10^{-48} \text{ cm}^2 @ 40 \text{ GeV}/c^2$ with 1,000 live days (and approaching neutrino floor)



Extra Slides



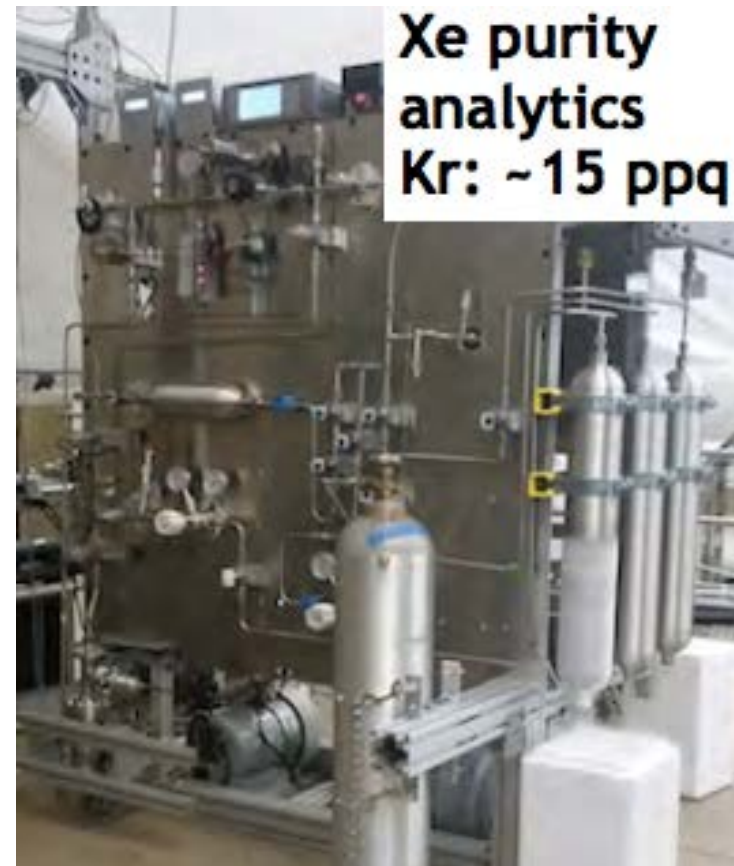
LZ Calibration (details)

Isotope	What	Purpose	Deployment
Tritium	beta, $Q = 18.6$ keV	ER band	Internal
^{83m}Kr	beta/gamma, 32.1 keV and 9.4 keV	TPC (x, y, z)	Internal
^{131m}Xe	164 keV γ	TPC (x, y, z), Xe skin	Internal
^{220}Rn	various α 's	xenon skin	Internal
AmLi	(α, n)	NR band	CSD
^{252}Cf	spontaneous fission	NR efficiency	CSD
^{57}Co	122 keV γ	Xe skin threshold	CSD
^{228}Th	2.615 MeV γ , various others	OD energy scale	CSD
^{22}Na	back-to-back 511 keV γ 's	TPC and OD sync	CSD
^{88}Y Be	152 keV neutron	low-energy NR response	External
^{205}Bi Be	88.5 keV neutron	low-energy NR response	External
^{206}Bi Be	47 keV neutron	low-energy NR response	External
DD	2,450 keV neutron	NR light and charge yields	External
DD	272 keV neutron	NR light and charge yields	External



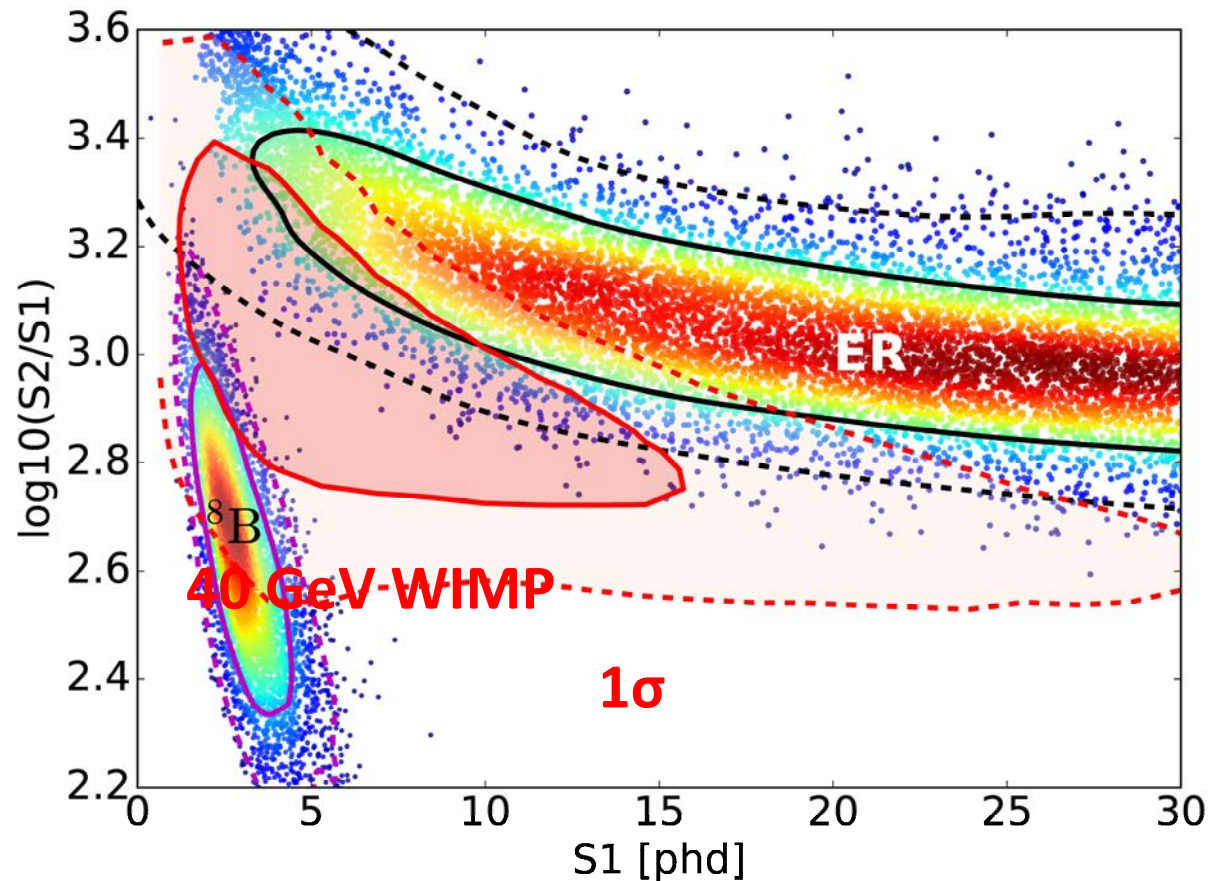
Xe purification and cryogenics

- Gas phase purification through getter: 10 tons/2.5 days
- Achieved sensitivity is 0.007 ppt (g/g), or 7 ppq (g/g).
- High-efficiency two-phase heat exchange
- LN₂ thermosiphon-based cryogenics: multiple cooling locations





Simulation of NR/ER backgrounds



Simulations of the most prominent ER and NR (from ^8B) backgrounds are plotted in the $\log_{10}(S_2/S_1)$ - S_1 plane. **The statistics shown represent 5x the expected ER background and 500x the expected NR background in the nominal LZ exposure.** The red tinted area shows the expectation for events from a 40 GeV/ c^2 -mass WIMP, falling between the two background populations with the region enclosed by the solid(dashed) line representing the 1σ (2σ) band.



Non-WIMP physics

- Effective Field Theory Interaction Decomposition
- Neutrinoless Double Beta Decay
- Axions/Axion-like-particles, leptophilic DM, fractionally charged particles
- External Neutrino Physics:
 - Solar
 - Supernova
 - Sterile Neutrino